

1 The seismic signature of lithospheric deformation beneath eastern North America
2 due to Grenville and Appalachian orogenesis

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17 **ABSTRACT**

18 The eastern margin of North America has been affected by two complete Wilson cycles
19 of supercontinental assembly and breakup over the past ~1.3 Ga. Evidence of these processes is
20 apparent in the surface geology; however, the geometry, strength, and extent of lithospheric
21 deformation associated with these events are poorly known. Observations of seismic anisotropy
22 in the continental lithosphere can shed light on past deformation processes, but information about
23 the depth distribution of anisotropy is needed. Here we investigate the azimuthal dependence of
24 transverse component receiver functions at broadband seismic stations in eastern North America
25 to constrain sharp contrasts in seismic anisotropy with depth. We examined data from six
26 permanent seismic stations, including three that lie just to the east of the Grenville Front and
27 three that lie within the Appalachian Mountains. A harmonic stacking modeling method was
28 used to constrain the presence of anisotropic interfaces within the crust and mantle lithosphere. A
29 comparison among stations located to the east and to the west of the Grenville Front reveals
30 evidence for different lithospheric anisotropy across the front, which in turn argues for
31 significant lithospheric deformation associated with the Grenville Orogeny. Stations located in

32 the Appalachians exhibit a striking signature of strong and multilayered anisotropy in the lower
33 crust, consistent with observations in modern orogens, as well as the lithospheric mantle. Our
34 observations constrain the existence and approximate depths of contrasts in anisotropy within the
35 lithosphere and may be used for future testing of specific hypotheses regarding lithospheric
36 deformation associated with orogenesis.

37

38 **INTRODUCTION**

39 The formation of mountain belts at convergent plate boundaries is one of the most
40 fundamental processes in plate tectonics. Despite the importance of orogenesis to the plate
41 tectonic system, our understanding of how deformation is accommodated in the crust and mantle
42 lithosphere during orogenesis remains poorly understood. This is true both in present-day
43 orogenic systems and in regions that have been affected by past orogenic events. Lithospheric
44 deformation during orogenesis, while poorly understood, is crucial for our understanding of the
45 evolution of topography, the partitioning of strain in collisional settings, and the evolution and
46 modification of continental margins such as eastern North America. Furthermore, lithospheric
47 deformation reflects the rheology of the crust and mantle lithosphere, which remain poorly
48 understood despite extensive study.

49 One type of observation that can inform our view of deformation in the lower crust and
50 upper mantle is seismic anisotropy, or the directional dependence of seismic wave propagation.
51 This is because there is a direct, albeit complicated, link between strain and the resulting seismic
52 anisotropy via the crystallographic preferred orientation (CPO) of anisotropic minerals. In the
53 upper mantle, seismic anisotropy primarily reflects the CPO of olivine (e.g., Karato et al., 2008),
54 while in the lower crust, there might be contributions from amphiboles, micas, and/or quartz (e.g.,

55 Ward et al., 2012; Liu et al., 2015; Ko and Jung, 2015). There are many published observations
56 of seismic anisotropy within continental settings in general (e.g., Fouch and Rondenay, 2006;
57 Tommasi and Vauchez, 2015), and beneath eastern North America in particular (e.g., Barruol et
58 al., 1997; Levin et al., 1999; Wagner et al., 2012; Long et al., 2010, 2016; Yang et al., 2017). A
59 challenge in their interpretation, however, is that it is often difficult to obtain good constraints on
60 the depth extent of anisotropy. For example, the splitting or birefringence of *SKS* waves, perhaps
61 the most common method for studying continental anisotropy, is a path-integrated measurement
62 involving nearly vertically propagating shear waves, so its depth resolution is poor (e.g., Long
63 and Silver, 2009). Complementary constraints on anisotropy can be obtained from surface wave
64 dispersion analysis (e.g., Deschamps et al., 2008; Yuan and Romanowicz, 2010) and by
65 combining different types of seismic data (e.g., Yuan and Levin, 2014; Bodin et al., 2016).

66 One analysis technique that can yield good resolution of the depth distribution of
67 anisotropy within the lithosphere is anisotropic receiver function analysis, which can reveal
68 sharp contrasts in anisotropic structure with depth (e.g., Levin and Park, 1998, 1999; Bostock,
69 1998; Frederiksen and Bostock, 2000). This technique has recently been applied to data from
70 continental settings; in particular, there are several recent studies that have applied it to study
71 both crustal anisotropy (e.g., Porter et al., 2011; Schulte-Pelkum and Mahan, 2014a; Liu et al.,
72 2015) and anisotropy within the lithospheric mantle (e.g., Yuan and Levin, 2014; Wirth and
73 Long, 2014; Ford et al., 2016). This analysis strategy, which looks for P-to-SH wave conversions
74 at a dipping and/or anisotropic interface, can provide unambiguous evidence for anisotropy
75 (usually under the assumption of hexagonal symmetry) and constrains the timing of converted
76 waves (and thus the likely depth of the interfaces). A disadvantage of the technique, however, is
77 that the full anisotropic geometry in each layer across an interface cannot be easily deduced from

78 the observations themselves. Detailed forward modeling can identify plausible anisotropic
79 models (e.g., Wirth and Long, 2012, 2014; McCormack et al., 2013), but forward modeling is
80 generally non-unique and computationally expensive, and the tradeoffs among different model
81 parameters are typically strong (e.g., Porter et al., 2011; Wirth et al., 2017). Simplified modeling
82 using harmonic decomposition or similar techniques (e.g., Shiomi and Park, 2008; Bianchi et al.,
83 2010; Schulte-Pelkum and Mahan, 2014a; Liu et al., 2015; Ford et al., 2016; Olugboji and Park,
84 2016; Park and Levin, 2016) can identify the presence and first-order features of dipping and/or
85 anisotropic interfaces without extensive computation of synthetic seismograms.

86 In this study we examine data from six long-running broadband seismic stations in the
87 eastern United States (Figure 1) that overlie lithosphere that has been affected by past mountain-
88 building events, namely the Grenville and Appalachian orogenies. The goal of this work is to
89 explore the nature and depth extent of deformation of the crust and mantle lithosphere due to
90 orogenesis using the technique of anisotropic receiver function analysis. We apply a harmonic
91 decomposition modeling approach to our computed receiver functions to identify interfaces
92 within the continental lithosphere with a dipping and/or anisotropic character beneath our
93 selected stations. Three of our stations lie just to the east of the Grenville Front, which marks the
94 mapped westward extent of deformation during the Grenville Orogeny, while three of our
95 stations are situated within the Appalachians. This particular study is motivated by previous
96 work by Wirth and Long (2014), who identified evidence for a contrast in anisotropy within the
97 mantle lithosphere at stations in the Granite-Rhyolite province of eastern North America. The
98 comparison of our new results with these previous results, obtained at stations located to the west
99 of the Grenville Front, allows us to investigate whether and how Grenvillian and Appalachian
100 orogenesis modified the anisotropic structure of the lithosphere.

101

102 **TECTONIC SETTING OF THE GRENVILLE AND APPALACHIAN OROGENIES**

103 The eastern margin of North America (Figure 1), today a passive margin that sits in the
104 interior of the North American plate, has been shaped by two complete supercontinent cycles
105 over the past ~1.3 Ga of Earth history (e.g., Thomas, 2006; Cawood and Buchan, 2007; Benoit et
106 al., 2014). This tectonic history includes multiple and protracted episodes of orogenesis,
107 including the Grenville orogenic cycle that culminated in the formation of the supercontinent
108 Rodinia (e.g., Whitmeyer and Karlstrom, 2007; McLelland et al., 2010) and the Appalachian
109 orogenic cycle that culminated in the formation of the supercontinent Pangea (e.g., Hatcher et al.,
110 2010; Hibbard et al., 2010). Dispersal of each of these supercontinents was accomplished via
111 continental rifting, including a major diachronous rifting episode between roughly 750-550 Ma
112 (e.g., Li et al., 2008; Burton and Southworth, 2010) that broke apart Rodinia and shaped the pre-
113 Appalachian eastern margin of Laurentia (e.g., Allen et al., 2010). The later breakup of Pangea
114 was also accomplished via a complex set of rifting processes that was complete by roughly ~190
115 Ma (e.g., Frizon de Lamotte et al., 2015) but began substantially earlier (~250 Ma), and was
116 accompanied (at ~200 Ma) by voluminous magmatism associated with the Central Atlantic
117 Magmatic Province (CAMP) large igneous province (e.g., McHone, 1996, 2000; Schlische et al.,
118 2003). These multiple tectonic episodes were likely accompanied by substantial deformation of
119 the crust and mantle lithosphere as the edges of the (progressively growing) North American
120 continent were modified.

121 The term “Grenville Orogeny” is used somewhat inconsistently in the literature (e.g.,
122 McLelland et al., 2010), but here we use the term to encompass both the Elzevirian orogeny at
123 ~1.3-1.2 Ga, which sutured the Elzevir Block to the Laurentian margin (e.g., Moore and

124 Thompson, 1980; Whitmeyer and Karlstrom, 2007; Bartholomew and Hatcher, 2010), as well as
125 later (~1.09-0.98 Ga; Whitmeyer and Karlstrom, 2007) continent-continent collision that formed
126 Rodinia (e.g., McLelland et al., 1996; Rivers, 1997) as the Grenville Province was joined to
127 Laurentia. The Grenville deformation front, which represents the westward extent of deformation
128 due to Grenvillian orogenesis (e.g., Culotta et al., 1990; Whitmeyer and Karlstrom, 2007) cuts
129 through eastern Michigan and western Ohio and extends to the south and west through Alabama
130 (Figure 1). In the context of the work presented here, the Grenville Front separates the seismic
131 stations within the Granite-Rhyolite province examined by Wirth and Long (2014) and those in
132 this study. Of the stations we examine in this study (Figure 1), station ACSO is located within
133 the Granite-Rhyolite province but to the east of the Grenville Front, while station ERPA is
134 located near the suture between the Elzevir block and the rest of the Granite-Rhyolite province
135 and station BINY is located within the Grenville Province, as defined by Whitmeyer and
136 Karlstrom (2007).

137 The later Appalachian orogeny, a detailed overview of which is given by Hatcher (2010),
138 involved three distinct phases of orogenesis over a period of several hundred Ma, and formed the
139 present-day Appalachian Mountains. The first, the Taconic orogeny, involved the accretion of
140 arc terranes onto the margin of Laurentia between roughly 496-428 Ma (Karabinos et al., 1998;
141 Hatcher, 2010), while the later phases (the Acadian/Neocadian and Alleghanian orogenies)
142 involved superterrane accretion (Carolina, Avalon, Gander, and Meguma superterranes, all of
143 peri-Gondwanan affinity) and continental collision. The Acadian orogeny began around ~410
144 Ma and primarily affected the northern Appalachians with transpressive, north-to-south collision
145 (e.g., Hatcher, 2010; Ver Straeten, 2010). The Alleghanian orogeny (e.g., Geiser and Engelder,
146 1983; Sacks and Secor, 1990; Hatcher, 2010; Bartholomew and Whitaker, 2010), which began

147 around ~300 Ma and was complete by roughly ~250 Ma, culminated in the final assembly of the
148 Pangea supercontinent. Taken together, the protracted series of orogenic events that make up the
149 Appalachian orogenic cycle produced widespread deformation, volcanism, and metamorphism,
150 and produced the topographic signature of the Appalachian Mountains that is still evident today
151 (Hatcher, 2010). Three of the seismic stations that we use in this study (SSPA, MCWV, and
152 TZTN) are located in the Appalachian Mountains, along the edge of the Laurentian core of the
153 North American continent, in regions that were likely affected by deformation associated with
154 Appalachian orogenesis.

155

156 **DATA AND METHODS**

157 In this paper we use anisotropic receiver function (RF) analysis to identify and
158 characterize contrasts in seismic anisotropy with depth within the continental crust and mantle
159 lithosphere, following our previous work (Wirth and Long, 2014; Ford et al., 2016). This
160 technique involves the computation of RFs – that is, a time series that reflects converted seismic
161 phases due to structure beneath a seismic station (e.g., Langston, 1979) – for both the radial
162 (oriented in a direction that points from receiver to source) and transverse (oriented orthogonal to
163 the radial) horizontal components of a seismogram. For the case of horizontal interfaces and
164 purely isotropic structure, all of the *P*-to-*S* converted energy would be expected to arrive on the
165 radial component, as the *P* and *SV* wavefields are coupled and (in this case) propagate
166 independently of the *SH* wavefield. When dipping interfaces or anisotropic structure are present,
167 some of the *P*-to-*S* converted energy will arrive on the transverse component due to *P*-to-*SH*
168 conversions (e.g., Levin and Park, 1997, 1998). For the case of a dipping isotropic interface,
169 these converted waves will manifest as an arrival on the transverse component RF trace that

170 changes polarity twice across the full backazimuthal range (that is, a two-lobed polarity flip). For
171 the case of a contrast in anisotropy (with a hexagonal symmetry and a horizontal symmetry axis),
172 these conversions will manifest as an arrival on the transverse component RF trace that changes
173 polarity four times across the full backazimuthal range (that is, a four-lobed polarity flip). For the
174 case of a contrast in anisotropy with a plunging axis of symmetry, a mix of two-lobe and four-
175 lobe polarity changes on the transverse component RFs are predicted. One way to distinguish
176 between a dipping (isotropic) interface and a contrast in anisotropy with a plunging symmetry
177 axis, at least in theory, is to examine the transverse components in the time range associated with
178 the direct P arrival. For a dipping interface, a change in polarity on the transverse components at
179 a time associated with the P -to- S conversion at an interface will be accompanied by an arrival
180 with the opposite polarity at $t = 0$; for a contrast in anisotropy at depth, this zero-time arrival is
181 not present. One caveat, however, is that this analysis is complicated when there are multiple
182 interfaces at depth, as multiple zero-time arrivals may interfere. Predicted RF patterns for a suite
183 of simple models that illustrate these ideas are shown and discussed in detail in Schulte-Pelkum
184 and Mahan (2014b) and Ford et al. (2016).

185 We selected six long-running broadband seismic stations in the eastern United States for
186 analysis in this study, all part of the permanent US (United States National Seismic Network) or
187 IU (Global Seismographic Network) networks (Figure 1). In order to achieve good azimuthal
188 coverage, we focus on a small number of high-quality stations with long run times. Of these,
189 stations ACSO, ERPA, and BINY are located in regions that were likely affected by Grenville
190 orogenesis, as described above; we collectively refer to these as the “Grenville” stations in this
191 study. Stations MCWV, SSPA, and TZTN are located in regions of present-day mountainous
192 topography that were likely affected by Appalachian orogenesis (“Mountain” stations). We

193 examined at least ten years' worth of data at each station, selecting events of magnitude 5.8 and
194 greater at epicentral distances between 30° and 100° (Figure 2) for analysis.

195 Our preprocessing methodology is identical to that of Ford et al. (2016); each
196 seismogram trace was cut to identical length, rotated into vertical, radial, and transverse
197 components, and bandpass filtered between 0.02-2 Hz for *P* wave picking. We visually inspected
198 each trace for a clear *P* wave arrival, discarding those traces that did not display an unambiguous
199 *P* wave, and manually picked the *P* arrival using the Seismic Analysis Code (SAC). Before
200 computing RFs, we rotated the components into the LQT reference frame (using a near-surface *P*
201 wave velocity of 3.5 km/sec) to account for the fact that the incoming *P* arrivals are not perfectly
202 vertical, again following Ford et al. (2016). We note that while all RFs were computed in the
203 LQT coordinate system, for simplicity we refer to “radial” and “transverse” RFs to follow
204 common terminology conventions (e.g., Levin and Park, 1997, 1998). We emphasize, however,
205 that the LQT and RTZ coordinate systems are not strictly equivalent, and that our rotation into
206 the LQT coordinate system is imperfect, as we used a single representative near-surface velocity
207 to calculate the rotations. Finally, before RF computation, we applied a final bandpass filter to
208 each waveform, using a highpass cutoff of 0.02 Hz and a variable lowpass cutoff of 0.5 or 1 Hz.

209 To compute radial and transverse component RFs, we used the frequency domain
210 multitaper correlation estimator of Park and Levin (2000), which is a commonly applied RF
211 computation technique (recent examples can be found in Wirth and Long, 2012, 2014;
212 McCormack et al., 2013; Yuan and Levin, 2014; Liu et al., 2015; Levin et al., 2016; Olugboji
213 and Park, 2016; Ford et al., 2016). Again following Ford et al. (2016), we computed individual
214 RFs and then stacked after corrections for variations in slowness of the direct *P* wave arrival,
215 which depends on epicentral distance. We computed single-station radial component stacks over

216 all backazimuths, as well as stacks of both radial and transverse component RFs as a function
217 both epicentral distance (which allows us to evaluate the effects of multiply reflected waves on
218 our results) and backazimuth (which allows us to identify dipping and/or anisotropic structure).
219 Our stacking and plotting conventions are shown in Figure 3, which illustrate a set of RF stacks
220 and plots at station SSPA as an example. In addition to the RF stacks shown in Figure 3a and 3b,
221 we also visualize the backazimuthal variability in transverse component RF energy within
222 specific time ranges (corresponding to specific depth ranges within the lithosphere) using rose
223 diagrams, following the convention of Wirth and Long (2014) and as illustrated in Figure 3d.
224 This plotting convention allows for visual inspection of the azimuthal variability in polarity
225 (represented with blue and red colors) at specific time (depth) ranges. The character of any
226 variability (two-lobed, four-lobed, or a mixture) sheds light on the type of interface (dipping,
227 anisotropic, or both). Distinguishing between a dipping interface and a plunging axis of
228 anisotropy requires an inspection of the transverse component RFs at the arrival time of the
229 direct P wave, as noted above, although in the presence of multiple interfaces this can be difficult.

230 Finally, we implement a harmonic decomposition technique (Shiomi and Park, 2008;
231 Bianchi et al., 2010; Park and Levin, 2016) that allows for the straightforward identification of
232 time ranges within the RF traces that are likely associated with conversions at interfaces of
233 different character. Specifically, harmonic decomposition identifies coherent energy on the RF
234 traces that correspond to arrivals that are constant across backazimuth, that vary as a function of
235 $\cos(\theta)$ or $\sin(\theta)$ (where θ is backazimuth), that vary as a function of $\cos(2\theta)$ or $\sin(2\theta)$, or a
236 combination of these. Harmonic decomposition modeling thus serves as a useful complement to
237 the simple visual inspection of backazimuthal RF gathers, and has been applied by several recent
238 studies (e.g., Liu et al., 2015; Olugboji and Park, 2016; Ford et al., 2016). A detailed description

239 of our implementation of harmonic decomposition can be found in Ford et al. (2016) and a recent
240 technical overview is given by Park and Levin (2016). Briefly, for each time window the stacked
241 RFs (both radial and transverse components) are modeled as a linear combination of $\sin(k\theta)$ and
242 $\cos(k\theta)$ terms (after application of a phase shift that depends on k), where $k = 0, 1, \text{ or } 2$. The $k =$
243 0 term, with no backazimuthal dependence, suggests an isotropic velocity change across a flat
244 interface; the $k = 1$ terms imply a dipping interface and/or a dipping axis of anisotropic
245 symmetry; the $k = 2$ terms correspond to contrast in anisotropy. The harmonic stacking technique
246 is illustrated for example station SSPA in Figure 3, which shows the modeled harmonic
247 decomposition as a function of time (along with an error estimate derived from bootstrap
248 resampling; see Ford et al., 2016 for details). Figure 3 also shows the unmodeled portion of the
249 signal, which cannot be represented as a linear sum of $\sin(k\theta)$ and $\cos(k\theta)$ terms. Here the
250 “unmodeled” components for each value of k represent the portion of the signal that has the
251 opposite phase shift as that of the harmonic expansion model (Shiomi and Park, 2008).

252

253 RESULTS

254 For each of the six stations examined in this study, we computed stacked radial RFs over
255 all backazimuths for two different lowpass frequency cutoffs (Figure 4) as well as backazimuthal
256 and epicentral distance gathers (Supplementary Figures S1-S12). Here we focus our presentation
257 of the results and discussion of our interpretation on the single-station stacked radial RFs (Figure
258 4) as well as the harmonic decomposition results for each station, grouped by region (Grenville
259 stations in Figure 5 and Mountain stations in Figure 6). We use the harmonic decomposition
260 modeling to guide our interpretation of major interfaces and their (an)isotropic character beneath
261 each station, as shown in Figures 7 and 8 and discussed in the next section. While we do not

262 discuss the epicentral distance gathers in detail, for each of the interfaces that we interpret in this
263 paper, we have checked the epicentral distance gathers (odd numbered Figures S1-S12) to ensure
264 that these arrivals do not exhibit moveout that suggests that they are contaminated by multiply
265 reflected waves from the crust. We note that while our use of the LQT coordinate system should
266 theoretically remove the direct P arrival from our RF stacks, some energy at or near $t = 0$ is often
267 visible (Figures 4 and S1-S12). This reflects minor inaccuracies in the velocity model used to
268 predict the geometry of the incident P wave; an offset from $t = 0$ in this arrival (e.g., stations
269 BINY and MCWV in Figure 4) may reflect the presence of thick sedimentary cover beneath the
270 stations or (more likely) shallow intracrustal interface(s).

271 The stacked radial RFs shown in Figure 4, which reflect events arriving across the full
272 backazimuthal range and which thus obscure any dipping or anisotropic structure, yield evidence
273 for complex structure with a number of interfaces beneath each of the stations examined in this
274 study. At our lower frequency range (lowpass cutoff at 0.5 Hz; Figure 4a), we observe at all
275 stations a strong, clear positive pulse (light blue lines in Figure 4) around 5-5.8 seconds after the
276 P arrival, which we interpret as corresponding to the positive (isotropic) velocity change across
277 the Moho. This time interval corresponds to a crustal thickness of around ~38-48 km for
278 reasonable velocity models for continental crust, and is generally consistent with previously
279 published estimates of crustal thickness in our study region (e.g., Abt et al., 2010). In addition to
280 the Moho, we also see evidence for (typically two or more) negative pulses (corresponding to a
281 decrease in velocity with depth) at later time ranges (roughly 8 s and greater; pink lines in Figure
282 4), which may correspond either to mid-lithospheric discontinuities (MLDs; Abt et al., 2010;
283 Ford et al., 2016) or, perhaps, to the base of the lithosphere. When higher frequency energy (up

284 to 1 Hz) is included, the radial RF stacks (Figure 4b) have similar character but more detail,
285 often hinting at intracrustal layering (pulses arriving before the Moho conversion).

286 Figures 5 and 6 show the results of the harmonic decomposition analysis for all stations,
287 computed for our low-frequency data (cutoff frequency of 0.5 Hz, same as the RFs shown in
288 Supplementary Figures S1-S12). Focusing first on the Grenville stations (ACSO, BINY, and
289 ERPA; Figure 5), we find that the harmonic stacking approach is generally able to model most of
290 the features of the RFs (note relatively small amplitudes on the “unmodeled” stacks) and yields
291 evidence for a number of interfaces with different character. Beneath station ACSO, the Moho is
292 expressed as a large-amplitude positive pulse on the constant term (at $t = 5.8$ s relative to direct
293 P); in the same time window, there is some energy (above the error estimate, shown on each
294 trace) on the $k = 1$ and $k = 2$ terms, although it is weaker. Still, the presence of energy on the $k =$
295 1 and $k = 2$ terms arriving at the same time as the constant Moho arrival suggests a contrast in
296 anisotropy across the Moho, perhaps with a dipping symmetry axis, and/or a dipping Moho
297 interface. We also see evidence on the constant term stack for an isotropic velocity drop at a
298 depth internal to the mantle lithosphere ($t = 9.5$ s, corresponding to roughly 90 km based on the
299 ak135 Earth model of Kennett et al., 1995); the $k = 1$ and $k = 2$ terms in this time range suggest a
300 contrast in anisotropy and/or a dip to this interface as well. This inferred MLD phase does not
301 show a clear moveout in the epicentral distance stacks (Figure S2), which argues against
302 significant contamination from crustal multiples; a similar argument can be made for each of the
303 mantle interfaces we discuss in this paper, with one exception noted below.

304 Returning to the harmonic decomposition analysis in Figure 5 and focusing purely on the
305 $k = 1$ and $k = 2$ components, we can identify peaks corresponding to interfaces at mantle depths
306 beneath ACSO that suggest a contrast in anisotropy (perhaps with a dipping symmetry axis), but

307 do not necessarily correspond to high amplitudes on the constant term (meaning an isotropic
308 velocity change is not required). In the time range associated with the direct P arrival, the
309 transverse component RFs show little or no energy at most backazimuths (Figure S1); we do not
310 observe any convincing zero-time arrivals that might result from a dipping interface, so contrasts
311 in anisotropy with a plunging symmetry axis may be more likely. We emphasize, however, that
312 the presence of multiple interfaces with large $k = 1$ terms in the harmonic expansion implies that
313 an examination of the zero-time transverse component arrivals to distinguish between dipping
314 interfaces and plunging anisotropy is not straightforward.

315 Beneath station BINY (Figure 5), the general character of the inferred interfaces is
316 similar to what we infer beneath ACSO, but the details are different. In the time window
317 corresponding to the Moho pulse ($t = 5.2$ s) on the constant term, there is clear energy on both
318 the $\sin(2\theta)$ and $\cos(2\theta)$ terms, again suggesting a contrast in anisotropy across the Moho. At later
319 times, corresponding to depths internal to the mantle lithosphere, we again see a number of
320 interfaces evident on both the constant and the higher order ($k = 1, 2$) terms, but the exact timing
321 of these arrivals, and the anisotropic and/or dipping geometry suggested by relative strength of
322 the $\cos(\theta)$, $\sin(\theta)$, $\cos(2\theta)$, and $\sin(2\theta)$ stacks, are different than for station ACSO. Similarly,
323 beneath ERPA, we see the same general pattern, with both isotropic and anisotropic/dipping
324 interfaces suggested by the harmonic decomposition, but with different specific characteristics.
325 Strikingly, ERPA exhibits an interface at a time of ~ 7.8 s (relative to direct P ; roughly 75-80 km
326 depth based on ak135) that shows both a strong (negative) constant component and a strong
327 $\cos(\theta)$ component. This strong $k = 1$ term may correspond to either a dipping MLD or a contrast
328 in anisotropy across the MLD that involves a plunging axis of anisotropy; the very weak signal at
329 zero delay time (relative to P) on the transverse component RFs for this station (Figure S5)

330 suggests that the latter possibility is more likely. There is another strong negative arrival on the
331 constant stack ($t = 12$ s; roughly 120 km depth) that also shows large amplitudes for $\cos(\theta)$,
332 $\sin(\theta)$, and $\sin(2\theta)$, again suggesting a contrast in anisotropy with a plunging symmetry axis.

333 Our harmonic stacking results for the Mountain stations (Figure 6) again reveal a series
334 of complicated interfaces, many with dipping and/or anisotropic character, beneath each station.
335 At each of the three stations, we observe a clear Moho arrival on the constant term, and at two of
336 the three stations (SSPA and TZTN), the higher order stacks also exhibit clear peaks in the same
337 time window, suggesting an anisotropic and/or dipping character to this interface. (At MCWV,
338 there is an anisotropic intracrustal interface whose conversions arrive roughly one second ahead
339 of the Moho P_s phase, but it appears to be distinct from the Moho itself.) The constant terms at
340 each of the three stations at time windows corresponding to mantle depths suggest one or more
341 interfaces with a velocity drop, implying the likely presence of multiple MLDs; typically, these
342 are also associated with an anisotropic or dipping character, as revealed by the higher order terms.
343 For the most part, the epicentral distance gathers argue against contamination of crustal multiples
344 in our interpretation of mantle interfaces beneath these stations, although we do observe some
345 moveout for the negative arrival at times near $t = 9$ sec for station TZTN (Figure S12); its
346 interpretation should therefore be treated with some caution. An examination of the transverse
347 RFs gathers at $t = 0$ for the Mountain stations (Figures S7, S9, and S11) reveals modest amounts
348 of energy, with complex backazimuthal patterns. For each of the three stations, there is nonzero
349 energy at $t = 0$, suggesting the presence of dipping interface(s) at depth, but the azimuthal
350 variability is not clear enough to identify individual two-lobed polarity changes and associate
351 them with specific interfaces.

352 As with the Grenville stations, we are also able to pick out features on the $k = 1$ and $k = 2$
353 stacks at individual Mountain stations that do not obviously correspond to features on the
354 constant stacks, suggesting that there is some anisotropic layering with the lithosphere that is not
355 accompanied by an isotropic change in velocity. Again as with the Grenville stations, the
356 collection of harmonic stacks for the Mountain stations, taken as a whole, suggest pronounced
357 differences in the details of both isotropic and anisotropic layering among different stations.
358 Visual inspection of the three stations shown in Figure 6 reveals major differences in the timing
359 and character (that is, the geometry of polarity reversals as revealed by the harmonic stacking
360 technique) of individual interfaces among different stations.

361

362 **INTERPRETATION AND DISCUSSION**

363 A key advantage of the harmonic stacking technique is that it allows us to identify
364 conversions associated with interfaces that exhibit a dipping and/or anisotropic character, a task
365 that is somewhat difficult relying only on visual inspection. From the harmonic stacks shown in
366 Figures 5 and 6, we have picked a series of time ranges (indicated by gray lines in Figures 5 and
367 6, with dashed lines for interfaces within the crust and solid lines for interfaces within the
368 mantle) that indicate dipping and/or anisotropic interfaces beneath each station. Our
369 methodology for picking these interfaces follows Ford et al. (2016) and involves a summation of
370 each of the four non-constant terms at each delay time in order to determine where coherent
371 peaks in energy occur. Another way of visualizing the character of these interfaces, which lends
372 itself well to comparison among different stations, is to plot rose diagrams (as illustrated in
373 Figure 3) for each of these time ranges, showing the variability in the transverse component RF
374 polarities and amplitudes as a function of backazimuth. In Figure 7, we show rose diagrams

375 interfaces at crustal depths for each station (either within the crust or co-located with the
376 isotropic Moho), grouped by tectonic setting (Grenville vs. Mountain stations). In Figure 8, we
377 show similar rose diagrams for interfaces within the mantle lithosphere; this figure also includes
378 a comparison with data from our previous study using stations within the Granite-Rhyolite
379 Province (Wirth and Long, 2014). We have also used the harmonic decomposition results to
380 calculate azimuthal phase information for individual interfaces (for both the $k = 1$ and $k = 2$
381 terms), which corresponds to the azimuthal location of polarity changes in the rose plots in
382 Figure 7 and 8 and which contains information about the possible dip or plunge directions for a
383 dipping interface or plunging symmetry axis (for the $k = 1$ components), and information about
384 contrasts in the orientation of horizontal azimuthal anisotropy (for the $k = 2$ components). These
385 orientations are shown in map view in Figures 9-11, and discussed in detail below.

386 Taken together, the collection of rose diagrams shown in Figures 7 and 8 illustrate that
387 our stations overlie highly complex lithosphere, with a number of interfaces both within the crust
388 and within the mantle lithosphere that dip and/or feature a contrast in anisotropy, perhaps with
389 dipping symmetry axes. This is true both at crustal depths (Figure 7) and within the mantle
390 lithosphere (Figure 8), and is similar to our previous findings within Archean cratonic regions of
391 North America (the Wyoming and Superior cratons; Ford et al., 2016). In that previous work, we
392 identified a surprising amount of both vertical and lateral heterogeneity in lithospheric structure,
393 suggesting that the complex tectonic processes which produce very heterogeneous structure in
394 the surface geology also produce similarly complex structure within the deep lithosphere. It is
395 strikingly evident from Figures 7 and 8 that there is significant lateral heterogeneity within the
396 crust and lithosphere in our eastern U.S. study region, even when we compare within groups of
397 stations. Stations that overlie lithosphere likely affected by Grenville orogenesis show little

398 similarity in the timing, character, or geometry of dipping/anisotropic structure across stations;
399 the same is true for the Mountain stations that overlie lithosphere likely affected by Appalachian
400 orogenesis.

401 A more detailed examination of the crustal interfaces illustrated in Figure 7 yields some
402 clues as to their geometry and possible origin. Beneath ACSO, we identify a dipping interface
403 within the crust, with conversions arriving ~ 3.9 s after the direct P arrival. The energy on the
404 transverse component RFs at $t = 0$ may result from this dipping interface, although the azimuthal
405 variability in its polarity is less clear-cut than the azimuthal variability in the P_s arrival at ~ 3.9
406 sec. Lacking a detailed crustal velocity model, we cannot accurately estimate the depth of this
407 interface, but its timing suggests mid-crustal depths. The arrival is modeled well with the $k = 1$
408 terms in the harmonic expansion (Figure 5), so anisotropy need not be invoked. Also beneath
409 ACSO, we see evidence for an anisotropic contrast across the Moho; in the time range (~ 5.8 s)
410 associated with the constant (isotropic) Moho arrival, the transverse component exhibits
411 significant azimuthal variability (Figure 7), with a nonzero $\sin(2\theta)$ term. Visually, the transverse
412 components associated with the Moho arrival look similar to those for the mid-crustal arrival
413 (upper left panels in Figure 7), but an additional anisotropic component is suggested by the
414 harmonic stacks. Beneath both BINY and ERPA, and similar to ACSO, there is variability on the
415 transverse component in the time range associated with Moho conversions (Figure 7) that
416 suggests the presence of a contrast in anisotropy (see non-zero $k = 2$ terms in this time range in
417 the harmonic stacks in Figure 5) as well as a dipping component. At all three Grenville stations,
418 then, we see evidence for a contrast in anisotropic properties between the lower crust and the
419 uppermost mantle, suggesting anisotropy in one or both layers.

420 Our inferences on crustal structure beneath the Mountain stations (Figure 7) include our

421 identification beneath station MCWV of two intracrustal interfaces (at ~ 2.2 s and ~ 4.3 s; both
422 time ranges are before the Moho arrival), each of which exhibit significant energy in both the $k =$
423 1 and $k = 2$ stacks (Figure 6) and thus require some contribution from anisotropic structure. This
424 observation strongly suggests at the presence of multiple anisotropic layers within the mid-to-
425 lower crust beneath MCWV, likely with dipping anisotropic symmetry axes (and/or with a dip to
426 the interfaces themselves). Similarly, beneath SSPA, we infer the presence of an interface in the
427 lowermost crust (arrival time ~ 4.3 s, just before the Moho arrival but with $k = 1$ and $k = 2$ peaks
428 that clearly arrive before the $k = 0$ Moho peak; Figure 6) that also requires a contrast in
429 anisotropy. The rose diagram for this interface (Figure 7) clearly shows four polarity changes
430 across the full backazimuthal range, although this pattern is modulated by $k = 1$ terms as well
431 (Figure 6). Beneath TZTN, the only identifiable crustal anisotropic contrast is clearly associated
432 with the isotropic Moho, suggesting a contrast between the lower crust and the uppermost mantle,
433 similar to what we observe beneath all three of the Grenville stations.

434 What are the implications of our findings about anisotropic layering at crustal depths for
435 our understanding of past tectonic deformation? First, our inference of a contrast in anisotropic
436 properties across the Moho at four of our stations (ACSO, BINY, ERPA, and TZTN) suggests
437 significant, “frozen-in” anisotropy in the relatively shallow portions of the lithosphere (lower
438 crust and/or uppermost mantle) due to deformation associated with past tectonic events. The
439 character of the transverse RFs in the time range associated with the Moho does not strictly
440 require anisotropy in the lowermost crust beneath these four stations, but it does imply
441 anisotropy in the deepest crust, the shallowest mantle, or both. In either case, our observations
442 suggest past deformation at depths near the Moho, with enough strain to develop coherent CPO
443 of anisotropic minerals. Second, our observations argue strongly for multiple anisotropic layers

444 within the crust itself, at mid-to-lower crustal depths, beneath two of our stations (MCWV and
445 SSPA), both located within the Appalachian Mountains. We interpret this finding as suggesting
446 that 1) the deep crust has undergone significant deformation in the past and has developed
447 anisotropy via the CPO of crustal minerals, most likely associated with Appalachian orogenesis,
448 and 2) the mineralogy, rheology, and/or deformation geometry varied as a function of depth
449 within the crustal column, resulting in multiple layers of crust with contrasting anisotropic
450 geometries. Our inference of layered crustal anisotropy in the ancient Appalachian orogeny is
451 similar to inferences from modern orogens such as Tibet (Liu et al., 2015) and Taiwan (Huang et
452 al., 2015), and provides additional support for the hypothesis that crustal deformation in orogenic
453 systems is complex, varies with depth, and extends to the deep crust.

454 Turning our attention now to the lithospheric mantle, we examine the character of
455 transverse component RFs at time ranges associated with anisotropic and/or dipping contrasts
456 within the lithospheric mantle, shown in Figure 8. Of particular interest are the possible
457 relationships between contrasts in anisotropy and the MLD(s) inferred from the single-station
458 radial RF stacks shown in Figure 3. The term “mid-lithospheric discontinuity,” coined by Abt et
459 al. (2010), refers to a sharp decrease in seismic velocity at a depth internal to the continental
460 mantle lithosphere. MLDs have been documented in a number of continental regions (e.g., Abt et
461 al., 2010; Ford et al., 2010; Foster et al., 2014; Hopper et al., 2014; Hopper and Fischer, 2015),
462 although their origin remains debated (e.g., Selway et al., 2015; Karato et al., 2015; Rader et al.,
463 2015). Several workers have suggested a link between anisotropic layering in the continental
464 lithosphere and the observation of MLDs (e.g., Yuan and Romanowicz, 2010; Sodoudi et al.,
465 2013; Wirth and Long, 2014). In particular, our previous work using data from the Granite-
466 Rhyolite province of North America (Wirth and Long, 2014) found evidence for a clear contrast

467 in anisotropy in the same depth range (~90 km) as the MLD, although our observations also
468 required a decrease in isotropic seismic velocity with depth and could not be explained solely by
469 anisotropic layering within the lithosphere.

470 Figure 8 demonstrates the presence of complex dipping and/or anisotropic layering
471 within the lithospheric mantle beneath the eastern U.S., with 2-3 clear interfaces that correspond
472 clearly to transverse RF polarity changes with backazimuth, arriving ~7-12 s after the P arrival,
473 identified at each station. This finding is similar to our previous work using data from the
474 Superior and Wyoming cratons in North America (Ford et al., 2016), where we typically
475 identified multiple layers of anisotropy within the mantle lithosphere. Comparing the rose
476 diagrams in Figure 8 to the single-station radial component RF stacks in Figure 3, we can
477 evaluate whether there is a correspondence (in converted wave arrival timing and thus in
478 interface depth) between isotropic MLD(s) (corresponding to red or negative pulses in Figure 3)
479 and contrasts in anisotropy and/or dipping interfaces (as inferred from the harmonic
480 decomposition and illustrated with rose diagrams in Figure 8). Beneath stations ACSO, ERPA,
481 MCWV, and SSPA, we are able to identify interfaces that apparently exhibit both an isotropic
482 drop in velocity with depth as well as an anisotropic or dipping character. For example, beneath
483 ACSO there is a clear discontinuity that produces a negative pulse on the stacked radial RFs at
484 9.5 s; the corresponding rose diagram in Figure 8 exhibits clear backazimuthal variations that
485 include an anisotropic component. Similarly, beneath SSPA there are two clear negative pulses
486 on the stacked radial RFs (at $t = 6.9$ s and 10.3 s) that also require a contrast in anisotropy, likely
487 with a dipping component (non-zero amplitudes in the $k = 1$ and $k = 2$ stacks in Figure 6).
488 However, again similar to the findings of Ford et al. (2016), we do not observe a simple, one-to-
489 one correspondence between the isotropic MLD arrivals and the anisotropic interfaces. Rather,

490 we observe complex layering within the lithosphere with both anisotropic and isotropic
491 interfaces, with only sporadic correlations between isotropic and anisotropic structure.

492 Constraints on the geometry of the RF polarity changes due to anisotropic and/or dipping
493 interfaces beneath our stations are visualized in map view in Figures 9-11, which show the phase
494 orientation information for the $k = 1$ and $k = 2$ terms in the harmonic expansion for a series of
495 interfaces at different depths. These directions are derived from measurements of the relative
496 amplitudes of the sine and cosine terms, and denote the possible orientations that may
497 correspond to the dip direction (for a dipping interface) or the plunge direction (for a plunging
498 symmetry axis) for the $k = 1$ terms, and contain information about contrasts across interfaces in
499 symmetry axis orientation (for azimuthal anisotropy) for the $k = 2$ terms. We caution that these
500 orientations do not uniquely constrain the geometry of anisotropy or of dipping interfaces; in the
501 absence of detailed forward modeling, assumptions must be made about the properties of
502 individual layers in order to estimate the anisotropic geometry. Nevertheless, the phase
503 information does contain some information about possible dip orientations and about the
504 geometrical relationships between azimuthally anisotropic symmetry directions across interfaces
505 (e.g., Shiomi and Park, 2008; Schulte-Pelkum and Mahan, 2014a,b; Olugboji and Park, 2016),
506 and regional variations in the possible orientations can be easily evaluated in map view.

507 Figure 9 shows the phase information for interfaces within the crust and across the Moho
508 for those stations where the Moho interface includes a dipping and/or anisotropic component.
509 This map reveals significant lateral variability in the character of crustal interfaces across the
510 study region, and little in the way of straightforward relationships between the phase orientations
511 and geologic and tectonic indicators. In Figure 10, we show phase orientations for interfaces
512 internal to the lithosphere that do not correspond to strong isotropic contrasts in velocity; rather,

513 these are shallow intralithospheric contrasts in anisotropy, typically with a dipping axis of
514 symmetry. As with the crustal interfaces in Figure 9, we observe a great deal of regional
515 variability. In Figure 11, we show orientations for interfaces in the mid-lithosphere that exhibit
516 both an isotropic drop in velocity with depth, thus corresponding to isotropic MLDs, as well as a
517 dipping and/or anisotropic character. There is some consistency between the $k = 1$ phase
518 orientations at stations ERPA, MCWV, and SSPA, two of which (ERPA and SSPA) exhibit
519 multiple anisotropic MLD interfaces. All of these have $k = 1$ orientations directed roughly NW-
520 SE, with more variability in the $k = 2$ phase information.

521 A direct comparison between our results and those of Wirth and Long (2014) is
522 instructive, particularly for stations ACSO and ERPA, which are located within the Granite-
523 Rhyolite province (ACSO) or in the adjacent and related Elzevir block (ERPA). Figure 8 shows
524 rose diagrams for arrivals in the time range between $\sim 7.8 - 10.5$ s (roughly corresponding to
525 depths between $\sim 80-105$ km) at Granite-Rhyolite stations; at each of these stations, Wirth and
526 Long (2014) identified evidence for layered anisotropy within the mantle lithosphere, with
527 conversions from one or more anisotropic interfaces arriving in the same time range as
528 conversions from the isotropic MLD. Furthermore, at three of the Granite-Rhyolite stations
529 (BWI, BLO, and CCM), the transverse component RFs associated with the main MLD interface
530 exhibited similar behavior, with four clear changes in polarity across the full backazimuthal
531 range. A forward model of the data at station WCI (Wirth and Long, 2014) invokes three
532 lithospheric layers of anisotropy; across the likely “main” MLD interface, the upper model layer
533 has a roughly N-S anisotropic fast direction and the lower layer is nearly E-W. This model
534 geometry predicts polarity changes similar to those observed at WCI (Figure 8, lower left), with
535 negative transverse component arrivals in the NE and SW quadrants and positive arrivals in the

536 NW and SE quadrants. A comparison of this specific geometric pattern with the transverse RF
537 behavior at stations ERPA and ACSO in similar time ranges (Figure 8, upper right) demonstrates
538 that the backazimuthal patterns are distinctly different beneath the Grenville stations in our study.
539 The fact that station ACSO in particular, which is also located within the Granite-Rhyolite
540 province (similar to nearby stations BLO and WCI from Wirth and Long, 2014) but to the east of
541 the Grenville Front, does not exhibit a similar geometric pattern implies that the lithosphere
542 beneath ACSO was modified via subsequent deformation. We suggest that this lithosphere was
543 deformed during Grenville orogenesis, thus modifying the preexisting lithospheric structure that
544 is still present elsewhere in the mantle of the Granite-Rhyolite province.

545 How do our observations compare to other previous inferences on the anisotropy of the
546 crust and mantle lithosphere beneath eastern North America? In general terms, our results are
547 consistent with previous work that has suggested a significant contribution to seismic
548 observations from anisotropy in the mantle lithosphere, based on surface waves (e.g., Deschamps
549 et al., 2010), *SKS* splitting patterns (e.g., Long et al., 2016), or on combinations of different types
550 of data (e.g., Yuan and Romanowicz, 2010; Yuan and Levin, 2014). Recent work on *Pn*
551 velocities and anisotropy beneath the continental U.S. (Buehler and Shearer, 2017) also suggests
552 significant anisotropy in the uppermost mantle beneath our study region; Buehler and Shearer
553 (2017) further propose that there must be multiple layers of anisotropy within the upper mantle,
554 potentially consistent with our observations of anisotropic layering within the mantle lithosphere.
555 Our finding of significant anisotropic layering with the mid-to-lower crust beneath at least some
556 of our stations is also generally consistent with results of Schulte-Pelkum and Mahan (2014a),
557 who suggested crustal anisotropy (often with a plunging symmetry axis) beneath a number of
558 USArray Transportable Array (TA) stations in the eastern U.S., particularly in the Appalachians

559 (see their Figure 8). Interestingly, our finding that the layered anisotropic structure of the mantle
560 lithosphere differs across the Grenville deformation front contrasts with the view provided by
561 shear wave splitting observations, which reflect an integrated signal from the entire upper mantle.
562 For example, Sénéchal et al. (1996) observed similar *SKS* splitting on either side of the Grenville
563 Front in Canada, while Long et al. (2016) noted that there is no obvious correlation between *SKS*
564 splitting patterns measured at TA stations and the geometry of the Grenville Front in the eastern
565 and central U.S.

566 A major limitation of the anisotropic RF technique is that while we can confidently infer
567 the presence (and roughly estimate the depths) of anisotropic interfaces within the lithosphere, as
568 well as some orientation information from the $k = 1$ and $k = 2$ phase values from the harmonic
569 stacking, it is difficult to infer the geometry of anisotropy within each layer without detailed
570 forward modeling. Such forward modeling is computationally intensive and non-unique, with a
571 large number of unknown parameters, so while it can identify plausible models, it often cannot
572 uniquely determine the anisotropy geometry in every layer. These difficulties can be ameliorated
573 by the use of model space search approaches (e.g., Porter et al., 2011; Wirth et al., 2017) or
574 Bayesian inversion schemes (e.g., Bodin et al., 2016), but there are still strong tradeoffs among
575 parameters. In lieu of detailed forward modeling, the harmonic decomposition modeling applied
576 here (or approaches similar to it; e.g., Schulte-Pelkum and Mahan, 2014a) can identify interfaces
577 and provide some first-order information about their geometry by examining the relative
578 amplitudes and phases of the $k = 0,1,2$ terms, and provides a natural way of comparing structure
579 among different stations. Furthermore, the RF observations and the harmonic decomposition
580 models presented in this paper can serve as a starting point for future detailed forward modeling,
581 ideally in combination with other data that constrain seismic anisotropy (such as *SKS* splitting

582 observations, surface wave dispersion, and/or Pn traveltimes). One example of this type of
583 modeling is discussed by Yuan and Levin (2014), who combined different types of anisotropy
584 observations, including backazimuthal RF gathers, into a forward model of multilayered
585 anisotropy beneath three stations in the eastern U.S. (including SSPA, one of the stations
586 examined in this study).

587 Despite the limitations of the RF technique, the observations presented in this paper allow
588 us to draw some straightforward inferences about lithospheric deformation beneath the eastern
589 U.S. and its likely causes. Our identification of extensive anisotropy in the deep crust and mantle
590 lithosphere beneath each of the stations examined in this study implies extensive lithospheric
591 deformation, with enough strain to generate significant CPO. While it is not possible to
592 conclusively identify the deformation events and their timing and geometry, we suggest that the
593 last major orogenic cycle to affect each region is the most plausible tectonic event to have caused
594 widespread lithospheric deformation (e.g., Meissner et al., 2002). Therefore, we propose that the
595 anisotropic structure of the mantle lithosphere beneath ACSO, ERPA, and BINY was shaped by
596 deformation associated with the Grenville orogenic cycle, while Appalachian orogenesis caused
597 lithospheric deformation beneath SSPA, MCWV, and TZTN. Our observations strongly suggest
598 anisotropic layering within the crust beneath two of the three Appalachian stations examined,
599 with an anisotropic geometry that varies with depth. This finding is similar to the documentation
600 of crustal anisotropy in modern mountain belts (e.g., Liu et al., 2015; Huang et al., 2015) and
601 suggests complex crustal deformation accompanying orogenesis. Our main finding of extensive
602 but layered deformation within the lithospheric mantle (at depths up to ~100 km) presents an
603 interesting challenge to the classic idea of vertically coherent deformation (Silver, 1996).
604 Although our observations suggest that the geometry of anisotropy (and thus deformation)

605 changes with depth, they also suggest significant deformation throughout much of the
606 lithosphere associated with orogenesis. This is consistent with the suggestion by Silver (1996)
607 that the lithosphere participates in deformation, leading to seismic anisotropy, but the finding of
608 changes in anisotropic geometry with depth suggests that the geometry of deformation and/or of
609 olivine fabric development is not coherent over the entire lithospheric mantle.

610 Our identification of lithospheric anisotropy beneath eastern North America is consistent
611 with observations in other continental regions, including those that have undergone (or are
612 currently undergoing) orogenesis (e.g., Meissner et al., 2002; Wüstefeld et al., 2010). Our main
613 conclusions are also consistent with the many previous suggestions that *SKS* splitting in and
614 around the Appalachians, which is typically parallel to the mountain belt (at least in the central
615 and southern portions), reflects a significant contribution from lithospheric anisotropy (e.g.,
616 Barruol et al., 1997; Wagner et al., 2012; Long et al., 2016). We note, finally, that our inference
617 of multiple layers of anisotropy within the mantle lithosphere is generally consistent with the
618 surface wave model of Deschamps et al. (2008) for a region just to the south and west of our
619 study area. Although there is no substantial geographic overlap between their study region and
620 the stations used in this study, Wirth and Long (2014) did discuss the similarities and differences
621 between their RF-derived model and the results of Deschamps et al. (2008).

622

623 **SUMMARY**

624 We have examined data from six long-running broadband seismic stations located within
625 regions of the eastern U.S. that were affected by the Grenville and Appalachian orogenic cycles.
626 We computed radial and transverse RFs and applied a harmonic stacking method that allows us
627 to identify interfaces within the crust and lithospheric mantle that have a dipping and/or

628 anisotropic character, in addition to interfaces that can be explained in terms of an isotropic
629 velocity contrast. We find evidence for layered anisotropy, often with a likely dipping axis of
630 symmetry, beneath all stations examined in this study. There is often a clear anisotropic contrast
631 associated with the Moho interface, requiring anisotropy in the lowermost crust, the uppermost
632 mantle, or both. Similar to recent findings within Archean regions of North America, we find
633 evidence for complex anisotropic layering with the mantle lithosphere. At two of the stations in
634 the Appalachian Mountains, we observe clear evidence for anisotropic contrasts within the crust,
635 suggesting mid-to-lower crustal deformation associated with Appalachian orogenesis. Finally, a
636 detailed comparison between our observations and previous anisotropic RF analysis at stations
637 located in the Granite-Rhyolite province but to the west of the Grenville Front suggests a
638 different geometry of anisotropy in the mantle lithosphere. This, in turn, suggests that the
639 lithospheric mantle just to the east of the Grenville Front was deformed during the Grenville
640 orogenic cycle. Our observations can provide a starting point for the future testing of detailed
641 deformation scenarios associated with past tectonic events, ideally in combination with other
642 types of data that constrain seismic anisotropy.

643

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653

654 **FIGURE CAPTIONS**

655 **Figure 1.** Map of station locations and major tectonic boundaries. Red triangles indicate
656 locations of stations used in this study; blue triangles show stations examined by Wirth and Long
657 (2014). Red lines indicate the boundaries of major Proterozoic terranes (Yavapai, Mazatzal,
658 Granite-Rhyolite, and Grenville), according to Whitmeyer and Karlstrom (2007). Dashed line
659 indicates the position of the Grenville deformation front, and the black line indicates the western
660 boundary of the Elzevier block, also from Whitmeyer and Karlstrom (2007).

661

662 **Figure 2.** Map of earthquakes used in our analysis at station SSPA. Station location is shown
663 with a triangle; event locations are shown with circles. Event distributions for other stations in
664 the study are similar.

665

666 **Figure 3.** Example of different stacking and plotting conventions for station SSPA. (A) Single
667 station stacked, radial component P_s receiver function. The Moho pick is shown with a cyan line,
668 while conversions that may correspond to potential MLDs are shown with magenta lines. (B)
669 Radial (top) and transverse (bottom) component receiver functions binned as a function of back
670 azimuth. Cyan and magenta lines on the top panel correspond to the arrivals picked in panel A.
671 Gray lines on the bottom panel correspond to anisotropic interfaces selected from the regression
672 in (C). (C) Harmonic expansion results. Top panel corresponds to the modeled portion of the
673 harmonic expansion, with (from right to left) the $k = 0$, $k = 1$, and $k = 2$ terms shown (as labeled

674 on the bottom of the plot). The bottom panel corresponds to the portion of the P_s receiver
675 function that cannot be modeled by the harmonic expansion. Receiver functions are plotted as a
676 function of delay time relative to the P wave arrival time. A target depth of 90 km was used in
677 the harmonic expansion, and the 90 km and 0 km marks are drawn as horizontal black lines in all
678 panels. Horizontal gray lines in the modeled stacks correspond to interfaces whose presence is
679 inferred from the harmonic stacking results, with numbers showing their arrival times relative to
680 the P wave arrival. Dashed gray lines indicate interfaces at crustal depths; solid gray lines
681 indicate mantle interfaces. Cyan line shows the Moho pick, as in panel A. Magenta bars show the
682 time range associated with likely MLD arrivals, as in panel A. (D) Example transverse
683 component P_s receiver function rose plots for inferred anisotropic boundaries at 6.9 and 10.3
684 seconds. Data is shown on the left of each time increment, while the associated model from the
685 harmonic expansion is shown on the right. We note the similarity between our backazimuthal RF
686 gathers for station SSPA (panel B) and those presented in Yuan and Levin (2014) (their Figure 7),
687 who also examined data from this station.

688

689 **Figure 4.** Station-stacked, radial component P_s receiver functions shown for all stations, filtered
690 at 0.5 Hz (left) and 1.0 Hz (right). Y-axis for each panel is delay time (relative to direct P arrival)
691 in seconds. Blue phases indicate positive amplitudes and an inferred velocity increase with
692 depth; red indicates negative amplitudes and an inferred velocity decrease with depth. Station
693 names are shown on the far left. The Moho picks are shown in cyan and the negative picks
694 interpreted as either MLDs or LAB are shown in magenta; dashed lines indicate interfaces with
695 smaller amplitudes.

696

697 **Figure 5.** Ps receiver functions binned as a function of harmonic expansion terms for Grenville
698 Province stations ACSO, BINY and ERPA. Top panels correspond to the modeled portion of the
699 harmonic expansion. The bottom panels correspond to the portion of the Ps receiver functions
700 which cannot be modeled via harmonic expansion. Both the modeled and unmodeled receiver
701 functions are plotted as a function of delay time relative to the direct *P* wave arrival. A target
702 migration depth of 90 km was used in the harmonic expansion (e.g., Bianchi et al., 2010; Ford et
703 al., 2016), and the 90 km and 0 km marks are drawn as horizontal black lines in all panels. Gray
704 solid lines mark the location of significant anisotropic boundaries within the mantle and are
705 labeled with the delay time relative to the direct *P* arrival. Gray dashed lines mark the location of
706 anisotropic boundaries at or above the Moho. Rose plots that display the transverse component
707 RF energy as a function of backazimuth for the time window associated with these boundaries
708 are shown in Figures 7 and 8. Magenta bars correspond to approximate MLD and/or LAB arrival
709 delay times (± 0.5 seconds). The cyan line corresponds to the Moho arrival.

710

711 **Figure 6.** Ps receiver functions binned as a function of harmonic expansion terms for
712 Appalachian stations MCWV, SSPA and TZTN. Plotting conventions are as in Figure 5. Rose
713 plots that display the transverse component RF energy as a function of backazimuth for the time
714 window associated with the boundaries marked in gray are shown in Figures 7 and 8.

715

716 **Figure 7.** Transverse component Ps receiver function rose plots for inferred anisotropic
717 boundaries within the crust, as inferred from the harmonic expansion shown in Figures 5 and 6.
718 Stations are grouped according to location (Grenville or Mountain stations). Each individual rose
719 plot shows the transverse component amplitudes as a function of backazimuth at the given delay

720 time. Individual rose plots were normalized such that the maximum amplitude within each plot is
721 equal to one.

722

723 **Figure 8.** Transverse component Ps receiver function rose plots for inferred anisotropic
724 boundaries within the mantle arriving at ~ 7 sec time delay and later, as inferred from the
725 harmonic expansion shown in Figures 5 and 6. The plotting convention is the same as in Figure 7.
726 Rose plots are grouped according to location (Grenville, Granite-Rhyolite, and Mountain
727 stations), and include results for stations within the Granite-Rhyolite Province from our previous
728 work (Wirth and Long, 2014).

729

730 **Figure 9.** Summary map of the properties of crustal interfaces, as derived from the harmonic
731 expansions shown in Figures 5 and 6. Solid lines show the azimuthal phase information (that is,
732 the azimuthal location of polarity changes in the transverse component RFs) for the $k = 1$ terms
733 in the harmonic expansion, which correspond to possible dip directions (for a dipping interface)
734 or plunge directions (for a dipping anisotropic symmetry axis). Dotted lines show the azimuthal
735 phase information for the $k = 2$ terms in the harmonic expansion (for interfaces with a strong $k =$
736 2 component), which are related to changes in symmetry axis orientations for contrasts in
737 azimuthal anisotropy. Note the 90° ambiguity in the phase orientations for the $k = 2$ terms,
738 reflecting the four-lobed polarity flip with backazimuth on the transverse component RFs. Blue
739 lines indicate intracrustal interfaces, while black lines correspond to dipping and/or anisotropic
740 contrasts across the Moho. Each of the $k = 1$ orientations is marked with the arrival time of the Ps
741 converted phase (relative to direct P wave arrival). Note that beneath station MCWV, evidence
742 for a dipping and/or anisotropic component to the Moho interface is ambiguous; however, this

743 station overlies two clear intracrustal interfaces with different geometries, as shown in Figure 7.
744 Similarly, we infer the presence of an anisotropic interface within the deep crust beneath station
745 SSPA. For station MCWV, we have also labeled arrival times for $k = 2$ components (in
746 parentheses) to distinguish the multiple interfaces. Station ACSO exhibits an intracrustal
747 interface that is well described with only a $k = 1$ component.

748

749 **Figure 10.** Summary map of the properties of shallow intralithospheric interfaces as derived
750 from the harmonic expansion. Here we plot anisotropic contrasts within the shallow mantle
751 lithosphere that do not correspond to isotropic velocity drops (those interfaces are shown in
752 Figure 11). Stations ERPA and SSPA do not overlie any such interfaces. As in Figure 9, solid
753 lines show the azimuthal phase information for the $k = 1$ terms in the harmonic expansion, while
754 dotted lines show the azimuthal phase information for the $k = 2$ terms in the harmonic expansion.
755 Each of the $k = 1$ orientations is marked with the arrival time of the P_s converted phase (relative
756 to direct P wave arrival).

757

758 **Figure 11.** Summary map of the properties of mid-lithospheric interfaces as derived from the
759 harmonic decomposition. Here we only show those interfaces that correspond to both an
760 isotropic velocity decrease (that is, that have a strong negative $k = 0$ term in the harmonic
761 expansion) and to a contrast in anisotropy and/or a dip, as inferred from the $k = 1$ and $k = 2$ terms.
762 Plotting conventions are as in Figure 10. Note that stations BINY and TZTN do not exhibit such
763 interfaces, while stations ERPA and SSPA both exhibit multiple interfaces. for ERPA, one of
764 those (arrival time at 7.8 sec) is described well with only $k = 1$ terms, so no $k = 2$ terms are

765 shown. At SSPA and ERPA, we label the *Ps* arrival time for $k = 2$ components in parentheses to
766 distinguish the multiple interfaces.

767

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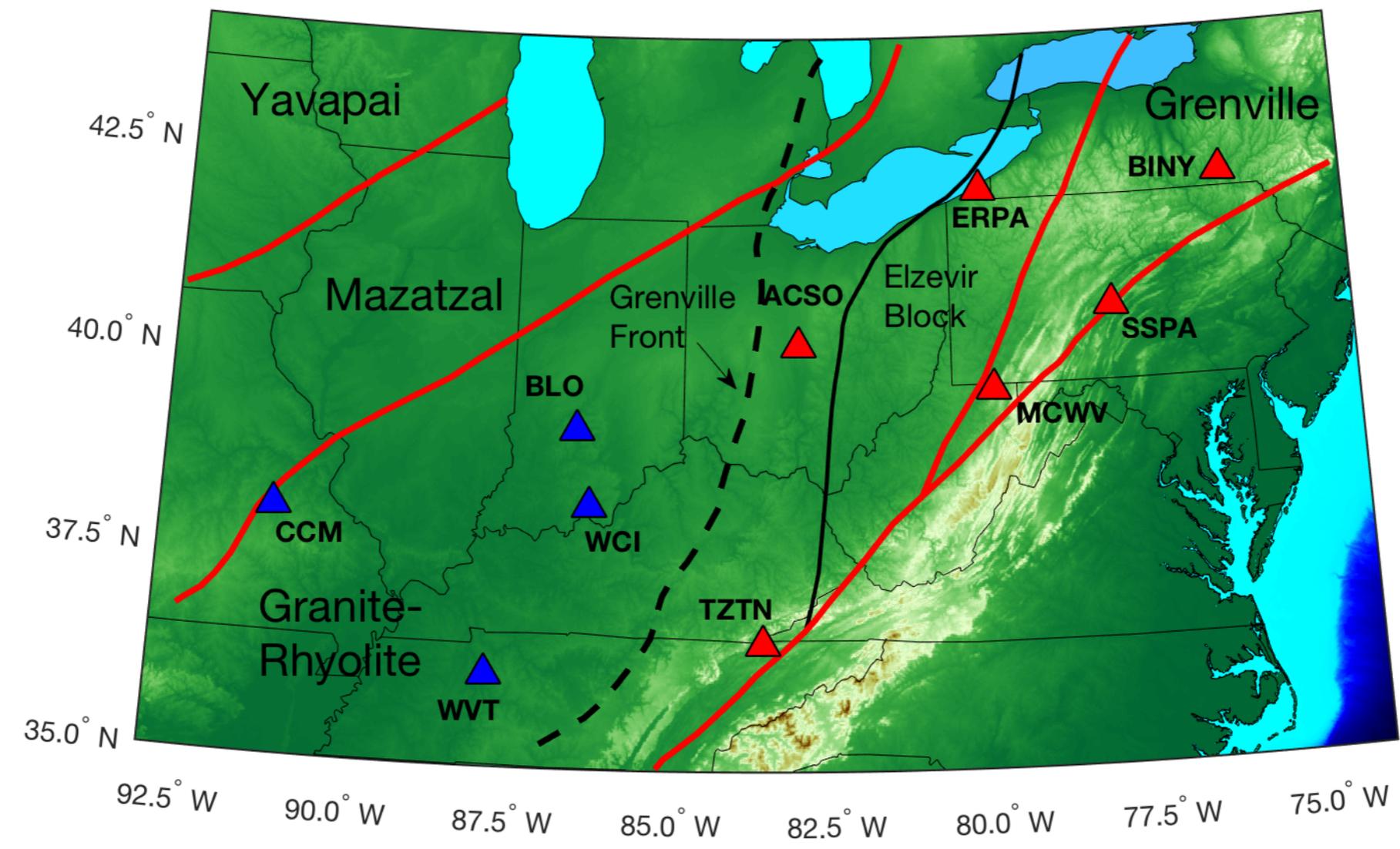


Figure 2

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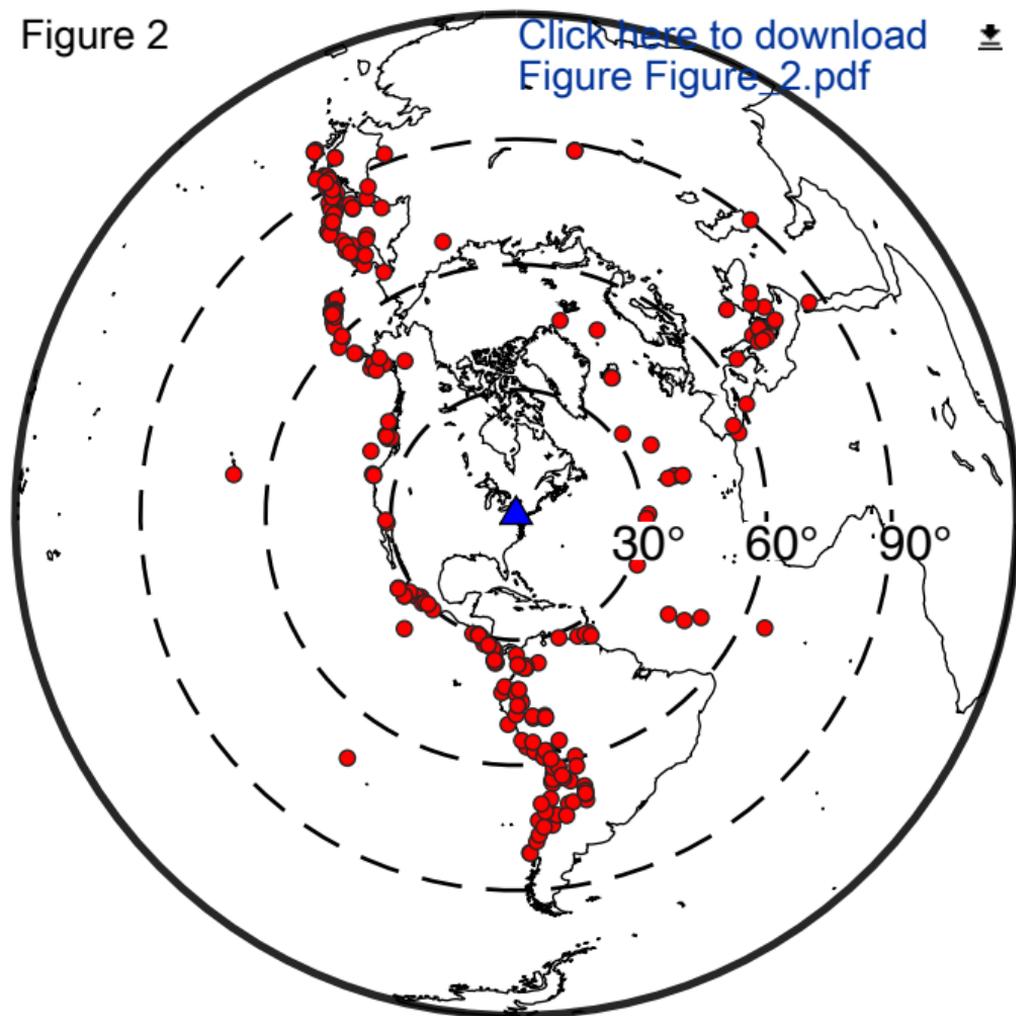
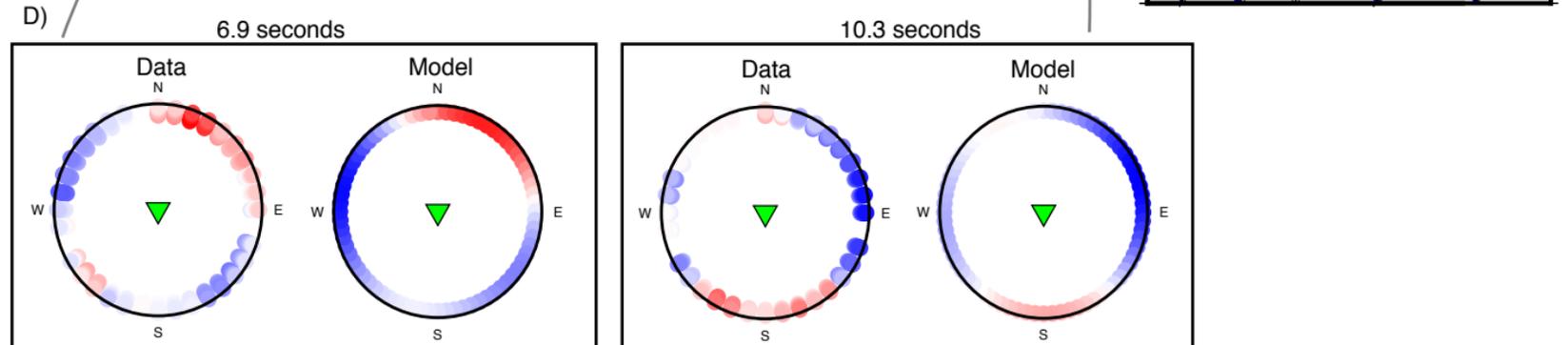
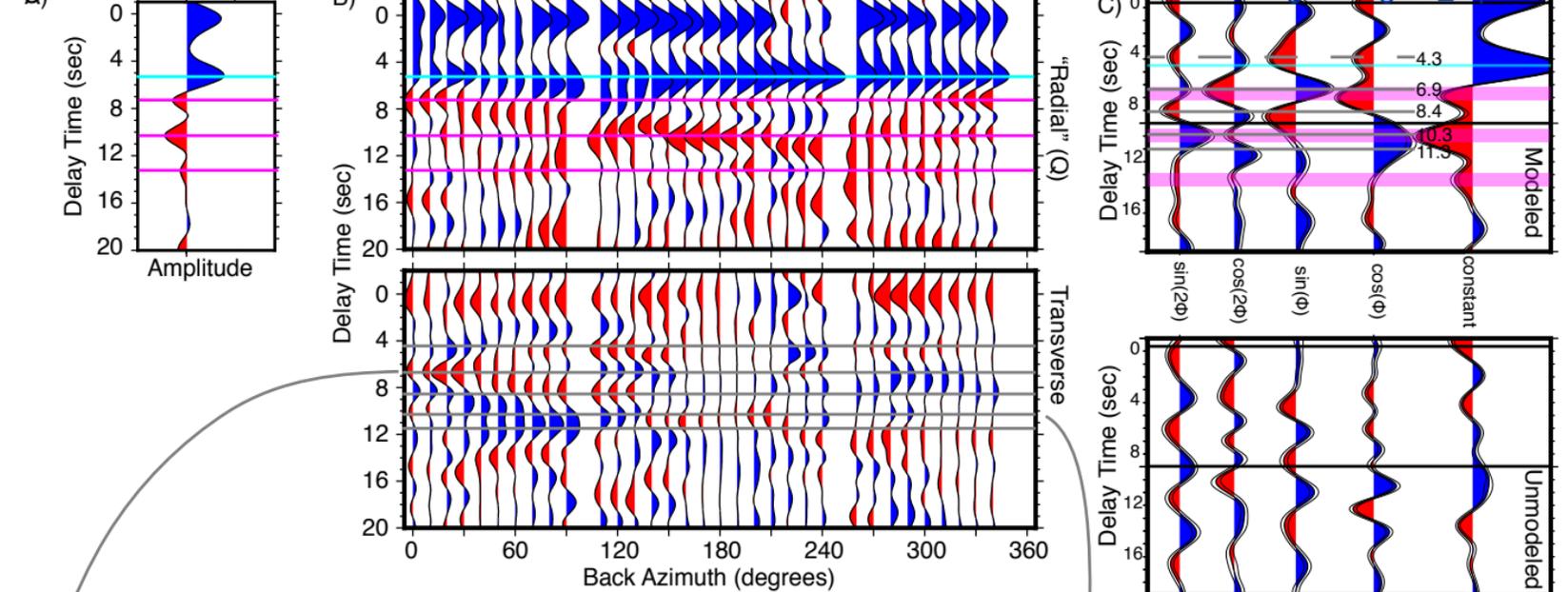


Figure 3



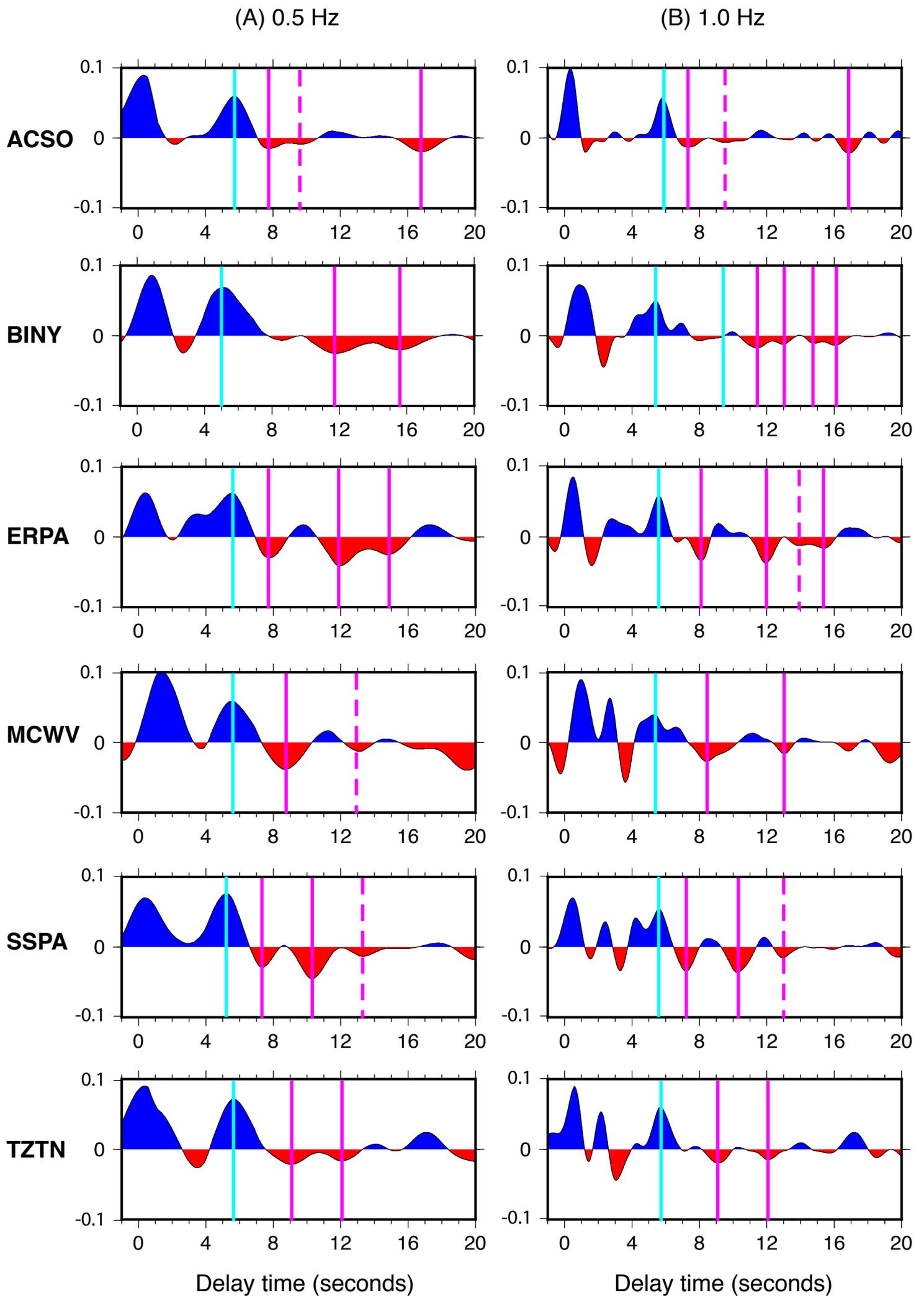


Figure 5

ACSO

BINY

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ERPA

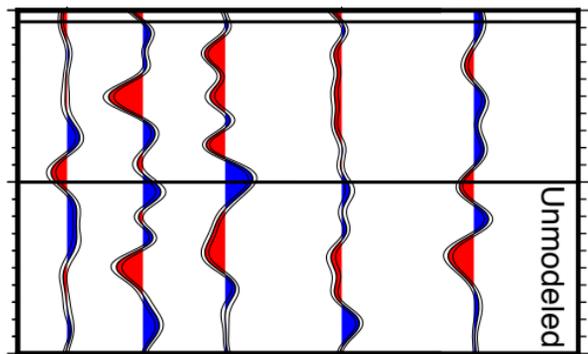
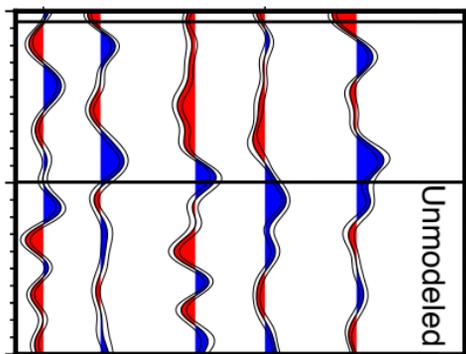
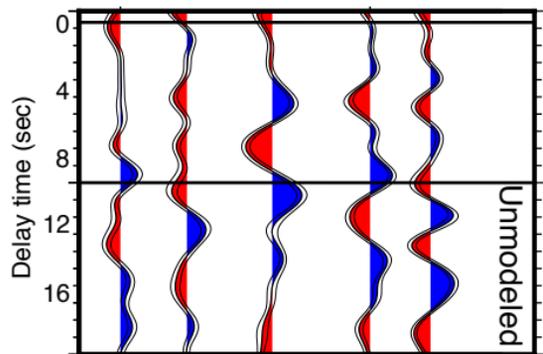
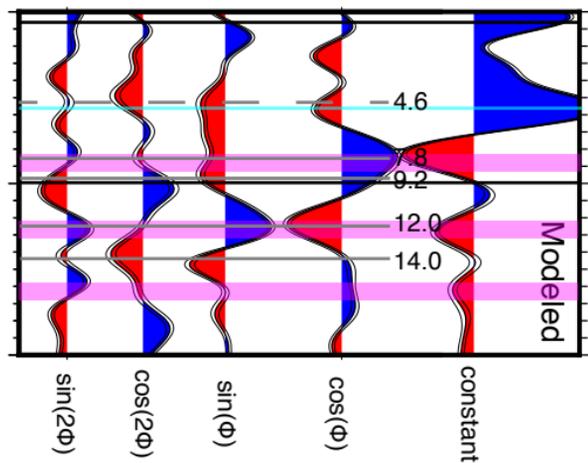
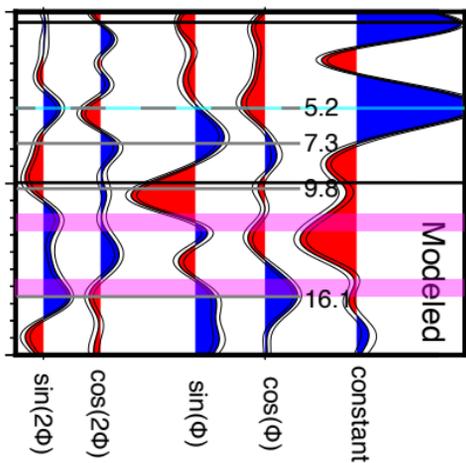
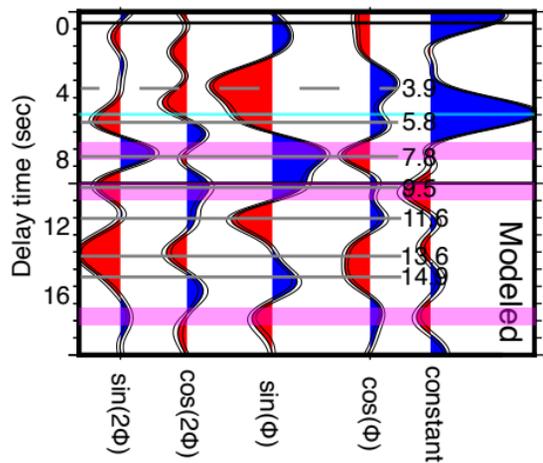


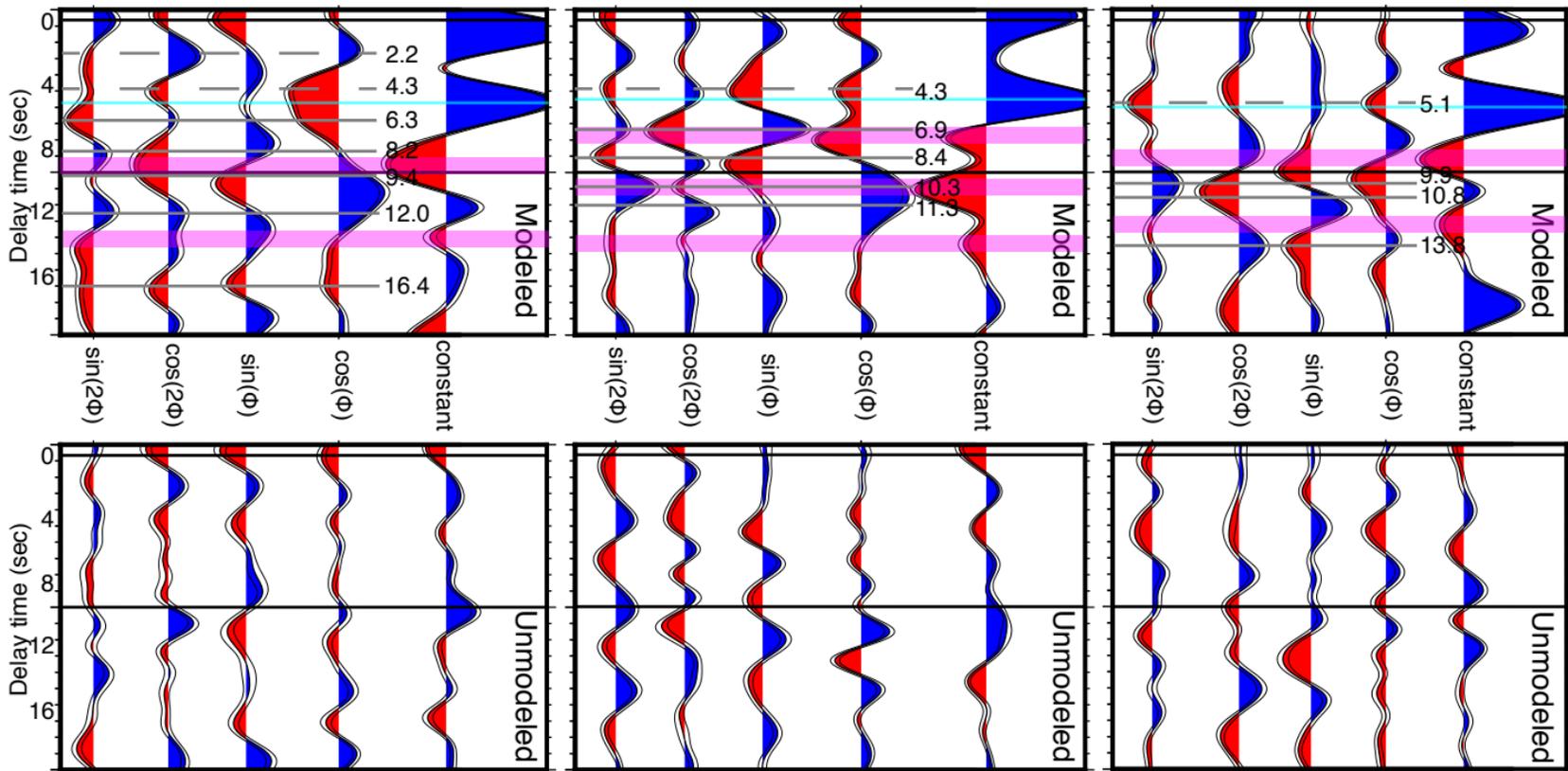
Figure 6

MCWV

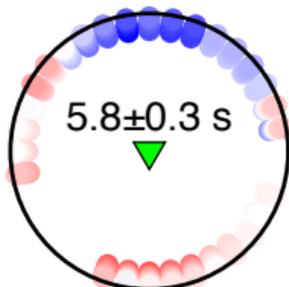
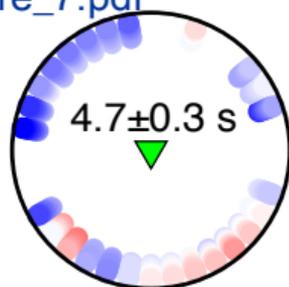
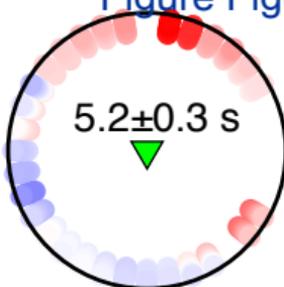
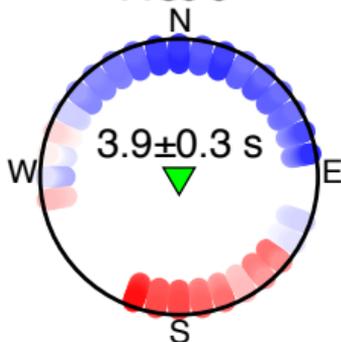
SSPA

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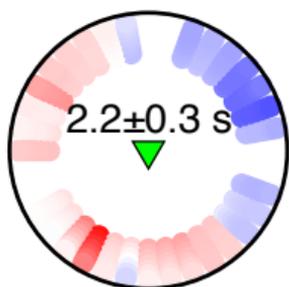
IZIN



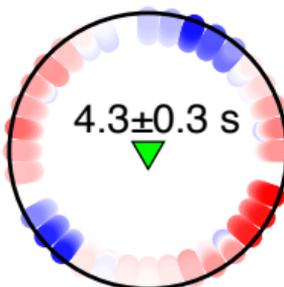
Grenville Stations



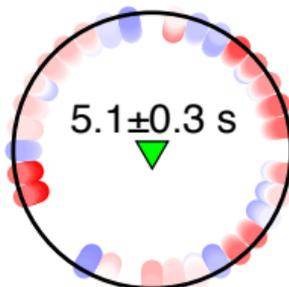
MCWV



SSPA



TZTN



Mountain Stations

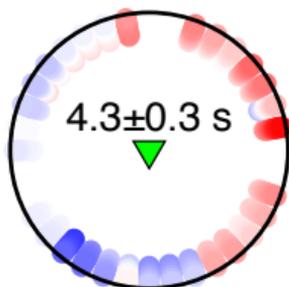
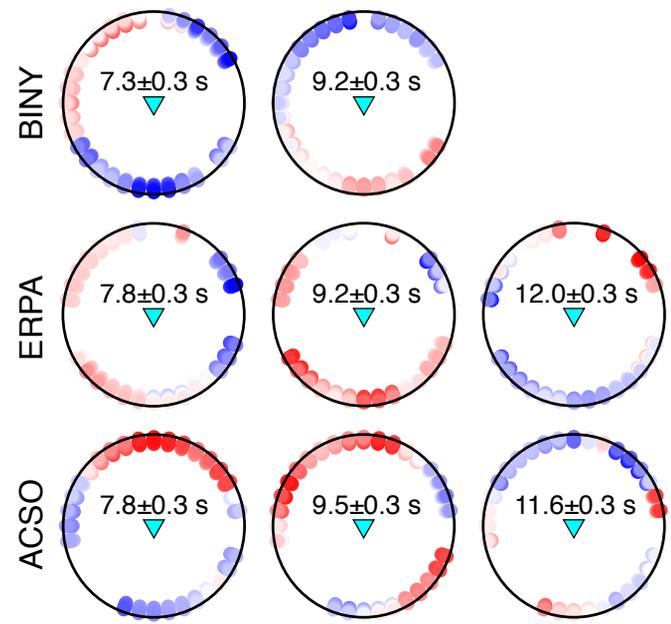
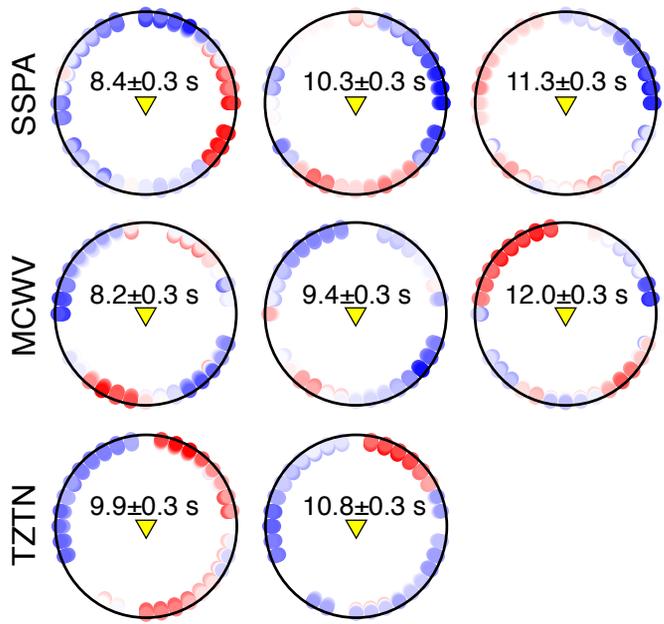


Figure 8

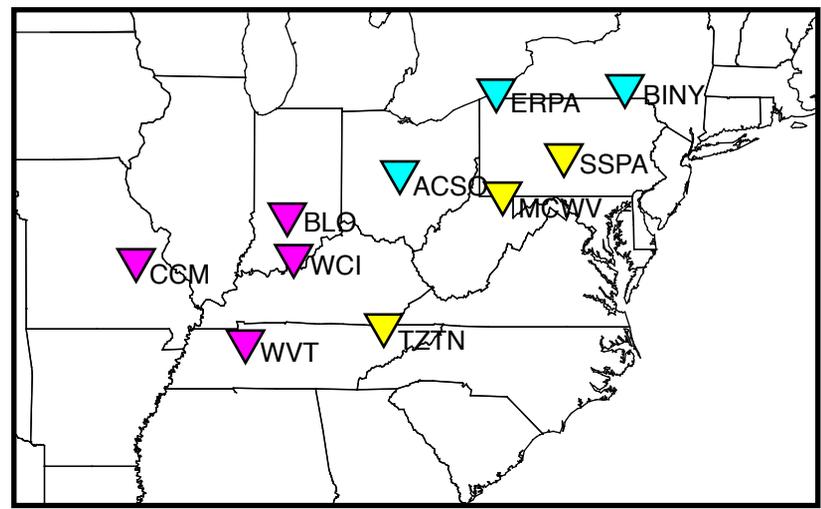
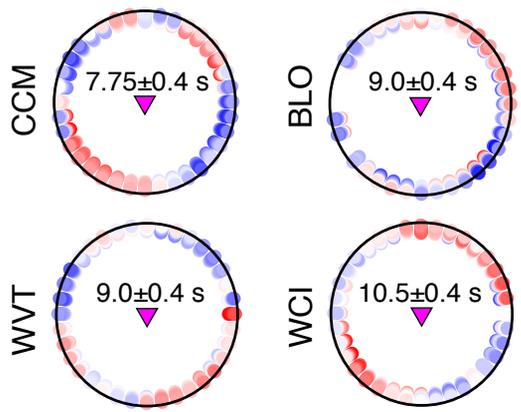
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Mountain Stations

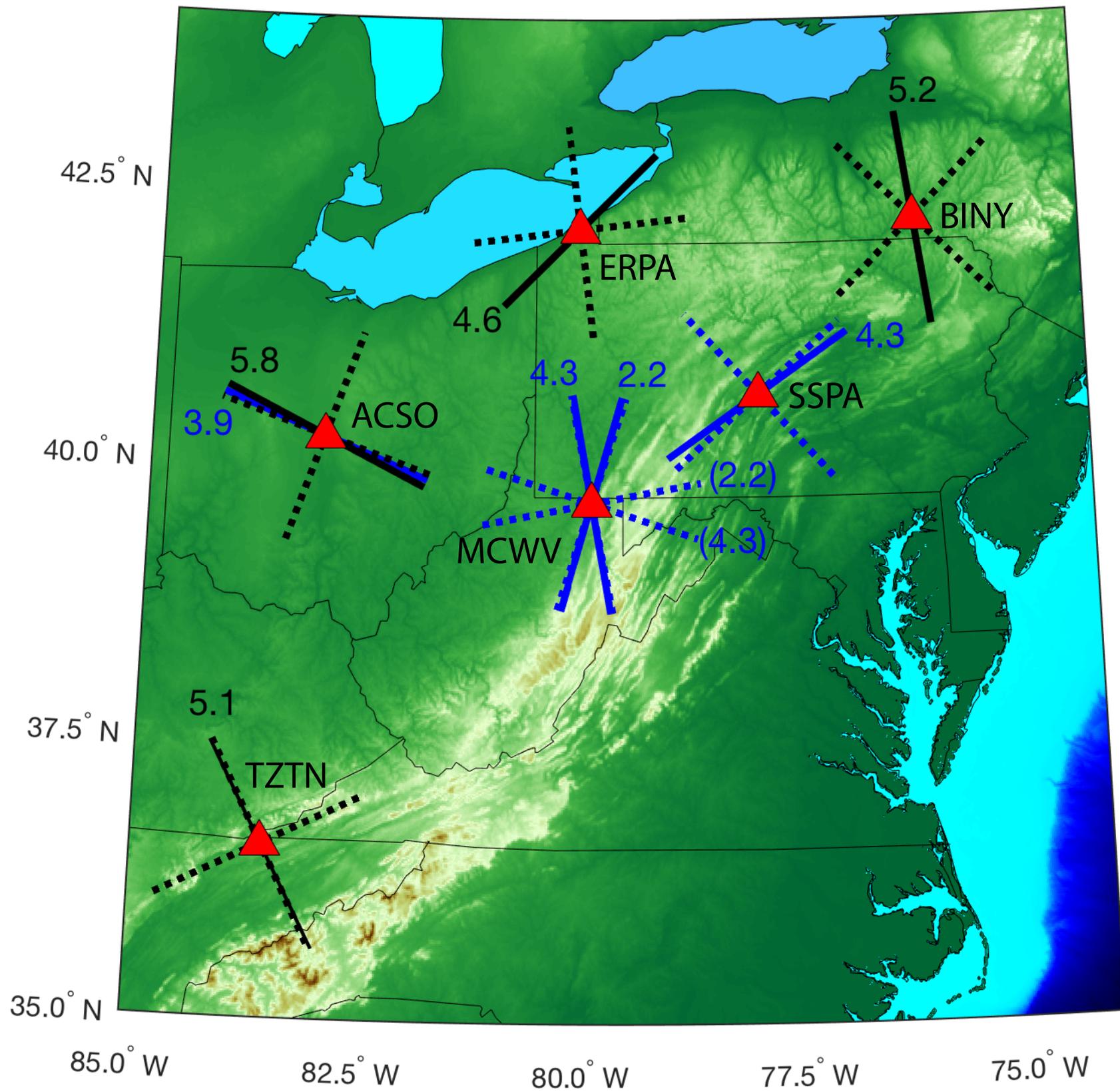
Grenville Stations



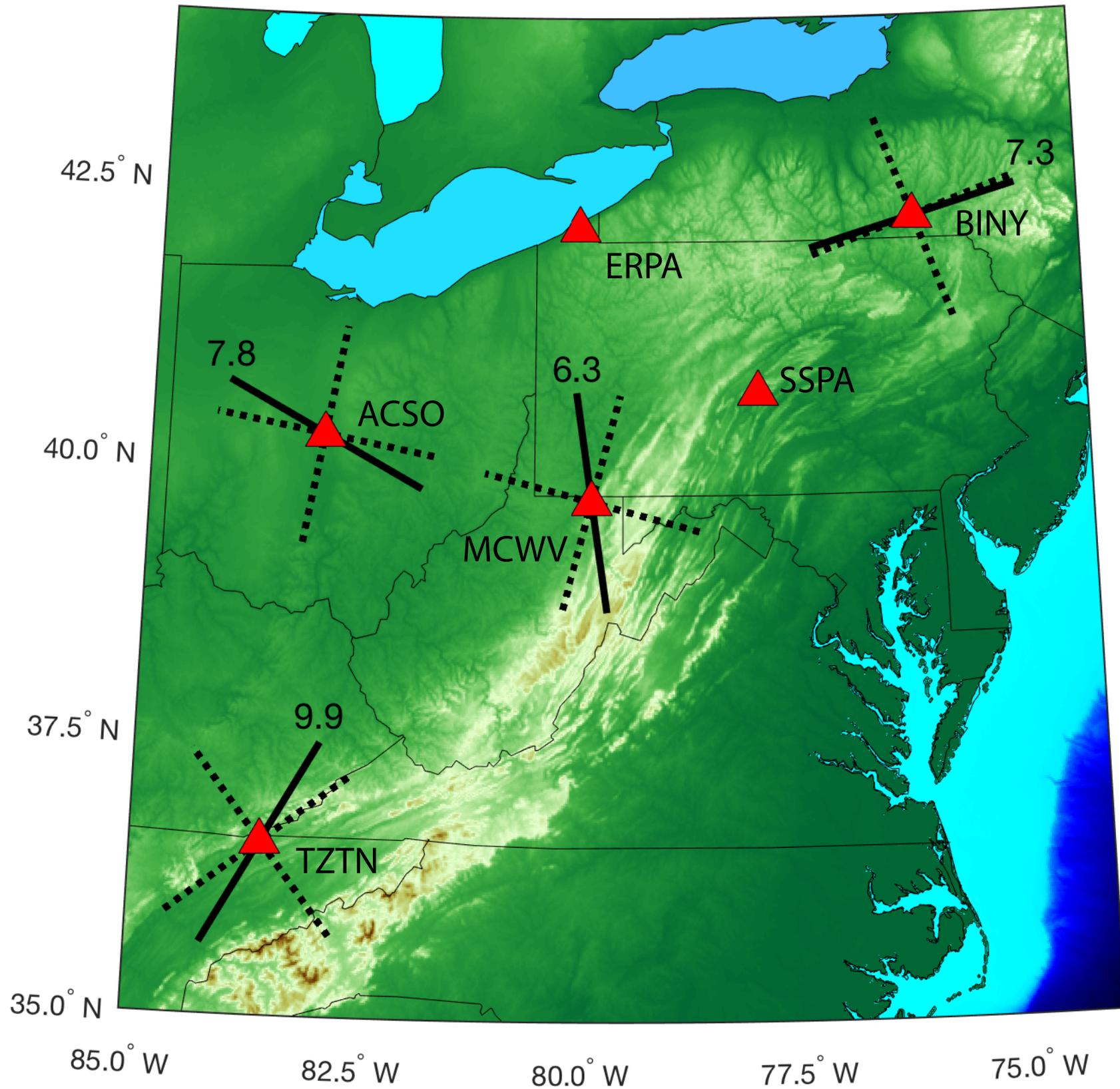
Granite-Rhyolite Stations



Crustal interfaces



Shallow intralithospheric interfaces



Co-existing with isotropic MLDs

