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Full wave sensitivity of $SK(K)S$ phases to arbitrary anisotropy in the upper and lower mantle

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SUMMARY

Core-refracted phases such as SKS and $SKKS$ are commonly used to probe seismic anisotropy in the upper and lowermost portions of the Earth's mantle. Measurements of $SK(K)S$ splitting are often interpreted in the context of ray theory, and their frequency dependent sensitivity to anisotropy remains imperfectly understood, particularly for anisotropy in the lowermost mantle. The goal of this work is to obtain constraints on the frequency dependent sensitivity of $SK(K)S$ phases to mantle anisotropy, particularly at the base of the mantle, through global wavefield simulations. We present results from a new numerical approach to modeling the effects of seismic anisotropy of arbitrary geometry on seismic wave propagation in global 3D Earth models using the spectral element solver AxiSEM3D. While previous versions of AxiSEM3D were capable of handling radially anisotropic input models, here we take advantage of the ability of the solver to handle the full fourth-order elasticity tensor, with 21 independent coefficients. We take advantage of the computational efficiency of the method to compute wavefields at the relatively short periods (5s) that are needed to simulate $SK(K)S$ phases. We benchmark the code for simple, single-layer anisotropic models by measuring the splitting (via both the splitting intensity and the traditional splitting parameters ϕ and δt) of synthetic waveforms and comparing them to well-understood analytical solutions. We then carry out a series of numerical experiments for laterally homogeneous upper mantle anisotropic

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6 23 models with different symmetry classes, and compare the splitting of synthetic waveforms to
7 24 predictions from ray theory. We next investigate the full wave sensitivity of $SK(K)S$ phases
8 25 to lowermost mantle anisotropy, using elasticity models based on crystallographic preferred
9 26 orientation of bridgmanite and post-perovskite. We find that $SK(K)S$ phases have signif-
10 27 icant sensitivity to anisotropy at the base of the mantle, and while ray theoretical approx-
11 28 imations capture the first-order aspects of the splitting behavior, full wavefield simulations
12 29 will allow for more accurate modeling of $SK(K)S$ splitting data, particularly in the pres-
13 30 ence of lateral heterogeneity. Lastly, we present a cross-verification test of AxiSEM3D against
14 31 the SPECFEM3D_GLOBE spectral element solver for global seismic waves in an anisotropic
15 32 Earth model that includes both radial and azimuthal anisotropy. A nearly perfect agreement is
16 33 achieved, with a significantly lower computational cost for AxiSEM3D. Our results highlight
17 34 the capability of AxiSEM3D to handle arbitrary anisotropy geometries and its potential for
18 35 future studies aimed at unraveling the details of anisotropy at the base of the mantle.

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21 36 **Key words:** Elasticity Tensor – Numerical Simulation – Spectral Element Method –
22 37 Anisotropy – Lower Mantle – Wave Propagation
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1 INTRODUCTION

Seismic anisotropy, the property of elastic materials to manifest directionally dependent seismic wave speeds (e.g., Anderson, 1989; Babuska & Cara, 1991), occurs in many regions of the Earth, including the crust (e.g., Barruol & Kern, 1996), the upper mantle (e.g., Silver, 1996; Savage, 1999), the transition zone (e.g., Foley & Long, 2011; Yuan & Beghein, 2013), the uppermost lower mantle (e.g., Lynner & Long, 2015; Ferreira et al., 2019), the D'' region at the base of the mantle (e.g., Nowacki et al., 2011; Creasy et al., 2017), and the inner core (e.g., Beghein & Trampert, 2003). Because mantle anisotropy reflects deformation processes, knowledge of its presence, style, and strength yields insight into past and present mantle flow (e.g., Long & Becker, 2010). The proper characterization of seismic anisotropy is therefore crucial for our understanding of the dynamics of Earth's mantle. Our ability to completely characterize anisotropy in the mantle is limited, however, in part due to limitations imposed by seismic data coverage, and in part due to theoretical or computational limitations to relate observations to Earth structure. It is common in many global seismological studies to either neglect anisotropy entirely, and consider an isotropic approximation to Earth structure, or to consider only simple anisotropic geometries, such as radial anisotropy.

Elastic anisotropy manifests itself in the seismic wavefield in many ways, including the difference in propagation velocity between vertically polarized Rayleigh waves and horizontally polarized Love waves (e.g., Anderson, 1961; Moulík & Ekström, 2014), the splitting of normal modes (e.g., Anderson & Dziewonski, 1982; Tromp, 1995; Beghein et al., 2008), the directional dependence of travel times of body waves such as P_n (e.g., Hess, 1964; Buehler & Shearer, 2017) or surface waves (e.g. Forsyth, 1975; Schaeffer et al., 2016), the scattering of energy from Love waves to Rayleigh waves via the coupling of spheroidal and toroidal modes (e.g., Park & Yu, 1993; Servali et al., 2020), the polarization of P waves (e.g., Schulte-Pelkum et al., 2001), and directionally dependent P -to- S conversions as manifested in receiver functions (e.g., Levin & Park, 1998; Wirth & Long, 2014). The most widely used technique for detecting anisotropy in the mantle, however, is shear wave splitting or birefringence (e.g., Silver, 1996; Savage, 1999; Long & Silver, 2009). The splitting of SKS and $SKKS$ phases is routinely measured to study anisotropy in both the

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upper mantle (e.g., Silver & Chan, 1991; Wolfe & Silver, 1998; Levin et al., 1999; Long & van der Hilst, 2005; Long, 2013; Roy et al., 2014) and in the lowermost mantle (e.g., Niu & Perez, 2004; Restivo & Helffrich, 2006; Long, 2009; Long & Lynner, 2015; Roy et al., 2014; Grund & Ritter, 2018; Reiss et al., 2019). Core traversing phases such as *SKS* and *SKKS* have several distinct advantages for shear wave splitting analysis. These include the known initial polarization of the shear wave, controlled by the *P* to *S* conversion at the core-mantle boundary (CMB), the lack of source-side effects, and the ability to observe clear *SK(K)S* phases that are often easily identifiable on seismograms. Shear wave splitting analysis also has several shortcomings, however; chief among these is the lack of vertical resolution of anisotropy, since it is a path-integrated measurement, and the need to obtain splitting measurements from multiple azimuths in order to fully characterize the anisotropic structure.

While a full 21 elastic parameters are needed to fully describe arbitrary anisotropy, it is common to use simpler parameterizations of anisotropy that invoke assumptions about anisotropic symmetry. For example, in global tomographic inversions that include radial anisotropy, under the assumption of hexagonal symmetry (e.g., Auer et al., 2014; Tesoniero et al., 2015), it is typical to use 5 parameters to describe the model, rather than the 2 needed for the isotropic case (e.g., Ritsema et al., 2011). Similarly, inversions of *SKS* splitting data for azimuthal anisotropy in the upper mantle typically rely on reduced parameterizations (e.g., Monteiller & Chevrot, 2011; Lin et al., 2014a; Mondal & Long, 2019). While such parameterizations may make sense in the context of practical limitations on observational data sets, they may not always be realistic for actual Earth materials. For example, olivine, the primary mineral constituent of the upper mantle and the major cause of upper mantle anisotropy, has orthorhombic symmetry, although deformed olivine aggregates may be approximated with higher symmetry classes (e.g., Karato et al., 2008). In any case, it is desirable to have computational tools that can simulate accurate wave propagation through anisotropic media of arbitrary symmetry efficiently; furthermore, azimuthal anisotropy is a well-known property of the upper mantle, so it is necessary for wavefield modeling schemes to be able to handle azimuthal anisotropy in addition to the more commonly invoked radial anisotropy.

Measurements of shear wave splitting are commonly interpreted in the framework of ray the-

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ory, either implicitly or explicitly. The most straightforward interpretation of SKS splitting measurements, for example, invokes a single layer of azimuthal anisotropy beneath a station whose properties (symmetry axis orientation, strength of anisotropy, and/or layer thickness) are related to the observed splitting parameters (typically fast splitting direction, ϕ and delay time, δt) via a simple ray theoretical approximation. In some cases, complex patterns of SKS splitting, in which apparent splitting parameters vary with backazimuth, are interpreted as reflecting multiple layers of anisotropy (e.g., Marson-Pidgeon & Savage, 2004; Eakin & Long, 2013), via analytical equations that were developed based on a ray theoretical approximation (Silver & Savage, 1994). While there has been some work on the nature of the frequency dependent sensitivity of SKS phases to upper mantle anisotropy (e.g., Favier & Chevrot, 2003; Favier et al., 2004; Chevrot, 2006; Long et al., 2008; Sieminski et al., 2008; Lin et al., 2014a; Mondal & Long, 2019), only a few observational studies have actually used finite-frequency sensitivity estimates to interpret (or invert) actual data (Monteiller & Chevrot, 2011; Lin et al., 2014b). Furthermore, the finite-frequency sensitivity of SKS and $SKKS$ phases to anisotropy in the lowermost mantle remains poorly understood. Given the increasing use of $SK(K)S$ phases in studies of deep mantle anisotropy, it is crucial to understand the nature of this sensitivity.

For both upper and lowermost mantle anisotropy studies, it is desirable to have a computationally efficient tool to simulate global seismic wave propagation for $SK(K)S$ phases in anisotropic media with arbitrary symmetry. The popular spectral-element based community software package SPEC-FEM3D_GLOBE (Komatitsch & Tromp, 2002a,b) is capable of handling arbitrary anisotropy, but its significant computational requirements make global simulations at the periods relevant for $SK(K)S$ phases (down to $\sim 5 - 10s$) impractical. In this study, we make use of the AxiSEM3D code (Leng et al., 2016, 2019), a coupled pseudo-spectral spectral element solver for 3D global wavefield propagation in realistic 3D Earth models. While previously released versions of AxiSEM3D only handled radially anisotropic input models, the actual solver is capable of handling the full fourth-order elasticity tensor C_{ijkl} with 21 independent coefficients. We have modified the formulation of the input models to handle arbitrary elasticity, and in this study we test and implement a range of anisotropic mantle models that include azimuthal anisotropy, relevant for $SK(K)S$ split-

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122 ting. The AxiSEM3D code combines the advantages of a full 3D spectral element method to model
123 complex, 3D structures with the computational efficiency of axi-symmetric methods; this allows
124 the user to model complex and realistic Earth structures with significant computational speedup
125 and without significant loss of performance (Leng et al., 2019).

126 The goal of this study is to evaluate the frequency dependent effects of anisotropy of different
127 symmetry classes (albeit in relatively simple, laterally homogeneous structures) on the splitting
128 of $SK(K)S$ phases, via global wavefield simulations for seismic periods down 5s. We generate
129 synthetic seismograms for a suite of Earth models with anisotropy in the upper and lowermost
130 mantle, and analyze the synthetic SKS and $SKKS$ waveforms by measuring shear wave split-
131 ting using both the traditional transverse component energy minimization method (Silver & Chan,
132 1988) and the splitting intensity method (Chevrot, 2000). We present benchmark results for sim-
133 ple cases, including one based on the spherically symmetric PREM model (Dziewonski & Ander-
134 son, 1981) and one that invokes vertically propagating shear waves traveling through a transverse
135 isotropic medium with a horizontal axis of symmetry (HTI). We then model global wave propa-
136 gation for more realistic upper mantle anisotropy scenarios, including approximations to olivine
137 crystallographic preferred orientation (CPO) that invoke both hexagonal and orthorhombic sym-
138 metry. Next, we investigate the effects of anisotropy in the D'' layer at the base of the mantle
139 on $SK(K)S$ splitting, via a series of global wavefield simulations for models that invoke CPO
140 of lower mantle minerals such as bridgmanite and post-perovskite. In the last part we provide a
141 verification of our implementation of full anisotropy via a benchmark solution against the full
142 3D spectral element solver SPECFEM3D_GLOBE. We then discuss the implications of our re-
143 sults from relatively simple models for future work on the interpretation of shear wave splitting
144 measurements in terms of mantle anisotropy, particularly as more complex models are considered.

145 **2 ANISOTROPIC MODELING STRATEGY WITH AXISEM3D**

146 AxiSEM3D (Leng et al., 2016, 2019) is a powerful hybrid spectral element solver for 3D global
147 wave propagation in realistic Earth structures. It aims to bridge the gap between computationally
148 expensive simulation methods for 3D Earth structure and faster simulation methods for spheri-

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6 149 cally symmetric Earth models. It is a fully convergent 3D method not unlike other discrete full
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8 150 3D methods e.g. SPECFEM3D_GLOBE. The method follows the axi-symmetric spectral element
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10 151 solver AxiSEM which assumes spherically symmetric or axi-symmetric structures. To accommo-
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12 152 date arbitrary 3D structures, AxiSEM3D is built upon a collection of coupled axi-symmetric 2D
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14 153 domains. This relies on the fact that wavefields in realistic 3D structures are relatively smooth in
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16 154 the azimuthal direction (Leng et al., 2016), and can thus be accurately represented in the azimuthal
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18 155 direction by honoring this smooth complexity rather than a complexity-blind, fine discretization
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20 156 just as in the in-plane dimension or for conventional 3D methods. AxiSEM3D couples the az-
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22 157 imuthal dimension by Fourier series, which can assume arbitrarily low or high expansion orders
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24 158 for each patch in the source-receiver plane. This hybrid formulation combines the strength of a
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26 159 full 3D discrete method to compute global wavefields with the significant speedup of computa-
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28 160 tional time of 2D methods. In other words, a 3D global wavefield in complex Earth models can be
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30 161 efficiently computed at a cost that scales almost as (multiple) 2D simulations, allowing us to run
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32 162 simulations of global wavefield propagation over a large frequency range with relatively limited
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34 163 computer resources. This computational efficiency is coupled with the capability of AxiSEM3D to
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36 164 readily allow the usage of 3D Earth models based on global tomography, including those available
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38 165 from the Incorporated Research Institutions for Seismology (IRIS) Earth Model Collaboration
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40 166 (EMC) in netCDF format. Independent modules can be invoked to handle certain geometrical
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42 167 properties, such as surface topography or topography on boundaries such as the CMB (Leng et al.,
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44 168 2019). AxiSEM3D automatically handles radial anisotropy if the velocity structure of the under-
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46 169 lying seismic velocity model is provided as horizontal and vertical compressional and shear wave
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48 170 velocity values. Detailed descriptions of the theoretical formulation of AxiSEM3D and benchmark
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50 171 tests with independent, fully 3D discrete spectral element solvers (e.g. SPECFEM3D_GLOBE) are
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52 172 given in (Leng et al., 2016, 2019).

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54 173 For this study, we take advantage of the accuracy and efficiency of the method to solve the
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56 174 elastodynamic wave equation in global media, and implement a new independent module which
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58 175 introduces general anisotropy of arbitrary symmetry. The mathematical formulation of the three-
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60 176 dimensional wave equation is based on the generalized Hooke's law, $\sigma_{ij} = C_{ijkl}\varepsilon_{kl}$, which relates

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6 177 the components of the stress tensor σ_{ij} to the strain tensor ε_{kl} through the fourth-order elasticity
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8 178 tensor C_{ijkl} . The elasticity tensor describes the elastic properties of the material, and in the most
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10 179 general case it can be fully described by 21 independent coefficients (see Malvern, 1969; Babuska
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12 180 & Cara, 1991, for details). In the new anisotropic module, the elastic properties of a specific
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14 181 anisotropic region in the domain are described by 21 independent elastic parameters, as opposed
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16 182 to the 2 Lamé constants (λ and μ) for the isotropic case or the 5 Love parameters (A, C, L, N, F,
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18 183 described further below) for the vertical transverse isotropy (VTI) case. In order to use the general
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20 184 anisotropy module in AxiSEM3D, the user must provide the solver with a netCDF file containing
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22 185 the information about the elasticity tensor and the anisotropic media's geographical location within
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24 186 the model domain.

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26 187 Introducing general anisotropy (an intrinsically 3D property of the medium) usually only ac-
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28 188 counts for a relatively small increment of the total computational cost. However, similarly to the
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30 189 case for lateral heterogeneity in (isotropic) velocity structure, the user must ensure that the Fourier
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32 190 degree of expansion is larger than the size of the heterogeneous structure, in order to ensure a
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34 191 correct representation of the seismic wavefield. Moreover, the user must take particular caution in
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36 192 taking into account the geometrical properties of the elasticity tensor. This is usually defined in a
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38 193 local coordinate system such that certain crystallographic axes (or a strain directions) align certain
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40 194 geographic coordinates. However, AxiSEM3D works with a global spherical coordinate system.
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42 195 Therefore, except at the geographical poles (where the coordinate systems coincide), a rotation
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44 196 matrix must be applied to the elasticity tensor (specified by the user in a local coordinate system)
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46 197 to translate its orientation to the global domain of AxiSEM3D.

47 48 49 50 198 **3 BENCHMARK TESTS FOR SIMPLE ANISOTROPIC MODELS**

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53 199 We begin by running two different numerical tests with the intention to benchmark the new
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55 200 anisotropic module of AxiSEM3D against well-understood wave propagation behavior for simple
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57 201 anisotropic models. In the first of these, we run a global wavefield simulation for the 1D PREM
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59 202 model (Dziewonski & Anderson, 1981), which includes radial anisotropy in the depth range be-
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203 tween $25km$ and $220km$ depth. In the second of these, we run wave propagation simulations for a

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204 simple, single layer of anisotropy with a horizontal axis of transverse isotropy (HTI), reproducing
205 a synthetic test presented in Chevrot (2000).

206 **3.1 Radially anisotropic PREM**

207 In the first test, we run global wavefield simulations for PREM under two different conditions:
208 (1) using the traditional model input format of AxiSEM3D, in which radial anisotropy is specified
209 via horizontal and vertical velocity values for both P and S waves, and (2) using the model input
210 format of the new anisotropy module, in which anisotropy is specified via the full elasticity tensor.
211 This allows us to test whether we have correctly implemented the specification of elasticity in
212 the new module, and whether we have correctly implemented the tensor rotations needed to map
213 the local coordinate system of the model input file to the global spherical coordinate system of
214 AxiSEM3D. In order to parameterize the PREM model in the new input file, we replace the elastic
215 constants in the depth range between $24km$ and $220km$ with equivalent elastic coefficients of the
216 6×6 elasticity tensor that describes the radial anisotropy contained in anisotropic PREM. At other
217 depths, we use the isotropic PREM values, which are specified via the Lamé constants λ, μ (and
218 the density ρ). The relationship between the full elasticity tensor C_{ijkl} and the Love parameters
219 that describe radial anisotropy are given by the following relationships (e.g., Babuska & Cara,
220 1991):

$$A = \rho V_{PH}^2 \quad (1)$$

$$C = \rho V_{PV}^2$$

$$L = \rho V_{SV}^2$$

$$N = \rho V_{SH}^2$$

$$F = \eta(A - 2L)$$

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$$C_{ijkl} = \begin{pmatrix} A & A - 2N & F & \cdot & \cdot & \cdot \\ A - 2N & A & F & \cdot & \cdot & \cdot \\ F & F & C & \cdot & \cdot & \cdot \\ \cdot & \cdot & \cdot & L & \cdot & \cdot \\ \cdot & \cdot & \cdot & \cdot & L & \cdot \\ \cdot & \cdot & \cdot & \cdot & \cdot & N \end{pmatrix} \quad (2)$$

221 with A, C, L, N being the parameters associated to the horizontal and vertical compressional and
 222 shear seismic wavespeed and F being the parameter associated to seismic waves propagating in
 223 intermediate directions through the coefficient η . We run the simulation for a hypothetical earth-
 224 quake in the South American subduction zone whose moment tensor is represented by the focal
 225 mechanism in Figure 1, which also shows the two sets of synthetic seismograms for different
 226 components at three representative stations (FFC, GRFO, and COCO). This test shows the general
 227 ability of our new anisotropy module to correctly specify the full elasticity tensor for models that
 228 include radial anisotropy; however, this test is not capable of capturing the effects that azimuthally
 229 dependent anisotropy have on the propagating wavefield.

3.2 Single anisotropic layer with vertical wave propagation

231 In order to test whether the new anisotropy model correctly represents the effects of azimuthal
 232 anisotropy for simple models, we carry out a second benchmarking exercise modeled on a syn-
 233 thetic test presented in Chevrot (2000). We consider wave propagation through a single layer of
 234 anisotropy of thickness $30km$, with a horizontal axis of transversely isotropic symmetry (HTI)
 235 at an azimuth of 45° North. The strength of anisotropy in the layer is 3%, and the value of the
 236 parameter η is 1.03. We use a density of $2450kg/m^3$ along with an average P wave velocity of
 237 $6.50km/s$ and an S wave velocity of $3.75km/s$ for the anisotropic layer. In order to mimic the syn-
 238 thetic test of Chevrot (2000) closely, and to avoid global waveform effects, we use a point source
 239 sitting directly beneath the receiver, just beneath the anisotropic layer to induce perfectly vertical
 240 S wave propagation. We use a monopole point source that radiates energy entirely polarized in a
 241 single direction, and consider different polarization directions of the source in order to explore the

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6 242 azimuthal dependence of the waveforms. To properly define the backazimuth, we set both source
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8 243 and receiver located along the same vertical direction, with the source slightly shifted along the
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10 244 latitudinal direction by 0.01° . This minor deviation does not substantially affect the polarization
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12 245 of the seismic energy which will remain confined in a single direction. As an alternative approach,
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14 246 we also tested a scheme that keeps source and receiver perfectly assigned and instead rotates the
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16 247 elasticity tensor by 15° for each simulation. Both approaches deliver the same results. We desig-
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18 248 nate the horizontal component parallel to the initial polarization direction of the S wave as the
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20 249 radial component, and examine how the behavior of the transverse component varies with azimuth
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22 250 (following Chevrot (2000)). We convolve the synthetic seismograms with a Gaussian source-time
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24 251 function with half-duration of $5s$, which corresponds to the resolution period of the mesh used for
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26 252 this simulation.

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28 253 Radial and transverse component seismograms as a function of backazimuth for this simula-
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30 254 tion are shown in Figure 2, which demonstrates the expected behavior of the transverse component
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32 255 waveform for azimuthal anisotropy (Vinnik et al., 1989; Chevrot, 2000). While the radial compo-
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34 256 nent does not change much with azimuth, the transverse component is azimuthally dependent, and
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36 257 its shape takes the form of the time derivative of the radial component, multiplied by a scalar fac-
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38 258 tor that depends on the angle between the initial polarization of the shear wave and the azimuth of
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40 259 the fast symmetry axis (as well as the strength of anisotropy). We observe the expected behavior
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42 260 of no energy on the transverse component in the case where the initial polarization is parallel or
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44 261 perpendicular to the fast axis orientation (Figure 2), and a maximum in transverse component en-
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46 262 ergy when the initial polarization is 45° from the symmetry axis. Following equations (7) and (A4)
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48 263 from Chevrot (2000), we measure the splitting intensity for each synthetic seismogram, which is
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50 264 defined as:

$$SI = -2 \frac{T(t)R'(t)}{\|R'(t)\|^2} \quad (3)$$

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57 265 The splitting intensity measured as a function of backazimuth forms the splitting intensity vector
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59 266 which is shown in Figure 3. Estimates of the splitting parameters (ϕ δt) can be derived from
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267 fitting a $\sin(2\theta)$ curve to the splitting intensity vector shown in Figure 3, using the relationship

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268 $SI \simeq -1/2\delta t \sin 2(\beta)$, where β corresponds to the angle between the initial polarization direction
269 and the azimuth of the fast symmetry axis. Our estimates of the splitting parameters derived from
270 Figure 3 are $\phi = 45^\circ$ and $\delta t = 0.228s$, which agree well with the values for the same test in
271 Chevrot (2000).

272 **4 UPPER MANTLE ANISOTROPY SIMULATIONS AND COMPARISON WITH RAY** 273 **THEORY**

274 **4.1 Background: Upper mantle anisotropy and shear wave splitting**

275 Seismic anisotropy in the upper mantle is generally interpreted in terms of the CPO of anisotropic
276 upper mantle minerals, primarily olivine. The relationships between strain and the resulting CPO
277 are complex and depend on many factors, potentially including stress, temperature, pressure, water
278 content, deformation history, and preexisting fabric (e.g., Karato et al., 2008; Skemer & Hansen,
279 2016). Despite these complications, however, a number of simplified relationships are often used
280 to interpret $SK(K)S$ splitting data, often with implicit assumptions made about the symmetry of
281 the anisotropic medium. For example, upper mantle anisotropy is often modeled using a hexag-
282 onal approximation to the actual elastic tensor (e.g., Browaeys & Chevrot, 2004; Becker et al.,
283 2006), even though olivine itself is orthorhombic, and even though actual aggregates of mantle
284 rocks may have even lower symmetry in practice. As discussed in the Introduction, the interpre-
285 tation of $SK(K)S$ data is usually done in the context of ray theory, with only a few exceptions.
286 While a few studies have examined the finite-frequency sensitivity of $SK(K)S$ splitting to upper
287 mantle anisotropy (e.g., Favier & Chevrot, 2003; Favier et al., 2004; Chevrot, 2006; Long et al.,
288 2008; Sieminski et al., 2008; Zhao & Chevrot, 2011; Lin et al., 2014a), it is relatively uncom-
289 mon to carry out global wavefield simulations to model the effects of upper mantle anisotropy
290 on $SK(K)S$ splitting. The purpose of the tests presented in this section is to establish the new
291 anisotropy module of AxiSEM3D as a viable tool for doing so, and to explore how anisotropy
292 models of increasing complexity in terms of their symmetry (from hexagonal to orthorhombic)
293 affect $SK(K)S$ waveforms.

294 A second goal of these tests is to understand to what extent shear wave splitting measurements

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performed on synthetic seismograms derived from global wavefield simulations depart from the predictions of ray theory. Work by Lin et al. (2014a) established that full-wavefield effects on *SK(K)S* phases can cause significant deviations from ray theoretical predictions of shear wave splitting, even for relatively simple (laterally homogenous) upper mantle anisotropy models. We aim to extend our understanding of this phenomenon here, particularly as it relates to the interpretation of *SKS* – *SKKS* splitting discrepancies. Relatively large discrepancies between the splitting of *SKS* and *SKKS* phases for the same event-station pair are typically interpreted as evidence for a contribution to the splitting of one or both phases from anisotropy in the lower mantle (e.g., Niu & Perez, 2004; Long, 2009; Lynner & Long, 2014; Grund & Ritter, 2018; Reiss et al., 2019). However, Lin et al. (2014a) showed that wavefield effects that depart from the predictions of ray theory can cause small discrepancies in splitting intensity values (up to $\sim 0.3s$) between *SKS* and *SKKS* phases for models that only include upper mantle anisotropy. The work of Lin et al. (2014a) used computations of finite-frequency sensitivity kernels based on the formulation of Zhao & Chevrot (2011). Here we build on this work using a complementary method (global wavefield simulations as opposed to explicit sensitivity kernel computations).

In order to compare the splitting of synthetic *SK(K)S* waveforms to the predictions of ray theory (and eventually to real data), we measure both the splitting intensity introduced by Chevrot (2000), as described in section 3.2, and the traditional splitting parameters (ϕ , δt) using the transverse component minimization method of Silver & Chan (1991). We use the implementation of the Silver & Chan (1991) method in **SplitRacer** (Reiss & Rumpker, 2017), a MATLAB-based graphical user interface for teleseismic shear wave splitting analysis. We compare the splitting measurements made on synthetic waveforms with predictions from ray theory, derived from the Python tool **christoffel** (Jaeken & Cottenier, 2016), which solves the Christoffel equation to predict the polarizations and velocities of the quasi-*S* phases for waves propagating over a range of directions.

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6 320 **4.2 Upper mantle case #1: Horizontal transverse isotropy (HTI)**

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8 321 We now introduce a global wavefield simulation for a case that includes azimuthal anisotropy,
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10 322 in the form of a (single, laterally homogenous) anisotropic layer in the upper mantle with HTI
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12 323 symmetry. We do this simply by modifying the radially anisotropic upper mantle layer in PREM
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14 324 from a VTI symmetry to an HTI symmetry by rotating the symmetry axis by 90° . The symmetry
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16 325 axis thus lies in the horizontal plane for this simulation. The properties of the anisotropic layer are
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18 326 shown visually in Figure 4, which shows representations of the anisotropic properties of various
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20 327 scenarios tested in this paper. The VTI elastic tensor of PREM is represented in Figure 4, top row,
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22 328 while the rotated HTI tensor used in this test is shown in Figure 4, second row. In order to avoid
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24 329 waveform complexity due to structural heterogeneity, we use a laterally homogeneous, 1D Earth
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26 330 model that corresponds to PREM except for in the depth range between $24km$ and $220km$, where
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28 331 we impose the HTI elastic tensor described above. We neglect attenuation and ellipticity in the
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30 332 simulation, even though this can be simulated with AxiSEM3D in general. We use a source and
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32 333 receiver configuration shown in the left panel of Figure 5, with the source located at the surface
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34 334 (to avoid depth phases resulting from reflections off the free surface). The receiver is located at
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36 335 the North Pole, while the sources are located at 15° azimuthal intervals at an epicentral distance of
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38 336 120° , at which both SKS and $SKKS$ phases should be visible (e.g., Long, 2009). The moment
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40 337 tensor for each source is chosen to maximize the radiation of S wave energy in the direction
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42 338 of the receiver, as shown in Figure 5. While our simulations reproduce the entire wavefield, we
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44 339 focus on the portions of the synthetic seismograms that show the SKS , $SKKS$, and S_{diff} phases;
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46 340 the ray-theoretical paths of these phases are shown in the right panel of Figure 5. The synthetic
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48 341 seismograms for the radial and transverse components plotted against the backazimuthal direction
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50 342 of the incoming wavefield are shown in Figure 6 in a time-window that includes SKS , $SKKS$,
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52 343 and S_{diff} phases. We also show synthetic waveforms as a function of backazimuth for epicentral
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54 344 distances of 100° and 110° in the Supplementary Information (Figures S1-S2).

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57 345 The splitting analysis results of this simulation are shown in Figure 7, which shows the mea-
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59 346 sured splitting intensity as a function of backazimuth from the synthetic seismograms for SKS
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347 and $SKKS$ phases, along with the predictions from ray theory. The ray theoretical predictions

Full wave sensitivity of $SK(K)S$ phases to arbitrary anisotropy in the upper and lower mantle 15

348 were obtained by solving the Christoffel equation for an elasticity tensor that was averaged over
 349 the depth range of the anisotropic layer, with the SKS and $SKKS$ propagation directions derived
 350 from **TauP** (Crotwell et al., 1999) for the PREM model. As expected, the measured splitting in-
 351 tensity values exhibit a variation with backazimuth that is close, but not identical, to that predicted
 352 by ray theory (solid orange and dashed red curves in Figure 7). The full wavefield solution yields
 353 splitting intensity values that are systematically smaller, by up to $\sim 0.25s$, than the ray theoret-
 354 ical predicted values. Similar to Lin et al. (2014b), we also find modest differences in SKS and
 355 $SKKS$ splitting intensity, of up to $\sim 0.2s$, at certain azimuths, with the difference being largest
 356 at a backazimuth that is 30° from the fast direction. We retrieve the best-fitting splitting parame-
 357 ters (ϕ , δt) for SKS and $SKKS$, respectively, by fitting a $\sin(2\theta)$ curve to the splitting intensity
 358 measurements in Figure 7, and obtain values of $\phi = 90^\circ$ and $\delta t = 0.780s$ for SKS and $\phi = 90^\circ$
 359 and $\delta t = 0.862s$ for $SKKS$. These values generally compare well to the ray theoretical values
 360 of $\delta t = 1.036s$ for SKS and $\delta t = 1.038s$ for $SKKS$, although again, ray theory overpredicts
 361 the amplitude of the splitting compared to the wavefield simulations. Finally, we compare the
 362 splitting parameters obtained from the splitting intensity curve to those obtained by the transverse
 363 component minimization method at two representative backazimuths, shown in Figure 8. As with
 364 the splitting intensity measurements, there are modest differences in measured splitting parameters
 365 between SKS and $SKKS$ phases at the same backazimuth, although in this case these differences
 366 are not significant considering the error estimates (Figure 8).

4.3 Upper mantle case #2: Orthorhombic symmetry

368 We now explore a case in which the upper mantle anisotropic layer has a lower symmetry class.
 369 Because CPO of olivine is generally understood to be the main cause of upper mantle anisotropy,
 370 and because single-crystal olivine has an orthorhombic symmetry, it is reasonable to test an or-
 371 thorhombic model. Single-crystal olivine has a strong anisotropy, up to $\sim 25\%$ for P waves and
 372 $\sim 22\%$ for S waves, according to laboratory measurements (Babuska & Cara, 1991; Isaak, 1992;
 373 Abramson et al., 1997; Mao et al., 2015). The bulk anisotropy in deformed olivine aggregates,
 374 however, is significantly lower than the single-crystal anisotropy (e.g., Ben Ismail & Mainprice,

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6 375 1998; Karato et al., 2008). We constructed a model for upper mantle anisotropy by taking an
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8 376 averaging approach that preserves the symmetry class of the orthorhombic single crystal but de-
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10 377 creases its anisotropy strength. We used the **MSAT** code (Walker & Wookey, 2012) to calculate
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12 378 a linear mixture consisting of 30% San Carlos single-crystal olivine at upper mantle conditions,
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14 379 as described by Abramson et al. (1997), and 70% isotropic PREM. The resulting elasticity tensor
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16 380 (Figure 4, third row) has a similar symmetry of a single crystal of olivine (we choose an orientation
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18 381 such that the [100] and [001] axes are in the horizontal direction, and the [010] axis is oriented ver-
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20 382 tically), but the anisotropy is substantially weaker ($\sim 5\%$ maximum). As in the HTI upper mantle
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22 383 simulation, the anisotropic layer is confined between $24km$ and $220km$ depth.

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24 384 We show the results of our wavefield simulation in Figure 9, which is similar to Figure 7 but
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26 385 for the orthorhombic symmetry case. Synthetic seismograms across a range of backazimuths at
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28 386 epicentral distances of 100° , 110° , and 120° for this simulation are shown in the Supplementary
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30 387 Information (Figures S3-S5). Because of the departure from hexagonal symmetry, the ray theoretic
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32 388 cal predictions (solid orange and dashed red lines in Figure 9) are no longer perfect $\sin(2\theta)$ curves,
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34 389 and in this case ray theory predicts some discrepancies between *SKS* and *SKKS* phases at cer-
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36 390 tain azimuths. The measured splitting intensity values from the synthetic seismograms show some
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38 391 departures from the ray theoretical predictions, as in the HTI simulation, but for the orthorhombic
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40 392 case the departures are more pronounced (compare Figure 9 with Figure 7). As with the HTI case,
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42 393 there are a few azimuths for which we find discrepancies in *SKS* – *SKKS* splitting intensities;
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44 394 somewhat surprisingly, this discrepancy is quite large (difference in splitting intensity of $\sim 0.4s$)
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46 395 at backazimuths that lie 45° away from the fast splitting direction. An examination of the trans-
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48 396 verse component minimization measurements at selected backazimuths (Figure 10) shows that the
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50 397 *SKKS* splitting measurement at a backazimuth of 45° is poorly constrained due to a stronger in-
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52 398 terference of a not well identifiable phase on the radial component that leads to the contamination
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54 399 of the splitting measurements.
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5 LOWERMOST MANTLE ANISOTROPY SIMULATIONS AND COMPARISON WITH RAY THEORY

5.1 Background: Lowermost mantle anisotropy and shear wave splitting

We now turn our attention to seismic anisotropy in the lowermost mantle, which has been the subject of a great deal of recent observational work (e.g., Wookey et al., 2005a; Long, 2009; Nowacki et al., 2010; Cottaar & Romanowicz, 2013; Lynner & Long, 2014; Ford et al., 2015; Creasy et al., 2017; Grund & Ritter, 2018; Reiss et al., 2019), but which is substantially more challenging to observe than upper mantle anisotropy. As summarized in Nowacki et al. (2011) and Creasy et al. (2019), there are a variety of possible mechanisms that may contribute to anisotropy in the D'' layer, including CPO of bridgmanite or post-perovskite (depending on which mineral dominates, which may in turn depend on the temperature in a given region of D'' (Houser, 2007)), CPO of ferropericlase, or shape-preferred orientation (SPO) of elastically distinct material such as partial melt. We do not test all of these possible mechanisms in this paper; instead, we test models of aligned post-perovskite and aligned bridgmanite, one of which is expected to be the most volumetrically important phase in any given region of D'' . Testing of other models for D'' anisotropy will be left to future work.

There are a variety of strategies for measuring D'' anisotropy using body waves, only some of which rely on $SK(K)S$ phases. One common approach is to measure the differential splitting of $S - ScS$ phases (introduced by Wookey et al. (2005a); see also Wookey & Kendall (2008); Ford et al. (2015); Rao et al. (2017)) or to examine the waveform behavior of phases such as S_{diff} (e.g., Cottaar & Romanowicz, 2013). The identification of significant discrepancies between the splitting of SKS and $SKKS$ phases for the same event-station pair, first documented by James & Assumpção (1996), have been attributed to anisotropy at the base of the mantle beginning with Niu & Perez (2004) and Restivo & Helffrich (2006). The measurement of $SKS - SKKS$ splitting discrepancies to study D'' anisotropy is becoming more common (e.g., Long, 2009; He & Long, 2011; Roy et al., 2014; Lynner & Long, 2014, 2015; Grund & Ritter, 2018), although there is some debate about the extent to which $SKS - SKKS$ splitting discrepancies require a contribu-

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6 427 tion to the lowermost mantle. Some studies (Monteiller & Chevrot, 2011; Lin et al., 2014a) have
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8 428 argued that $SKS - SKKS$ splitting discrepancies can be explained solely in terms of upper man-
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10 429 tle anisotropy. Recently, some studies of $SKS - SKKS$ splitting discrepancies have come to rely
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12 430 on the measurement of splitting intensity, rather than the traditional splitting parameter estimation
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14 431 methods, on the grounds that the splitting intensity is a more robust and stable measurement (Deng
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16 432 et al., 2017; Reiss et al., 2019). Despite ample observational evidence that $SK(K)S$ phases can in-
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18 433 deed be affected by anisotropy in the lowermost mantle, there is still no consensus in the literature
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20 434 about how strong the finite-frequency sensitivity of these phases is to anisotropy in D'' (Sieminski
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22 435 et al., 2008; Zhao & Chevrot, 2011). Furthermore, there are disagreements in the literature about
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24 436 to what extent documented $SKS - SKKS$ splitting discrepancies reflect a contribution to split-
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26 437 ting from lowermost mantle anisotropy to one or both phases (e.g., Niu & Perez, 2004; Restivo
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28 438 & Helffrich, 2006; Long, 2009), or whether such discrepancies can mainly be attributed to upper
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30 439 mantle anisotropy (e.g., Monteiller & Chevrot, 2011; Lin et al., 2014a). There is, therefore, a need
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32 440 for global wavefield simulations that include lower mantle anisotropy.

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34 441 Another argument for the study of full-wave effects of D'' anisotropy on $SK(K)S$ phases
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36 442 comes from recent work that has sought to carry out forward modeling of D'' anisotropy, relying
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38 443 on results from mineral physics (e.g., Walker et al., 2011; Cottaar et al., 2014; Ford et al., 2015;
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40 444 Ford & Long, 2015; Creasy et al., 2017, 2019; Walker et al., 2018; Tommasi et al., 2018). These
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42 445 types of modeling studies tend to rely exclusively on ray theory, with very few exceptions (e.g.,
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44 446 Nowacki & Wookey (2016), who looked at full-wave effects for complex anisotropic models but
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46 447 who only considered ScS phases). Given the popularity of these types of modeling approaches, it
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48 448 is crucial to understand how well the ray theoretical approximation captures the true behavior of
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50 449 $SK(K)S$ waves.

53 450 **5.2 Lowermost mantle case #1: Bridgmanite crystallographic preferred orientation**

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56 451 We first consider a case that invokes a laterally homogeneous, $250km$ thick D'' layer at the base
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58 452 of the mantle, with elasticity that is designed to capture the first-order characteristics of aligned
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60 453 bridgmanite. Because the dominant slip systems at lowermost mantle conditions remain uncertain,

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we follow the strategy of recent modeling studies (e.g., Ford et al., 2015; Creasy et al., 2017) and consider tensors based on single-crystal elasticity. We chose an orientation such that the [100] and [010] axes of bridgmanite are oriented in the horizontal direction, while the [001] axis is vertical. We create an elastic tensor via a linear mixture of 30% single crystal anisotropy of post-perovskite, using the elastic constants from Wookey et al. (2005b), with 70% of isotropic PREM at D'' depths. This mixing ratio yields an elastic tensor with an orthorhombic symmetry and with a maximum anisotropy of 6%, as shown in Figure 4, fourth row. As with the upper mantle cases, we propagate the seismic wavefield through this model using the source and receiver geometries shown in Figure 5, and measure the splitting intensity, as well as the traditional shear wave splitting parameters, as a function of backazimuth. Also as with the upper mantle cases, we compare the synthetic splitting measurements with predictions from ray theory. In contrast to the upper mantle case, in the lowermost mantle the SKS and $SKKS$ phases under study have significantly different ray propagation paths. Using **TauP** (Crotwell et al., 1999), we calculated nominal propagation angles (from the horizontal) through the D'' region of 67° for SKS and 42° for $SKKS$.

Results from our bridgmanite simulation are shown in Figures 11 and 12. Synthetic seismograms across a range of backazimuths at epicentral distances of 100° , 110° , and 120° for this simulation are shown in the Supplementary Information (Figures S6-S8). Because the raypaths of SKS and $SKKS$ phases depart so significantly from the vertical, we no longer expect a simple $\sin 2(\theta)$ variability in the splitting intensity curve, and this is borne out by the measurements on synthetic seismograms. For this reason, we do not fit a $\sin 2(\theta)$ curve to our measurements, as we did for the upper mantle cases. We do predict the ray theoretical splitting intensity curves, as shown on Figure 11; these are based on the predicted apparent splitting parameters at each azimuth, visualized in Figure 4, fourth row. As with the upper mantle cases, particularly the HTI simulation, we find that the full-wavefield simulations predict splitting intensity behavior that is generally similar to the ray theoretical predictions, but not identical to it. We see deviations of up to $\sim 0.2s$ in splitting intensity values between the ray theoretical predictions and the synthetic seismogram measurements at certain azimuths, again similar to what is observed for the upper mantle cases. Importantly, the bridgmanite model predicts spitting intensities of up to $\sim 0.8s$ for SKS

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6 482 and *SKKS* phases, nearly as large as the maximum values predicted from ray theory, indicating
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8 483 that these phases have significant sensitivity to anisotropy at the base of the mantle. Notably, both
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10 484 the synthetic seismogram measurements and the ray theoretical predictions indicate that at certain
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12 485 azimuths, significant discrepancies between *SKS* and *SKKS* splitting intensities are expected,
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14 486 even though the underlying model is laterally homogeneous.

17 487 **5.3 Lowermost mantle case #2: Post-perovskite crystallographic preferred orientation**

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20 488 Finally, we test a model that invokes aligned post-perovskite as the cause for D'' anisotropy. The
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22 489 phase transition from bridgmanite to post-perovskite, which was discovered experimentally by
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24 490 Murakami et al. (2004), is thought to be the cause of the D'' discontinuity, and is expected to
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26 491 dominate in relatively cold regions of the lowermost mantle (e.g., Hernlund et al., 2005), and
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28 492 perhaps throughout (e.g., Koelemeijer et al., 2018). Post-perovskite is generally favored as the
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30 493 most likely mechanism for D'' anisotropy by many authors (e.g., Nowacki et al., 2010; Walker
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32 494 et al., 2011; Cottaar et al., 2014; Ford et al., 2015). Similar to our test for bridgmanite, we create
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34 495 an elastic tensor that is a linear mixture of 30% single crystal anisotropy, using elastic constants
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36 496 for post-perovskite from Wookey et al. (2005b), and 70% isotropic PREM. In this scenario, we
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38 497 assume that the [100] and [010] axes are oriented in the horizontal directions, while the [001] axis
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40 498 is vertical. Predicted ray theoretical splitting patterns for this tensor are shown in Figure 4, bottom
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42 499 row.

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44 500 Results from our post-perovskite wavefield simulation, along with ray theoretical predictions
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46 501 for this scenario at the relevant azimuths, are shown in Figure 13, and Figure 14 shows repre-
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48 502 sentative transverse component minimization splitting measurements at two azimuths. Synthetic
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50 503 seismograms across a range of backazimuths at epicentral distances of 100° , 110° , and 120° for
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52 504 this simulation are shown in the Supplementary Information (Figures S9-S11). Interestingly, for
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54 505 this orientation of the post-perovskite elasticity tensor, *SKS* phases are split only very weakly,
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56 506 with maximum splitting intensities of $\sim 0.2s$. This is consistent with the ray theoretical prediction
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58 507 for *SKS* splitting behavior (Figure 13; see also Figure 4, bottom row). The behavior of *SKKS*
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60 508 phases provides a striking contrast; for the propagation angles relevant for *SKKS*, this elastic

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tensor scenario predicts strong splitting, with maximum splitting intensities of $\sim 2.5s$. The azimuthal variation in the $SKKS$ splitting intensities measured from the synthetic seismograms is to first order captured by the ray theoretical prediction (Figure 13), although the synthetic splitting intensities are generally lower (sometimes by as much as $\sim 1.0s$) than the values predicted by ray theory. For the traditional splitting measurements (Figure 14), the splitting of SKS phases is so slight that there is not enough transverse component energy to obtain a stable measurement; as expected, these measurements do not yield meaningful splitting parameter estimates. For the $SKKS$ phases, in contrast, the transverse component minimization method yields robust measurements with strong splitting, consistent with the splitting intensity measurements (Figure 13) and as predicted by ray theory.

6 CROSS-VERIFICATION AGAINST AN INDEPENDENT 3D SPECTRAL ELEMENT SOLVER

Here we present a verification of our implementation of full anisotropy in AxiSEM3D by a benchmark solution against SPECFEM3D_GLOBE (Komatitsch & Tromp, 2002a,b). It is important to point out that the following benchmark test should be regarded as a cross-verification rather than a validation per se, as SPECFEM3D_GLOBE has not yet been benchmarked for global wavefield propagation in arbitrary anisotropic Earth models. We use the global-scale 3D anisotropic model of Montagner (2002), based on surface wave tomography, which includes both radial and azimuthal anisotropy. This model spans between the Moho (24km) and the 670km discontinuity and has a 5° horizontal resolution. Its elasticity tensor is determined by 13 independent parameters, numerically implemented as a fully anisotropic tensor with 21 independent parameters in both SPECFEM3D_GLOBE and AxiSEM3D. The benchmark problem, shown in Figure 15, is based upon simulating an earthquake source located in Virginia at a depth of $12km$, corresponding to the 2011 $Mw = 5.8$ earthquake in Mineral, VA. The source time function is an error function with half-duration of 10s. The synthetic seismograms are computed at 129 stations of the Global Seismographic Network (GSN) distributed across the Earth's surface (Figure 15). Attenuation is

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6 535 turned off and the record length of the seismogram is 3600s. A nearly perfect agreement has been
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8 536 achieved between SPECFEM3D_GLOBE and AxiSEM3D in this simulation.

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10 537 In Figure 16, we show the vertical components of a set of synthetic seismograms for sta-
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12 538 tions (blue triangles in Figure 15) in the $\sim 90^\circ - 130^\circ$ range of epicentral distance. In this
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14 539 set of synthetic traces, which show a range of body wave arrivals as well as surface waves, the
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16 540 SPECFEM3D_GLOBE and AxiSEM3D traces are virtually indistinguishable. For reference, we
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18 541 also show corresponding synthetic seismograms computed for the radially anisotropic PREM,
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20 542 demonstrating the effect of the anisotropic structure in the Montagner (2002) model on the wave-
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22 543 forms. Because the Montagner (2002) model only includes structure in the upper mantle, the major
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24 544 differences with PREM manifest in the upper mantle phases (SS , SSS) and the surface waves.

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26 545 In Figure 17, we show the same set of synthetics, but we zoom in on the time window shown in
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28 546 Figure 16, to highlight the SKS and $SKKS$ arrivals. We show both radial (top panel) and trans-
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30 547 verse (bottom panel) component records; the latter shows the effect of azimuthal anisotropy in
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32 548 the Montagner (2002) model. The same version of this figure for the vertical component is shown
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34 549 in the Supplementary Information (Figure S12). For the transverse component traces in Figure
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36 550 17 (bottom panel), the PREM model predicts no energy associated with the SKS or $SKKS$
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38 551 time window; because PREM does not include azimuthal anisotropy, no splitting of $SK(K)S$
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40 552 phases is predicted. The SPECFEM3D_GLOBE and AxiSEM3D synthetics, in contrast, show
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42 553 significant transverse component $SK(K)S$ energy at several of the selected stations. Again, the
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44 554 SPECFEM3D_GLOBE and AxiSEM3D traces are virtually indistinguishable in the time windows
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46 555 associated with $SK(K)S$ arrival. We also report in the Supplementary Information the full record
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48 556 section for the radial, transverse and vertical component of the all the stations localized in the
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50 557 $\sim 90^\circ - 130^\circ$ range of epicentral distance (Figures S13-S15). We do note some extremely small
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52 558 differences on the transverse components between the SPECFEM3D_GLOBE and AxiSEM3D
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54 559 traces in PS and PPS phases at some stations (e.g. II.MSVF). These differences are not visible
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56 560 on the corresponding radial components. These small differences are due to the fact that in general,
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58 561 phases with small amplitudes are more vulnerable to numerical errors (such as floating point errors
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and discretisation errors). Because the two different methods have entirely different discretisation schemes, some small differences for low-amplitude phases are expected.

Compared to a fully discretised 3D method such as SPECSEM3D_GLOBE, AxiSEM3D has only one more parameter: the Fourier expansion order of the solution n_u (Leng et al., 2016, 2019). This parameter controls both the accuracy and the computational cost of AxiSEM3D. For a spherically symmetric Earth model such as PREM, we have $n_u = 2$ everywhere in the 2D computational domain of AxiSEM3D; in this case, AxiSEM3D degenerates to a pure axisymmetric spectral element method (Nissen-Meyer et al., 2014). For a 3D model, the AxiSEM3D solution converges to the real 3D solution as n_u increases, and its value for solution convergence is always much smaller (usually by orders of magnitude) than a fully-discretised one for global wave propagation in a realistic 3D Earth models (Leng et al., 2016). It is most efficient to vary n_u with depth and epicentral distance to maximise the performance of AxiSEM3D. Convergence tests are common practise in AxiSEM3D and are thoroughly explained in Leng et al. (2016, 2019). We refer the reader to (e.g., Leng et al., 2016, Fig.11) and to figure S16 in the Supplementary Information for a visual illustration of the convergence behaviour. In order to achieve the agreement demonstrated in Figures 16 with SPECSEM3D_GLOBE, we have used $n_u = 400$ in the uppermost 200km and $n_u = 100$ elsewhere. We enlarge n_u near the surface to have better accuracy for surface wave propagation. Such n_u field leads to highly-accurate waveforms for all the phases at all the stations in Figure 15. In terms of computer performance, with this n_u field, we have obtained a speedup of ~ 6 compared to SPECSEM3D_GLOBE. This speedup can be further increased in case neither surface waves nor multiple bouncing body waves (SSS , $SSSS$) are of interest. The technique of wavefield scanning (Leng et al., 2019) can be used to fully optimise n_u for better performance.

A key aspect of AxiSEM3D is its ability to correctly account for off great-circle scattering at minimal computational cost, via the Fourier expansion approach. This capability has been extensively discussed in previous papers (Leng et al., 2016, 2019), which include previous benchmark tests of the AxiSEM3D code against SPECSEM3D_GLOBE. Because the Montagner (2002) model involves strong upper mantle heterogeneity, and thus causes strong scattering from out-of-

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6 599 plane structure, the benchmark test presented here provides a clear demonstration of the ability of
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8 590 AxiSEM3D to capture the complex properties of scattering by anisotropic heterogeneities in 3D.
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10 11 591 **7 DISCUSSION**

12 13 14 592 **7.1 A tool for efficient global wavefield modeling in arbitrary anisotropic media**

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16 593 The series of benchmark tests and global wavefield simulations presented in this paper establishes
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18 594 the AxiSEM3D code as an efficient tool for modeling of the global wavefield in Earth models that
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20 595 include anisotropy with arbitrary symmetry. In contrast to other 3D wavefield simulation methods,
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22 596 AxiSEM3D simulations can be run at periods that are relevant for studies of $SK(K)S$ splitting
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24 597 (down to $\sim 5 - 10s$ period) at relatively modest computational cost due to the sparse sampling
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26 598 of smoothness in the azimuthal wavefield. For example, the computational requirements for the
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28 599 global simulations presented in this paper involved about 3 hours and 30 minutes (wall-clock
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31 600 time) on 40 cores to produce 2000s long synthetic seismograms that can resolve seismic periods
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33 601 of 5s.
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35 602 The benchmarking tests presented in section 3 serve to validate the implementation of arbitrary
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37 603 trary anisotropy in AxiSEM3D. In the first case, we ran a global simulation for the PREM model,
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39 604 including radial anisotropy, in two ways: one in which the model was specified in terms of ver-
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41 605 tical and horizontal wave velocities in the anisotropic layer, and one in which the same model
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43 606 was specified in terms of the equivalent full elastic tensor C_{ijkl} . The two solutions were found to
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45 607 be numerically identical, as expected. In the second test, we replicated a synthetic test presented
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47 608 in Chevrot (2000) that included azimuthal anisotropy. This test successfully demonstrated the az-
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49 609 imuthal dependence of transverse component waveforms, with the expected waveform shape, and
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51 610 reproduced the expected splitting parameters (as measured via the splitting intensity) found by
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53 611 Chevrot (2000) for this model case. The simulations for upper and lowermost mantle anisotropy
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55 612 presented in sections 4 and 5 demonstrate the feasibility of interrogating wavefield behavior due to
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57 613 anisotropy effects in the context of global models using AxiSEM3D, and the cross-validation tests
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59 614 with SPECFEM3D_GLOBE presented in section 6 yield a remarkable agreement between the two
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615 solutions, with a significant computational advantage obtained by AxiSEM3D.

7.2 Full wave sensitivity of $SK(K)S$ phases to upper and lowermost mantle anisotropy

The upper mantle anisotropy models presented here build on previous work on the finite-frequency sensitivity of $SK(K)S$ phases to azimuthal anisotropy in the upper mantle by a number of authors (e.g., Favier & Chevrot, 2003; Favier et al., 2004; Lin et al., 2014a; Mondal & Long, 2019) using a different set of tools (global wavefield simulations rather than sensitivity kernel computations). Similar to previous workers, we have found here that for laterally homogeneous models, ray theoretical approximations to the effects of upper mantle anisotropy on $SK(K)S$ splitting are generally accurate to first order. However, similar to Lin et al. (2014a), we also found that at certain azimuths and epicentral distance ranges, full waveform effects can produce modest deviations from the predictions of ray theory. Unlike Lin et al. (2014a), we examined models in our study with orthorhombic symmetry (in addition to transversely isotropic models) and found that the deviations from ray theoretical predictions are more pronounced at certain azimuths for this symmetry class.

There are several potential reasons for the (generally relatively minor) deviations between the synthetic splitting results and the ray theoretical predictions for our upper mantle models. The most important one is that AxiSEM3D considers full wave sensitivity to Earth structure and accounts for finite frequency effects, in contrast to ray theoretical methods. Other effects may include the fact that the ray theoretical calculations are based on a geometrical approximation of the path of the seismic ray traveling through an average, depth-weighted elasticity tensor; in AxiSEM3D, the elasticity is discretized in 33 different layers and varies with depth. Finally, like Lin et al. (2014a,b) we also noticed some complexities in the apparent splitting parameters due to interference with other seismic phases that could perturb the splitting measurement. This phenomenon is particularly recognizable in the HTI case, where a second wiggle is very close to the $SKKS$ phase, especially in the $30\text{--}60^\circ$ ($120\text{--}150^\circ$) backazimuthal ranges (Figure 6) and it can also be observed in the upper mantle orthorhombic case at 45° (Figure 10). These other seismic phases coming in between SKS and $SKKS$ correspond to $PPPP$ phases: it can be observed from the seismograms directly that this phase is not split (no energy on the transverse component), consistent with it being a P phase.

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5 643 The amplitude of this phase is strongly attenuated in real data seismograms, but since attenuation
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7 644 is turned off in our simulations they show up with high amplitude in our synthetics.
8

9 645 In general, while ray theory is generally an adequate approximation for $SK(K)S$ splitting due
10
11 646 to upper mantle anisotropy, it is more likely to be inadequate when the anisotropic models under
12
13 647 study are complicated (either in terms of their symmetry, as in this study, or when there are lateral
14
15 648 variations in anisotropic structure, which was not considered here but which will be considered in
16
17 649 future works). Similar to Lin et al. (2014a), we found that modest differences in splitting intensity
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19 650 for SKS vs. $SKKS$ phases (up to $\sim 0.2s$, with values up to $\sim 0.4s$ for specific symmetries
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21 651 and at certain azimuths) for the same event-station pair can be caused by waveform effects for
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23 652 models that include upper mantle anisotropy only. We note, however, that the $SKS - SKKS$
24
25 653 discrepancies due to upper mantle anisotropy documented in this study are modest compared to
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27 654 the larger discrepancies observed for the lowermost mantle cases (Section 5). For example, for the
28
29 655 bridgmanite case, discrepancies larger than $0.5s$ and up to $\sim 1s$ are observed in the $30^\circ - 75^\circ$
30
31 656 ($105^\circ - 135^\circ$) backazimuthal range (Figure 11), while discrepancies as high as $\sim 2.5s$ are ob-
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33 657 tained for post-perovskite models (Figure 13). Our findings reinforce the need for $SKS - SKKS$
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35 658 discrepancy studies of mantle anisotropy to exercise caution when interpreting weakly discrepant
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37 659 splitting, and to only attribute strong differences in splitting intensity values to anisotropy in the
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39 660 lowermost mantle (e.g., Deng et al., 2017; Reiss et al., 2019).
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42 661 For the lower mantle, our results have shown that lowermost mantle anisotropy does have a
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44 662 significant effect on the splitting of SKS and $SKKS$ phases, in agreement with the sensitivity
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46 663 kernel computations of Zhao & Chevrot (2011) but in contrast to the suggestion of Sieminski
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48 664 et al. (2008) that $SK(K)S$ phases may have limited sensitivity to anisotropy in the lower mantle.
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50 665 A comparison between our shear wave splitting measurements on synthetic waveforms and the
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52 666 predictions of ray theory (for models that represent bridgmanite and post-perovskite elasticity)
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54 667 reveals that, as for the upper mantle models, ray theory generally correctly captures the first-order
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56 668 aspects of the splitting behavior. However, ray theory typically slightly overpredicts the magnitude
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58 669 of shear wave splitting when compared to the behavior of the full waveform. From these results,
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60 670 we conclude that forward modeling approaches for lowermost mantle anisotropy models that rely

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on ray theory (e.g., Walker et al., 2011; Cottaar et al., 2014; Ford et al., 2015; Ford & Long, 2015; Creasy et al., 2017, 2019; Tommasi et al., 2018) are generally valid, but are less accurate than full-wavefield simulations. This implies that moving to approaches that consider the full waveform behavior is desirable in future modeling work aimed at D'' anisotropy problems, as also suggested by Nowacki & Wookey (2016).

7.3 Limitations and future work

While this study has established the feasibility of full-wave solutions using AxiSEM3D for Earth models that include anisotropy of arbitrary symmetry, it is important to highlight some caveats and limitations. One area in which particular concern must be taken in future studies of mantle anisotropy has to do with the design of the mesh and the size of anisotropic regions. In spectral element methods, the domain is subdivided into a number of non-overlapping quadrangular shaped elements, and on each element the nodes are chosen to be the Gauss-Lobatto-Legendre (GLL) points where the wave equation is evaluated (Peter et al., 2011). A good mesh design must consider this characteristic and allow for seismic discontinuities to lie on the GLL points. This will make sure that the points at the discontinuity have the elastic properties of both materials, which in turn will allow for the correct partitioning of seismic energy at the interface and avoid contamination of the numerical solution by artifacts due to staircase sampling effects. A proper mesh design is an essential requirement for all spectral element methods and not just AxiSEM3D. It is indeed a more complex task to achieve for 3D discrete methods as 3D meshes that honor these conditions have to be generated, while for AxiSEM3D this condition applies to 2D meshes. With this in mind, in future studies that interrogate anisotropic Earth models, it will be important to ensure that anisotropic discontinuities are also honored by the mesh design. This is relatively easy to accomplish in the case of vertical stratification of anisotropy (that is, horizontal interfaces), but it turns out to be more complex for cases of lateral variation in anisotropic properties, which are likely relevant for many regions of the Earth.

Consideration of laterally heterogeneous anisotropic structures has been intentionally neglected in this study, whose main purposes are to establish AxiSEM3D as a tool for modeling seismic

698 waveforms in arbitrary anisotropic global Earth models and to interrogate the behavior of relatively
699 simple anisotropy scenarios. However, more realistic models that include lateral heterogeneity are
700 needed to more fully explore the origin of $SKS - SKKS$ splitting discrepancies and the pos-
701 sible contributions of lowermost mantle anisotropy to $SK(K)S$ phases in the real Earth. This is
702 particularly true given that in the lowermost mantle, the length scale of anisotropic heterogeneity
703 may be smaller than the Fresnel zone of $SK(K)S$ phases; in practice, therefore, it is not clear
704 how large the contribution to $SK(K)S$ splitting may be for realistic models of lowermost man-
705 tle anisotropy. Now that AxiSEM3D has been established as a computationally efficient tool for
706 modeling wave propagation in an anisotropic Earth, work to understand the full-wave sensitivity
707 of $SK(K)S$ phases to anisotropy in more realistic, heterogeneous Earth models is ongoing. In the
708 case of 1D anisotropic models, such as those considered in this study, the Fourier spectral order
709 required for a correct approximate solution to the 3D wave equation is dependent on the complex-
710 ities of the underlying seismic velocity model. For a fully 3D anisotropic model, it is crucial to
711 ensure that the Fourier spectral order of expansion is large enough to correctly account for the ge-
712 ometrical properties of the anisotropic model. As the cost of 3D anisotropic wavefield simulations
713 increase with growing Fourier order, careful attention must be paid to the geographic dimensions
714 of anisotropic domains in 3D models in order to preserve the computational advantages of the
715 method with respect to 3D discrete spectral element methods.

716 **7.4 Implications for the interpretation of $SKS - SKKS$ splitting discrepancies**

717 Observations of discrepant $SKS - SKKS$ splitting for pairs of phases from the same event-
718 station pair have been puzzling to shear wave splitting analysts since they were first documented
719 by James & Assumpção (1996). A global study by Niu & Perez (2004) found that 95% of $SKS -$
720 $SKKS$ observations globally were non-discrepant; that is, the SKS and $SKKS$ phases showed
721 similar splitting behavior. In the remaining 5% of cases, however, the pairs exhibited discrepant
722 splitting. Niu & Perez (2004) argued that because SKS and $SKKS$ phases have nearly identical
723 ray paths in the upper mantle but diverge significantly in the lower mantle, discrepancies in $SKS -$
724 $SKKS$ splitting should be attributed to anisotropy in the lower mantle. Restivo & Helffrich (2006)

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725 suggested that anisotropy in the D'' layer is the most likely explanation for such discrepancies, and
726 argued that topography on structures at or near the CMB may generate polarization anomalies,
727 and thus splitting anomalies. A key aspect of the argument made by Restivo & Helffrich (2006)
728 is that lateral gradients in structure at the base of the mantle are responsible for $SKS - SKKS$
729 splitting discrepancies. This notion was invoked by subsequent studies of $SKS - SKKS$ splitting
730 discrepancies and lowermost mantle structure (e.g., Wang & Wen, 2007; Long, 2009), who argued
731 that $SKS - SKKS$ splitting discrepancies should be interpreted in terms of a lateral gradient in
732 seismic anisotropy at the base of the mantle (that is, with the SKS and $SKKS$ phases sampling
733 a different geometry and/or strength of lowermost mantle anisotropy). More recent papers that
734 have been informed by forward modeling studies (e.g., Ford et al., 2015; Creasy et al., 2019) have
735 explicitly acknowledged the possibility that $SKS - SKKS$ splitting discrepancies can arise from
736 homogenous anisotropy at the base of the mantle (e.g., Long & Lynner, 2015; Deng et al., 2017;
737 Wolf et al., 2019; Reiss et al., 2019). In this case, the anisotropy must be present in a geometry
738 that splits SKS and $SKKS$ phases differently.

739 While many authors have interpreted $SKS - SKKS$ splitting discrepancies as reflecting a
740 contribution from anisotropy in the lower(most) mantle, other workers have questioned the extent
741 to which anisotropy in the upper mantle may give rise to such discrepancies (or to discrepancies in
742 splitting between SKS phases measured at the same station at similar backazimuths and incidence
743 angles). For example, Monteiller & Chevrot (2010) documented differences in transverse compo-
744 nent waveforms for pairs of SKS and $SKKS$ phases from nearby earthquakes, and pointed out
745 that the variability between phases coming from nearby directions is of similar magnitude to the
746 variability documented in previous studies of $SKS - SKKS$ splitting discrepancies that invoked
747 D'' anisotropy as an explanation. Monteiller & Chevrot (2010) suggested that noise on the trans-
748 verse component waveforms was the most likely explanation for these discrepancies, although
749 other workers have argued that strong discrepancies measured on seismograms with relatively low
750 noise levels cannot be due (solely) to noise (Long, 2009; Lynner & Long, 2014; Long & Lynner,
751 2015). In a later paper, Monteiller & Chevrot (2011) documented differences in estimated SKS
752 splitting parameters for stations in southern California between their measurements, obtained with

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5 753 the splitting intensity method, and previously published measurements; again, they attributed these
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7 754 discrepancies to the presence of noise and to differences in processing and measurement methods.
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9 755 Finally, Lin et al. (2014a) showed that full-wave effects can cause splitting discrepancies between
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11 756 *SKS* and *SKKS* phases for the same event-station pairs of up to $\sim 0.3s$ at certain azimuths for
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13 757 models that only include homogeneous upper mantle anisotropy, as discussed in section 4.

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15 758 What insights do our simulations give us into the interpretations of *SKS* – *SKKS* splitting
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17 759 discrepancies? Our documentation of predicted *SKS* – *SKKS* splitting discrepancies for both
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19 760 upper and lowermost mantle anisotropy models demonstrates that such discrepancies can arise
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21 761 from a physical cause, although of course the effects of noise must be considered for actual data.
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23 762 Similar to Lin et al. (2014a), we found that modest splitting intensity discrepancies (typically
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25 763 up to $0.2s$, with a few values up to $0.4s$) can arise from homogeneous upper mantle anisotropy
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27 764 models, possibly influenced by the interference of other phases in the *SK(K)S* time window.
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29 765 Our simulations showed that much larger *SKS* – *SKKS* splitting intensity discrepancies can
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31 766 arise from lowermost mantle anisotropy models (up to $\sim 1.0s$ or greater for post-perovskite;
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33 767 see Figure 13). Therefore, our work shows that both upper and lowermost mantle anisotropy may
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35 768 potentially contribute to *SKS* – *SKKS* splitting intensity discrepancies for simple, homogeneous
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37 769 anisotropic models. Ongoing work that extends to models that include realistic 3D heterogeneity in
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39 770 anisotropic structure will answer the question of whether our general finding that lowermost mantle
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41 771 anisotropy models predict substantially stronger *SKS* – *SKKS* splitting discrepancies than upper
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43 772 mantle models holds for more complex models. Particularly because *SK(K)S* phases have large
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45 773 zones of sensitivity at the base of the mantle (e.g., Zhao & Chevrot, 2011), and because there
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47 774 may be heterogeneity in lowermost mantle anisotropy on length scales shorter than the Fresnel
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49 775 zones of the *SK(K)S* waves under study, lowermost mantle anisotropy may contribute less to
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51 776 the splitting of *SK(K)S* phases in practice than the results from laterally homogeneous models
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53 777 suggest. Furthermore, for actual, noisy data that reflects complex Earth structure, the effects of
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55 778 noise, phase interference, and complex wave propagation effects on apparent splitting parameters
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57 779 must be carefully considered.

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60 780 Our lowermost mantle simulations also reinforce the notion that discrepant splitting behavior

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between SKS and $SKKS$ phases can arise even in models that feature laterally homogeneous anisotropic structure in the lowermost mantle. While $SKS - SKKS$ splitting discrepancies are often taken to imply a lateral gradient in anisotropy between the respective D'' pierce points of the SKS and $SKKS$ phases (e.g., Long, 2009), our simulations (and previous work by others, including work based on ray theoretical approximations) show that homogeneous anisotropy can give rise to such discrepancies. $SKS - SKKS$ splitting discrepancies, when interpreted in the context of lowermost mantle anisotropy, should therefore be taken to imply a contribution to splitting from anisotropy sampled by one or both phases (e.g., Lynner & Long, 2014; Long & Lynner, 2015; Deng et al., 2017; Reiss et al., 2019), rather than to require a lateral gradient in D'' anisotropy.

8 CONCLUSION

We have presented a new wavefield modeling strategy to introduce the effects of general anisotropy in global models using the pseudo-spectral element code AxiSEM3D. The implementation of arbitrary anisotropy is accomplished by describing the elastic properties of the seismic domain in terms of the full elastic tensor with 21 independent coefficients. We have carried out global wavefield simulations for models that include anisotropy in the upper and lowermost mantle, reaching frequencies as high as $0.2Hz$ with relatively modest computational resources. We benchmarked our implementation against known reference solutions for simple upper mantle models, and then investigated the behavior of $SK(K)S$ phases for models that include upper mantle anisotropy (in HTI and orthorhombic geometries) as well as those that include anisotropy at the base of the mantle (for possible bridgmanite and post-perovskite CPO scenarios). We carried out shear wave splitting analysis (both the traditional transverse component minimization method and the splitting intensity method) on synthetic waveforms for the suite of global anisotropic models. These tests revealed that shear wave splitting, as manifested in the full-wavefield simulations, behaves similarly to the predictions of ray theory to first order, but some departures from ray theoretical behavior (due to full waveform effects) are found. Our results indicate that although some $SKS - SKKS$ splitting intensity discrepancies arise from anisotropic upper mantle models, par-

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5 808 ticularly when low symmetry classes (e.g. orthorhombic) are considered, these are of modest am-
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7 809 plitude (generally less than $0.2s$, with a few discrepancies up to $0.4s$). On the other hand, we find
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9 810 that realistic lowermost mantle anisotropy scenarios can cause significant splitting (up to $\sim 1s$)
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11 811 of $SK(K)S$ waveforms when full wave propagation is taken into account, with $SKS - SKKS$
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13 812 splitting intensity discrepancies up to $\sim 1s$ or greater. The cross-validation test with the discrete
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15 813 spectral element solver SPECFEM3D_GLOBE further highlights the capability of AxiSEM3D to
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17 814 handle increasingly complex Earth models, including those with arbitrary anisotropic symmetry,
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19 815 without loss of accuracy and with accessible computing resources. In the future, AxiSEM3D will
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21 816 be used to investigate the behavior of $SK(K)S$ phases in the presence of lateral heterogeneity
22
23 817 in anisotropic structure, paving the way for more realistic consideration of full wavefield effects
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25 818 when interpreting shear wave splitting measurements, particularly due to D'' anisotropy.
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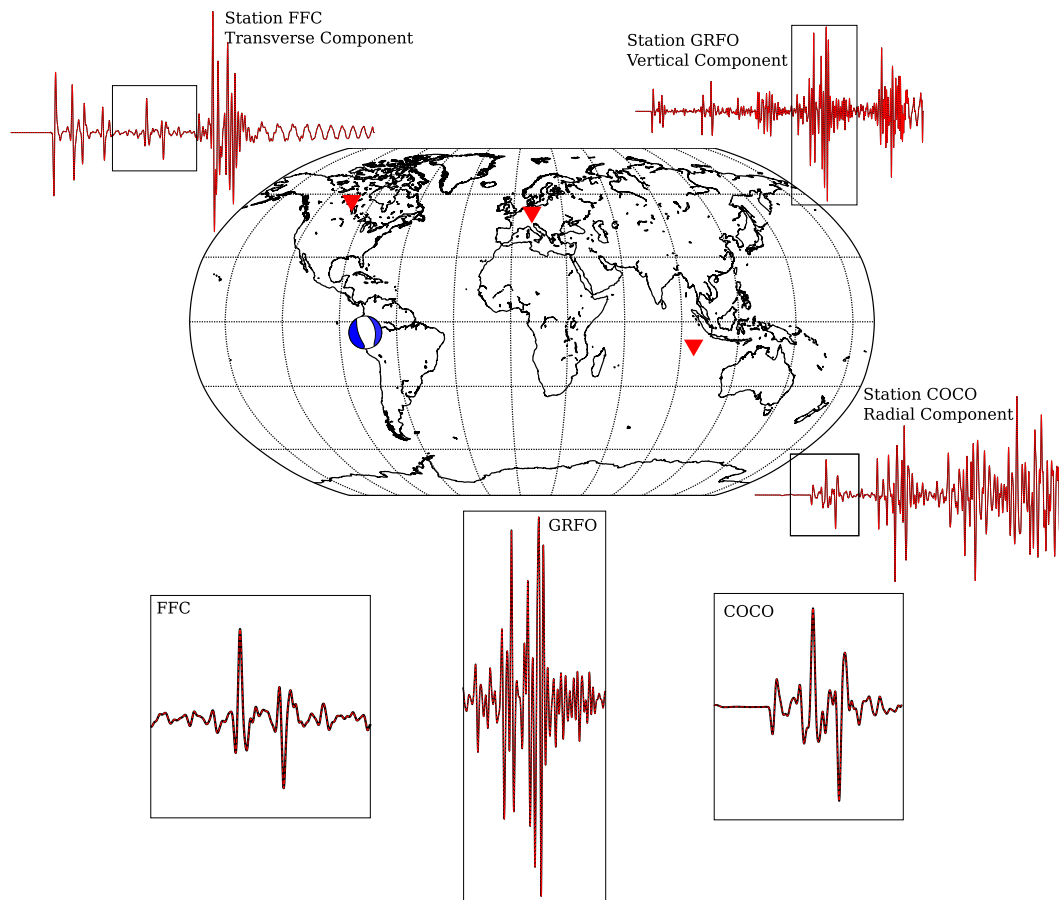


Figure 1. Map showing the source-receiver configuration settings and the synthetic results for the self-benchmarking tests computed using AxiSEM3D with the anisotropic PREM model (black seismograms) and the solution obtained using a description of the anisotropic PREM as C_{ijkl} (red seismograms). Three representative pairs of synthetic seismograms are plotted at three different locations. A close-up view of the synthetics is shown in the three boxes at the bottom of the figures for three different sections of the time series.

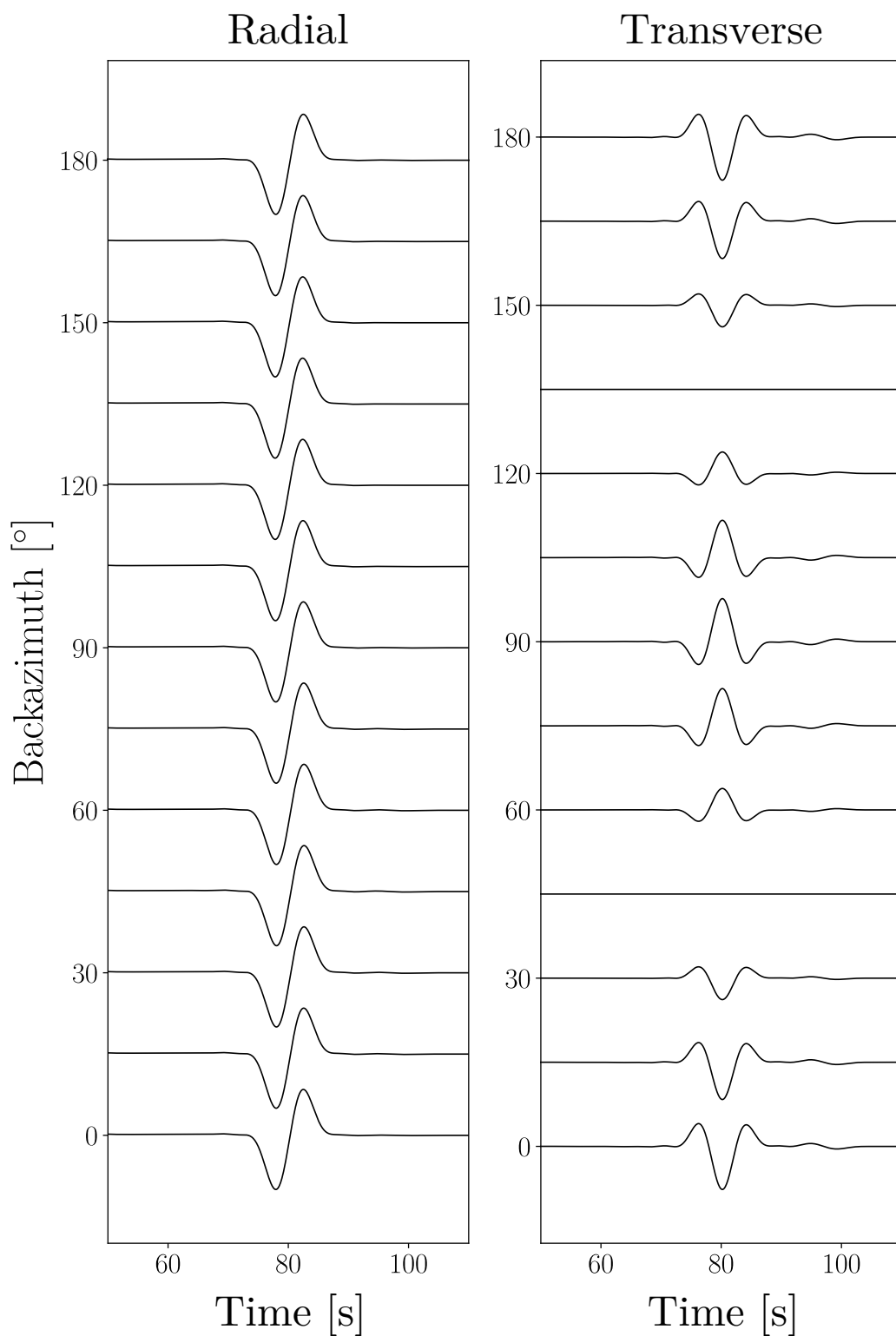


Figure 2. Synthetic seismograms computed for a vertically propagating shear wave in a 30 km thick horizontally transverse isotropic (HTI) layer with 3% anisotropy. The amplitude of the transverse component is amplified by a factor of 10.

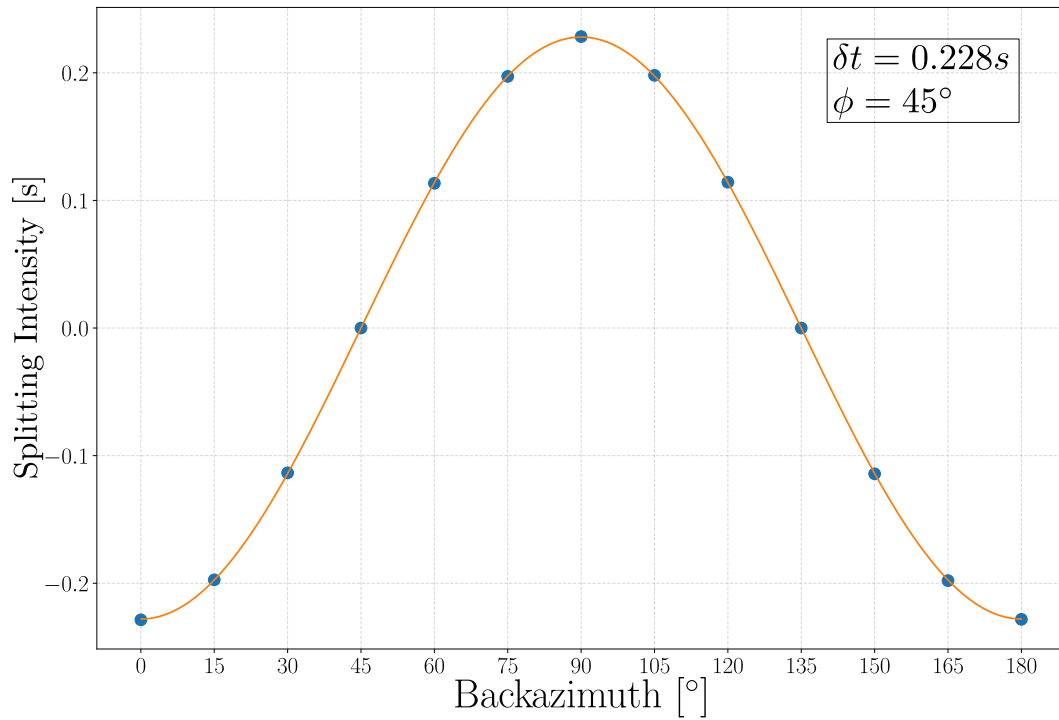


Figure 3. Result of the shear wave splitting intensity analysis performed on the synthetic seismograms of Fig. 2. The splitting intensity vector (blue dots) is computed following Chevrot (2000, eq. A4). The orange curve is the least square fit solution $s = \sin 2\theta(\phi - \phi_0)$, calculated over possible pairs of splitting intensity parameters ϕ and δt , that minimizes the misfit between the synthetic and the theoretical splitting intensity solution. The amplitude of the curve is $0.228s$ and its phase is 45° .

Compare this figure with (Chevrot, 2000, Figure 3).

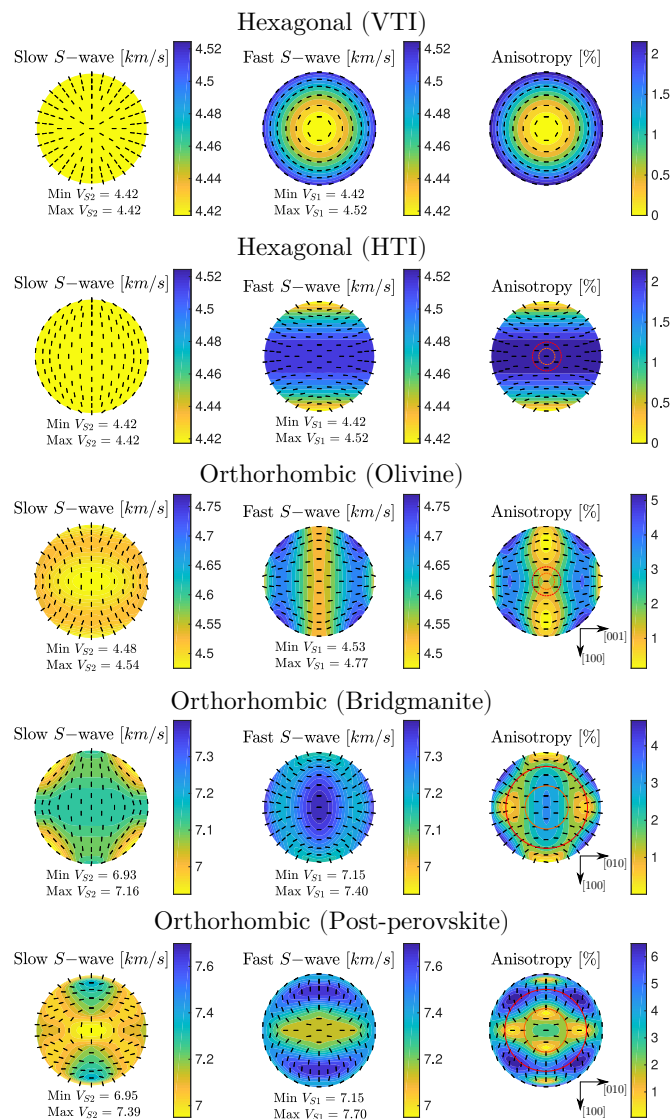


Figure 4. Polar view representation of different elasticity tensors belonging to different symmetry classes (from top to bottom hexagonal (VTI, HTI) and orthorhombic (olivine, bridgmanite, post-perovskite) used to describe the elastic and geometrical properties of the anisotropic regions tested in the different simulation settings. The magnitude of the slow and fast shear wave velocity and its polarization direction (black thick marks) are shown on the left and middle plots respectively. A measure of the anisotropy is shown on the right-hand side plot along with the fast shear wave polarization direction (black thick marks). The VTI, HTI, and olivine tensors are used in examples where the anisotropy is localized in the upper mantle whereas the bridgmanite and post-perovskite are used in the D'' . The orientation of the orthorhombic crystals are specified by the 2 orthogonal crystallographic axes that lie in the horizontal plane. The dashed concentric circles in the HTI and olivine tensors mark the region of the piercing points for the SKS (orange) and $SKKS$ (red) phases in the upper mantle, while the solid circles indicate the region of the piercing points for the bridgmanite and post-perovskite crystals in the D'' layer.

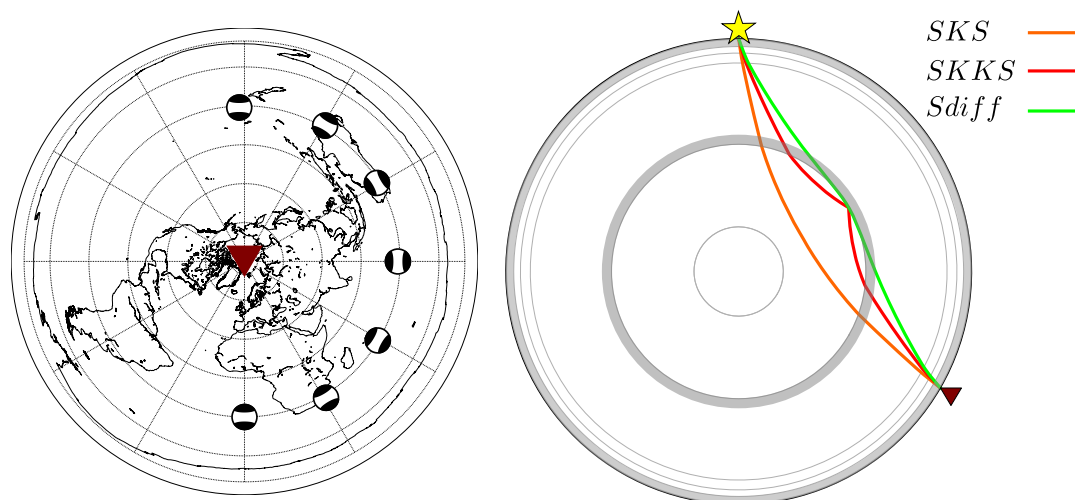


Figure 5. Left: Source-receiver distribution for global seismic wavefield simulations in anisotropic Earth-like models. The red triangle marks the center of the array of receivers located at the North Pole at a distance of 120° from the seismic sources. The sources span a fan of 180° along the azimuthal direction with inter-spacing of 15° .

Right: Diagram of ray-theoretical paths of seismic S phases commonly used in shear wave splitting analysis. The paths for SKS , $SKKS$ and S_{diff} are shown as reference for a source (star) and the receiver (red triangle) at 120° of epicentral distance. The shaded grey areas represent the two anisotropic regions between 24km and 220km in the upper mantle and 250km at the top of the Core Mantle Boundary in the D'' . Paths are calculated using **TauP** (Crotwell et al., 1999) for the PREM velocity model (Dziewonski & Anderson, 1981)

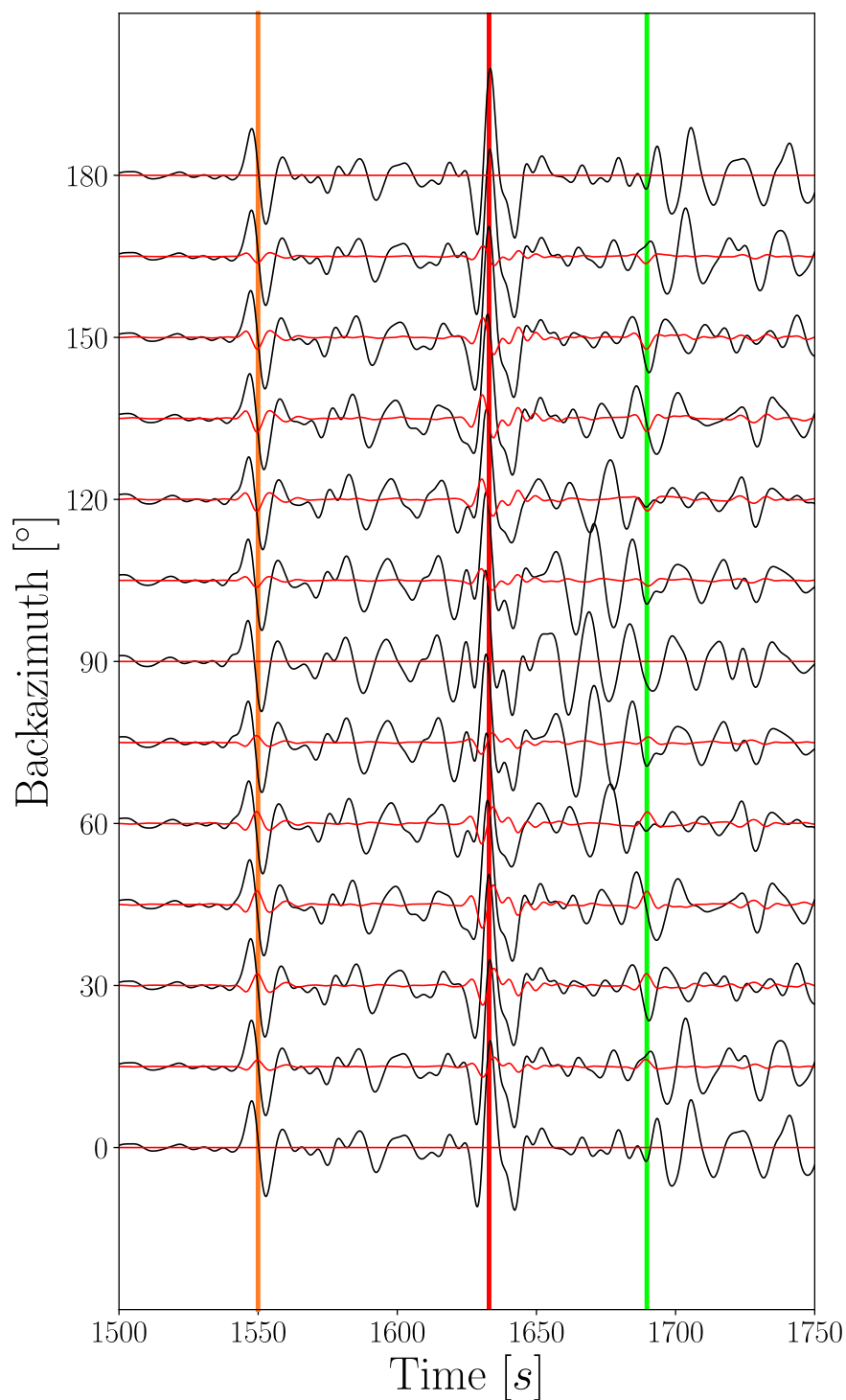
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Figure 6. Synthetic seismograms computed for the PREM model with a HTI layer localized between 24 and 220km plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 120°. The radial component is shown in black and the transverse is shown in red. Predicted arrival time for *SKS*, *SKKS* and *S_{dif}* are shown as the vertical orange, red and green bars respectively.

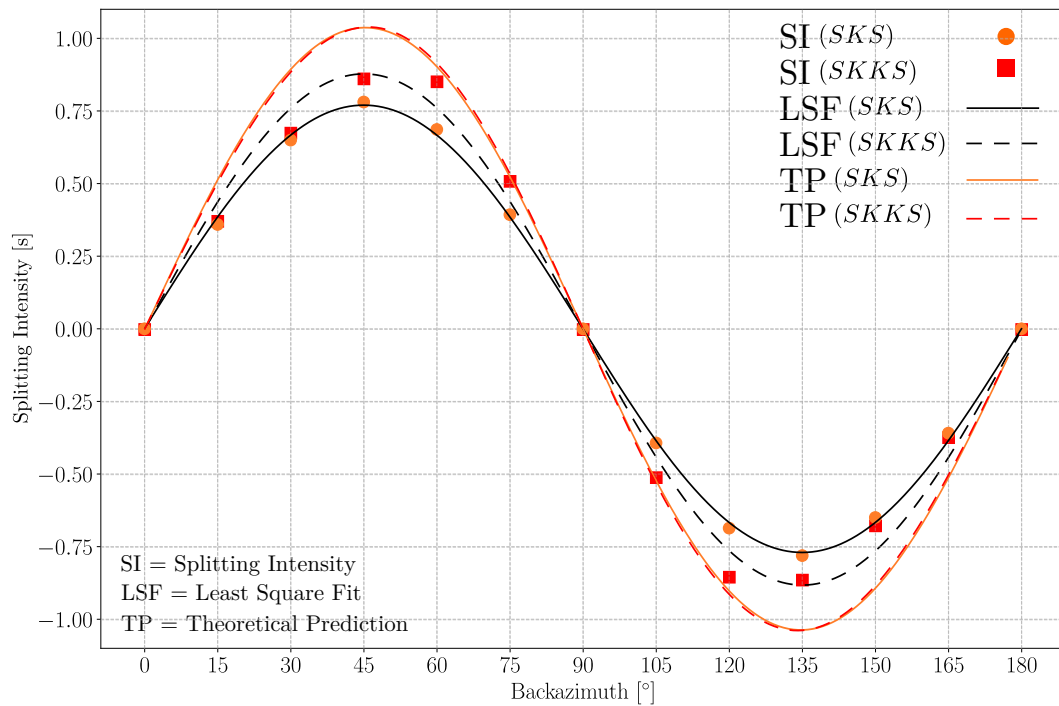


Figure 7. Shear wave splitting intensity analysis performed on the synthetic seismograms evaluated for the PREM model with HTI (see Figure 4, second row) localized between 24 and 220 km depth. Values of the splitting intensity vector calculated for SKS are shown as orange circles while for $SKKS$ are shown as red squares. The black solid and dashed curves represent the two least square fit solutions for SKS and $SKKS$ respectively. The orange solid and dashed red lines are the ray-theoretical predicted splitting intensity vectors for SKS and $SKKS$ respectively, evaluated by solving the Christoffel equation for an elasticity tensor that was averaged over the depth range of a ~ 200 km thick anisotropic layer in the upper mantle.

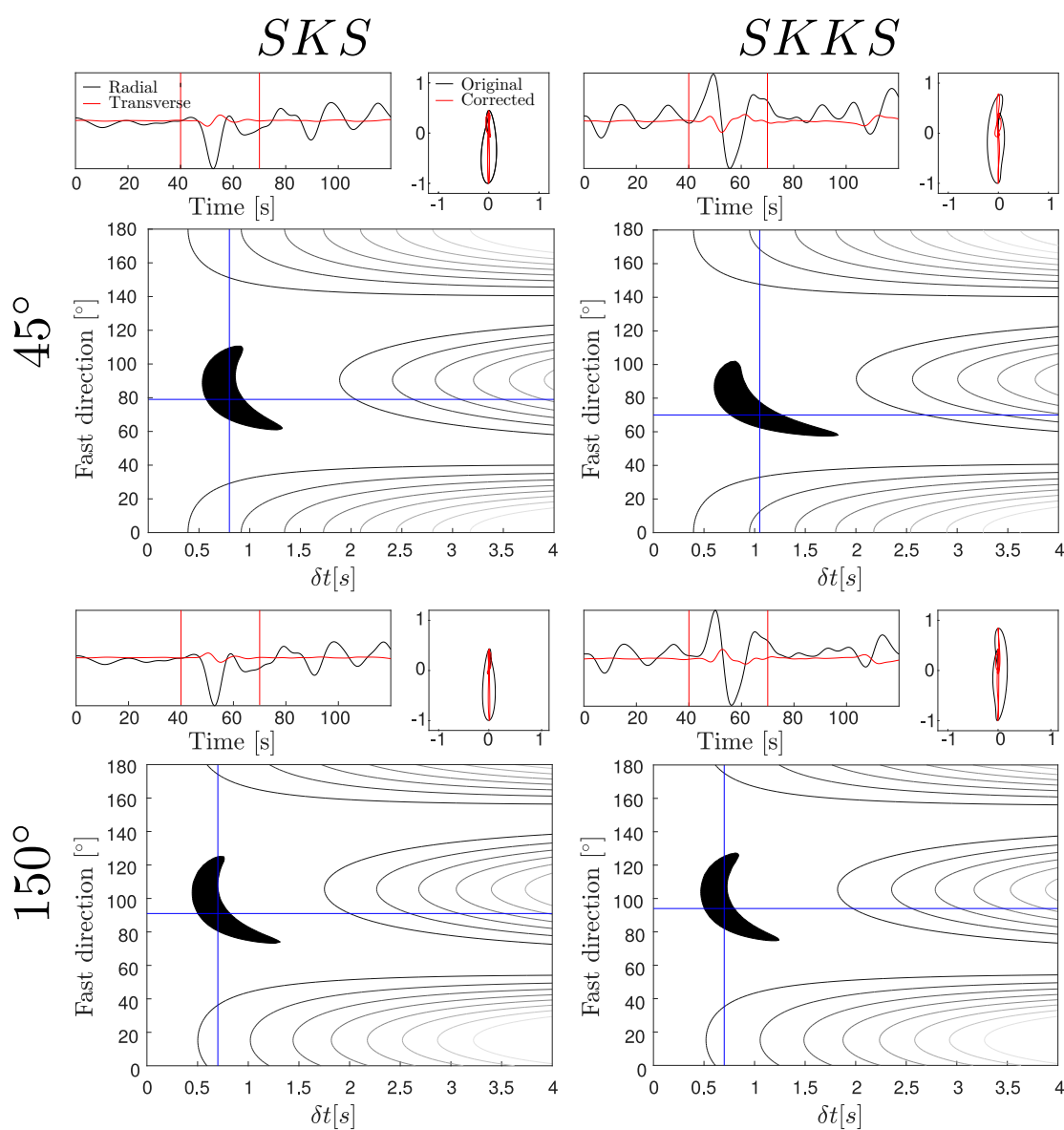
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Figure 8. Shear wave splitting analysis performed using the transverse component minimization method (Silver & Chan, 1991) on the *SKS* (left) and *SKKS* (right) phases for 2 representative backazimuthal propagation directions of the incoming seismic energy (45° and 150°). The synthetic seismograms are evaluated for a PREM model with a HTI layer localized between 24 and 220 km as in Figure 7. Each panel is composed by 3 plots: On the top left a 120 s long time-window with the radial (black) and transverse (red) component synthetic seismograms. The vertical red bars mark the phases on which the analysis is performed. On the top right corner we show the original and the corrected particle motion. In the bottom plot we show the energy map with the calculated splitting parameters and the 95% confidence interval as the shaded black area. For this configuration *SKS* and *SKKS* yield a comparable results for both fast direction axis and delay-time.

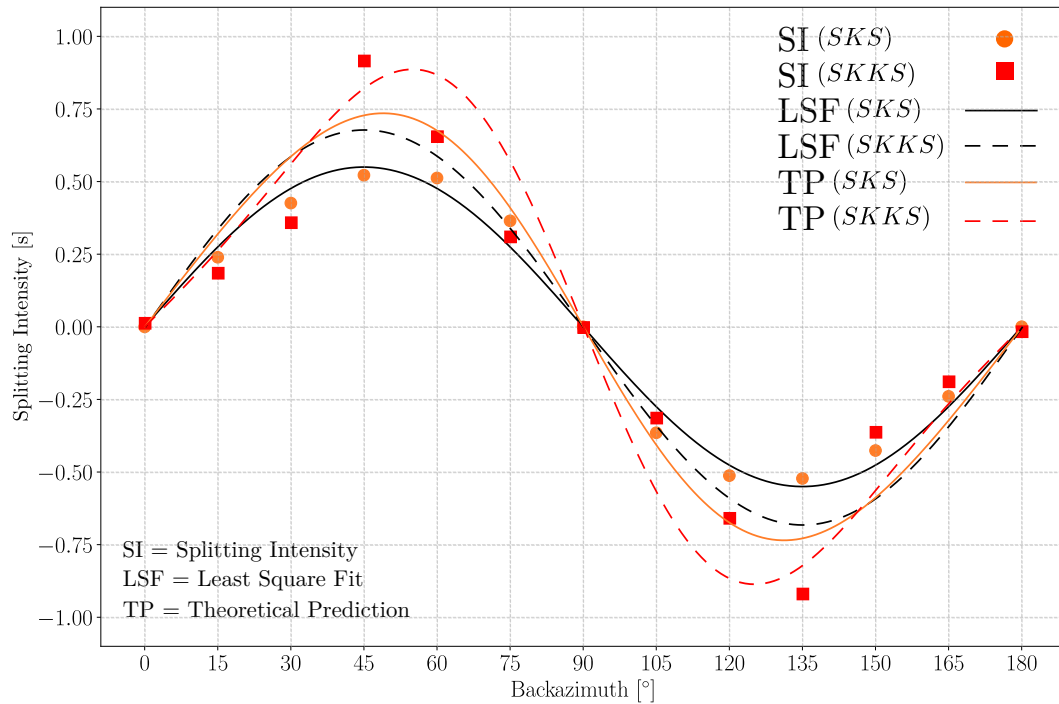


Figure 9. Same as Fig. 7 but with the anisotropic region in upper mantle between 24 and 220km described with an orthorhombic elasticity tensor (see Fig. 4, third row). The elastic properties of the tensor are the result of a linear mixture of 30% of San Carlos single crystal olivine from Abramson et al. (1997) and 70% PREM. The least square fit solutions and the theoretical predicted splitting intensity vectors are also plotted with the same color code used in Fig. 7.

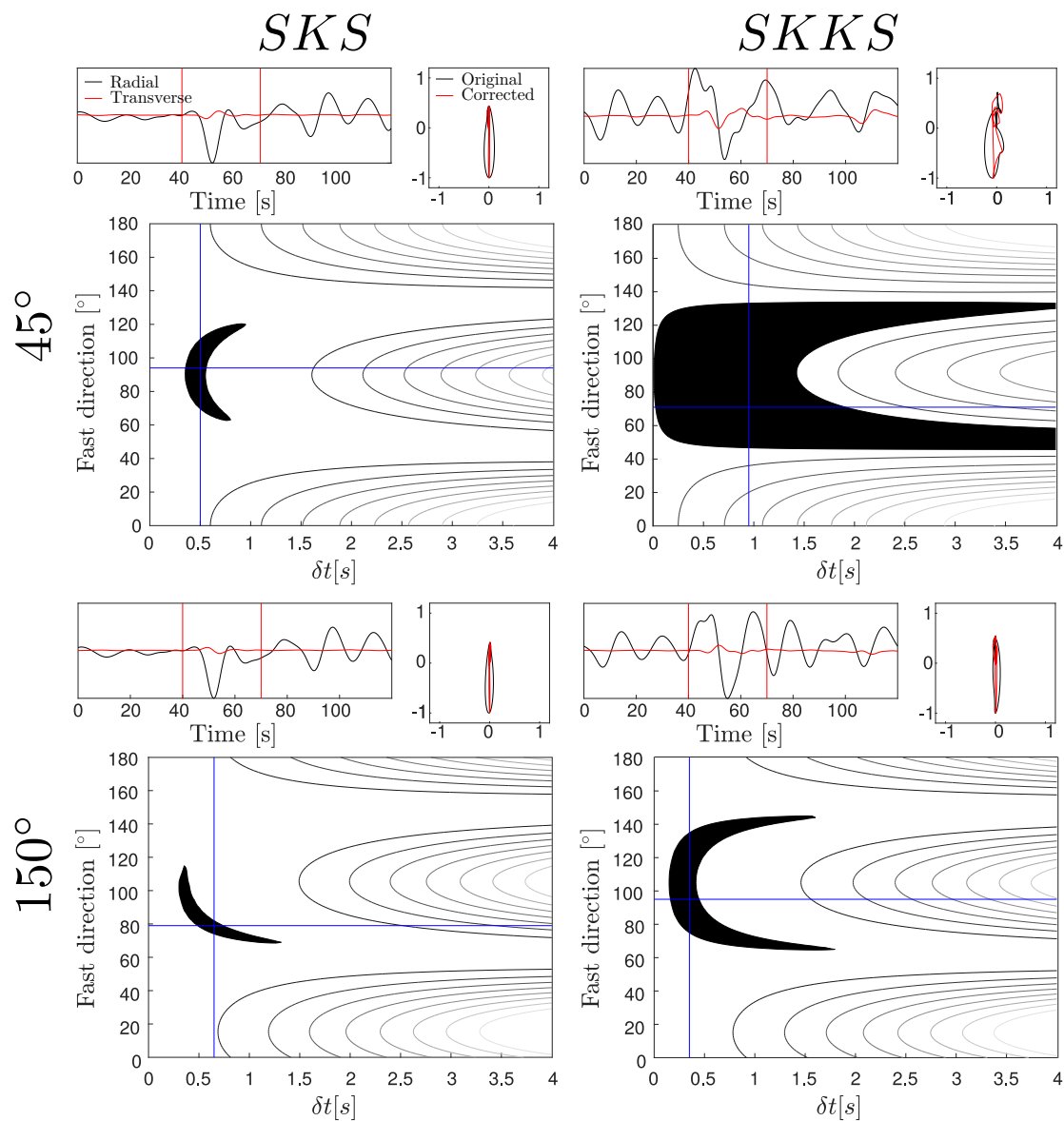
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Figure 10. Same as Fig. 8 but with the anisotropic region in the upper mantle between 24 and 220 km being described with an elasticity tensor of orthorhombic class which mimics the elastic properties of a horizontally strained olivine (see Figure 4, third row). The splitting parameter estimates are less well constrained, with larger formal errors, than for the hexagonal upper mantle simulation shown in Fig. 8

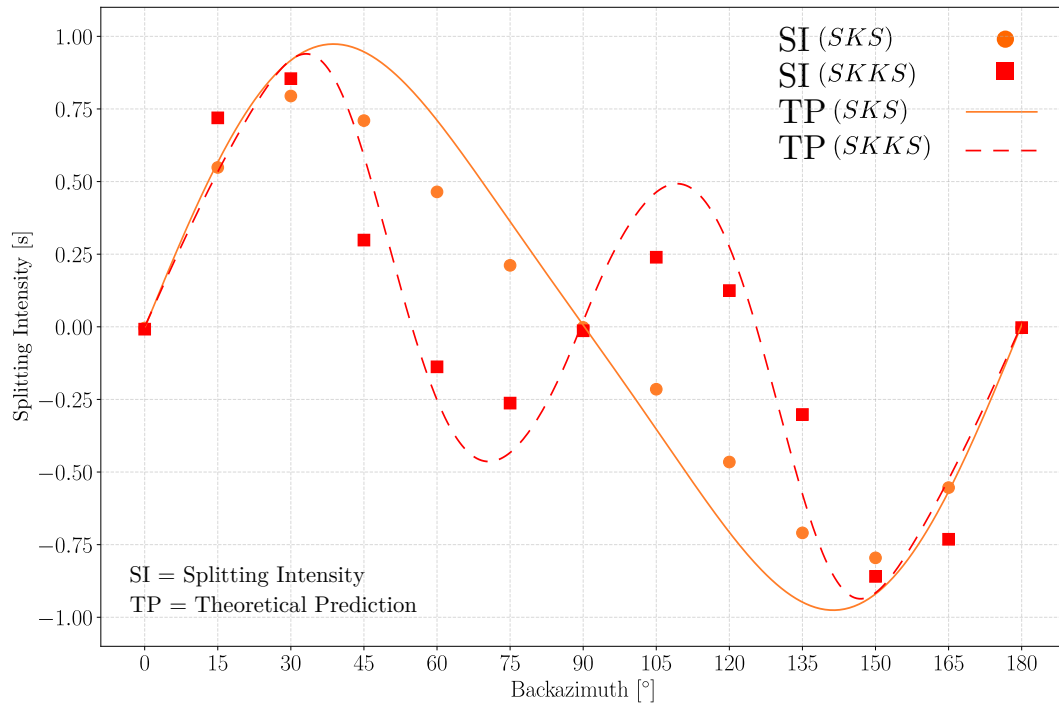


Figure 11. Splitting intensity analysis performed on synthetic seismograms evaluated for the PREM model with anisotropy localized at the base of the lower mantle in a 250km thick D'' layer. The elasticity tensor is composed of a linear mixture of 30% bridgmanite from Wookey et al. (2005b) and 70% isotropic PREM (see Figure 4, fourth row). The splitting intensities for SKS (orange circles) and $SKKS$ (red squares) are similar for some specific backazimuth propagation (0° to 30° and 150° and 180°) but show discrepancies between 45° and 135° .

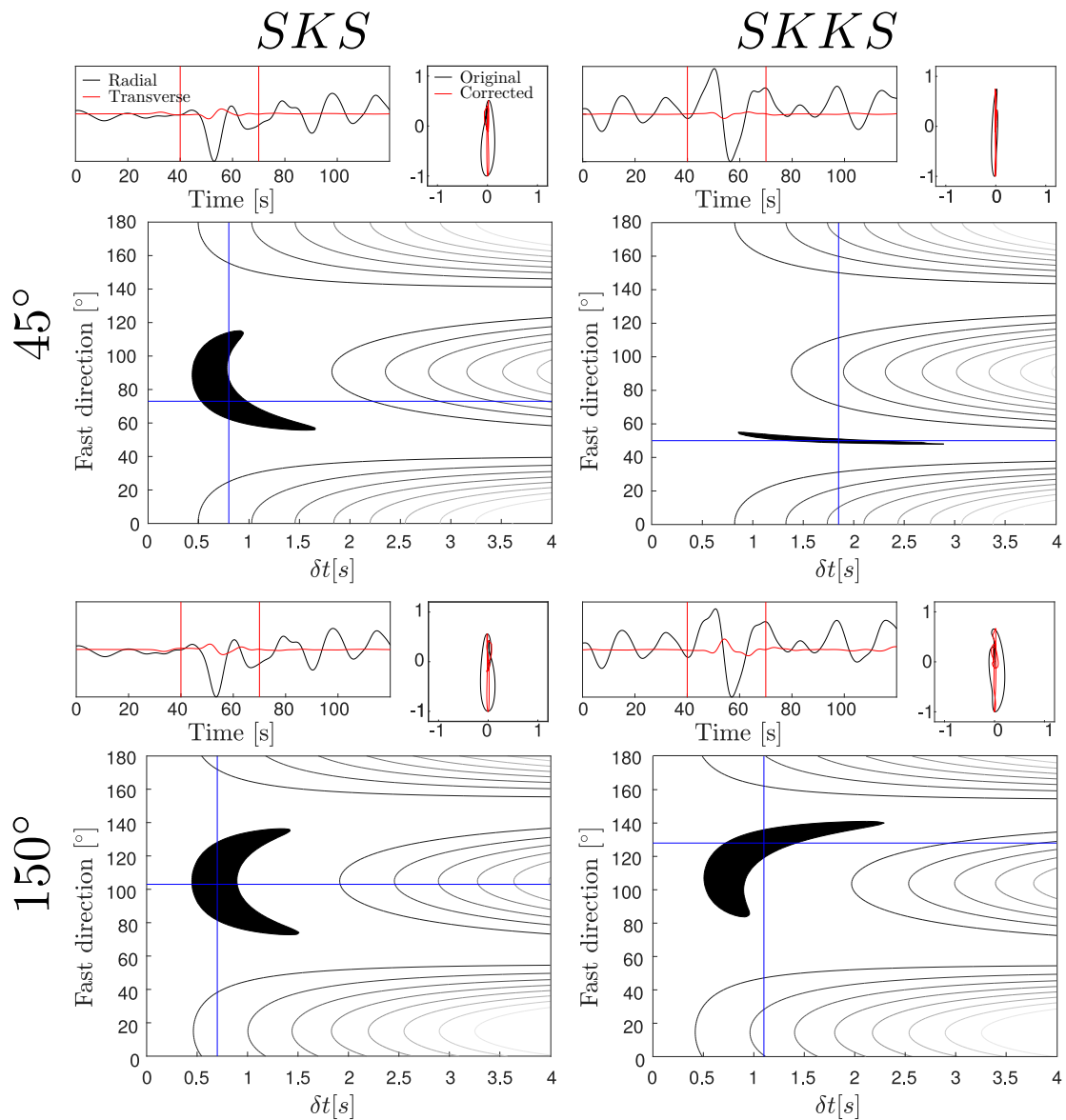
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Figure 12. Same analysis presented in Figure 8 but with orthorhombic style anisotropy in the form of a bridgmanite crystal localized in a 250km thick D'' layer. As for the splitting intensity analysis, we can see a similar result for *SKS* and *SKKS* in the case of seismic energy coming from 150° backazimuthal direction and discrepant results for 45° . We caution, however, that for the 45° backazimuth the *SKKS* splitting measurement is not well constrained.

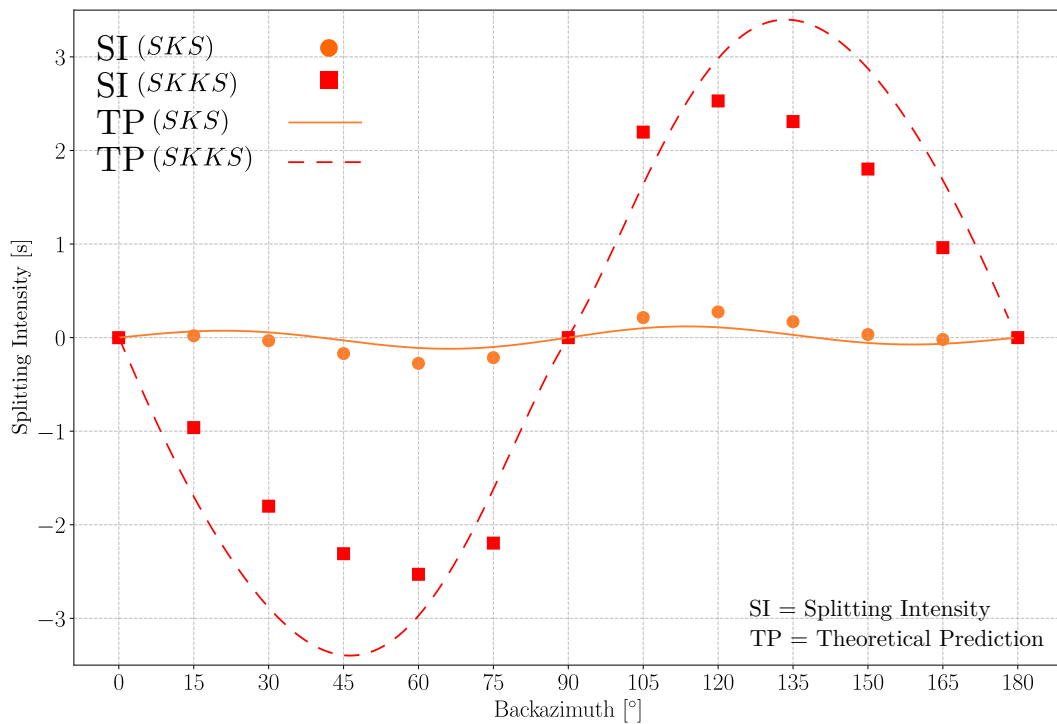


Figure 13. Same as Figure 11 but with an elasticity tensor describing a post-perovskite mineralogy (see Figure 4, bottom row). Strong discrepancies between SKS and $SKKS$ are observed in this case for all the backazimuthal propagation directions with a very strong splitting observed for the $SKKS$ and no splitting (null measurement) observed for the SKS .

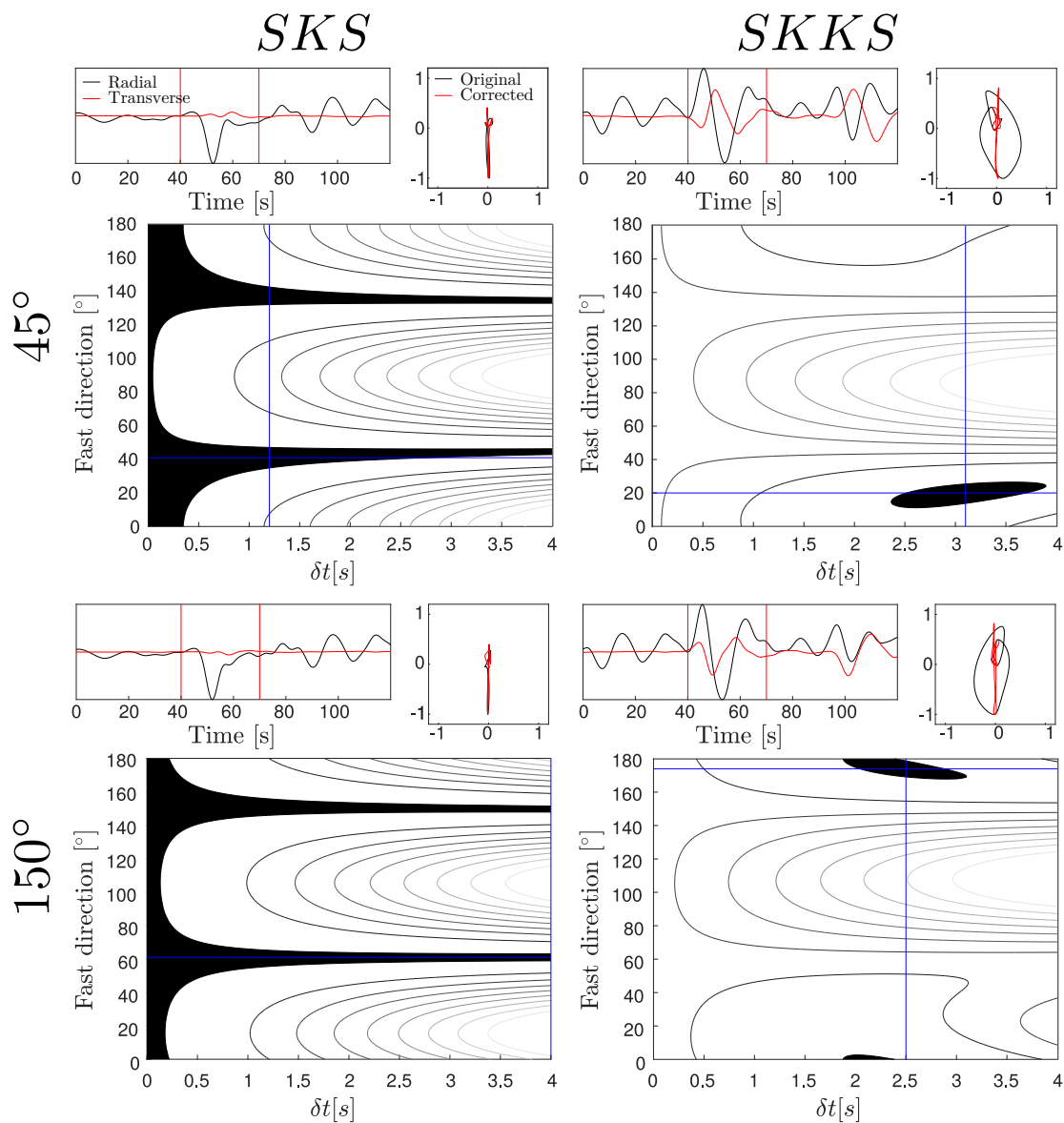


Figure 14. Same as Figure 8 but with the anisotropy localized in a 250km thick D'' layer. The elasticity tensor used for this test is the one shown in Fig. 4, bottom row. Also in this case we notice a strong discrepancies between SKS and $SKKS$, with the SKS yielding null measurements.

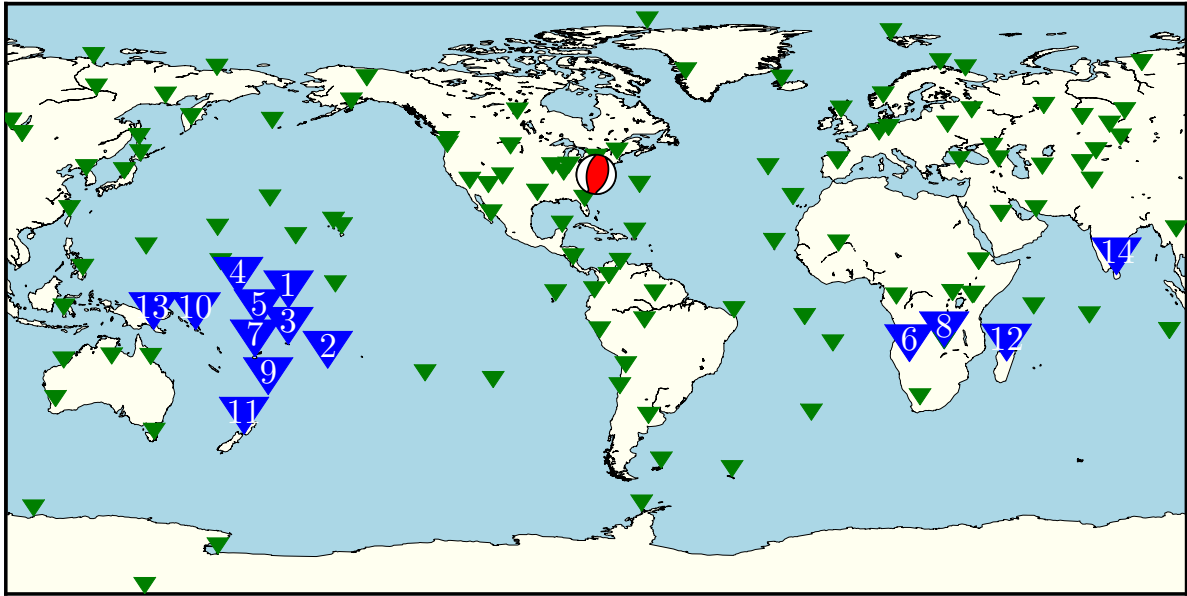


Figure 15. Source and receiver configuration for the AxiSEM3D-SPECFEM3D_GLOBE cross-verification exercise. The earthquake source is located in Virginia, US (37.91°N , 77.93°W), at a depth of 12km (event ID 201108231751A). Receivers of the Global Seismographic Network (GSN) are shown as green triangles. The large blue triangles with numbers are 14 representative stations where the synthetic seismograms displayed in Figure 16 are computed.

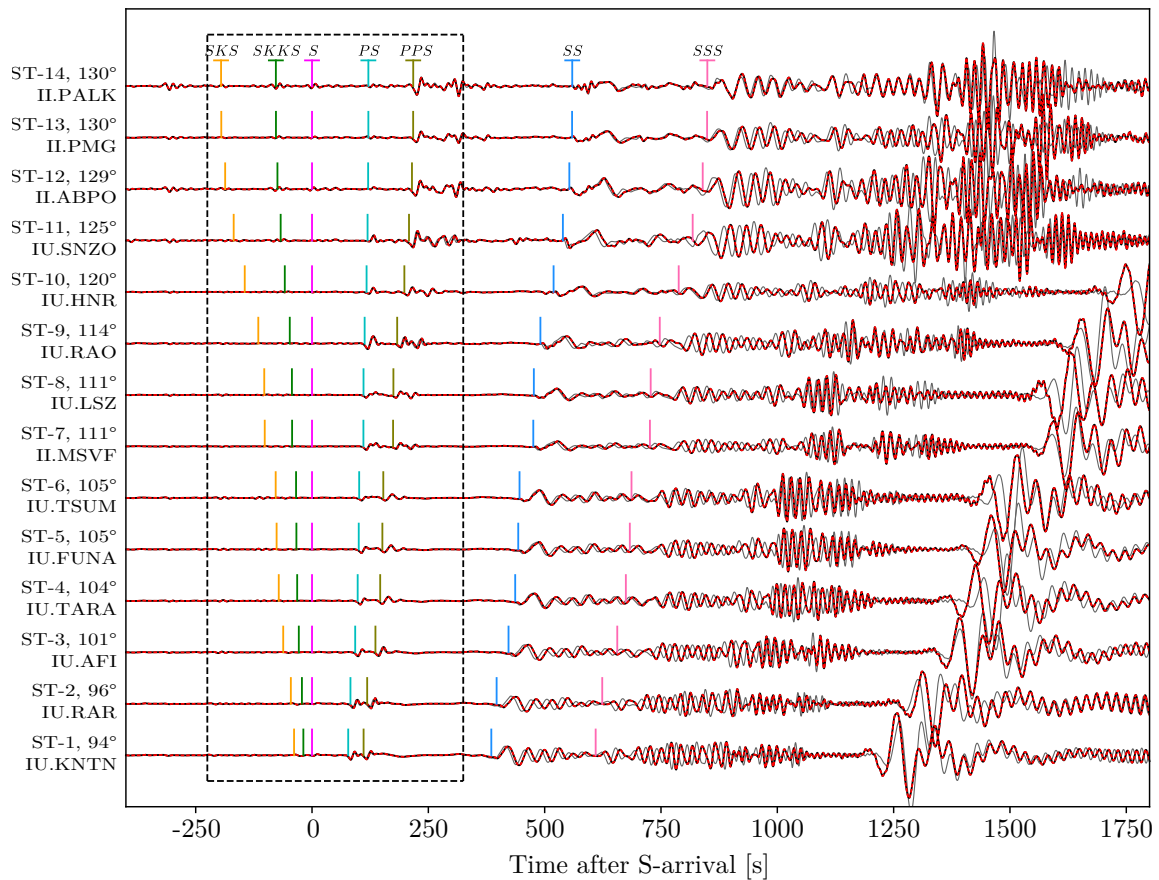


Figure 16. Comparison of synthetic waveforms computed by SPECFEM3D_GLOBE (black) and AxiSEM3D (dashed red) for the tomographic model of Montagner (2002), which includes 3D anisotropy. Synthetic seismograms for the radially anisotropic PREM model are shown in light grey for comparison. We show vertical displacements for all three cases. The receiver locations are shown with blue triangles in Figure 15. Radial and transverse components are shown in Figure 17 in the time-window bounded by the dashed black line that includes *SKS* and *SKKS* phases.

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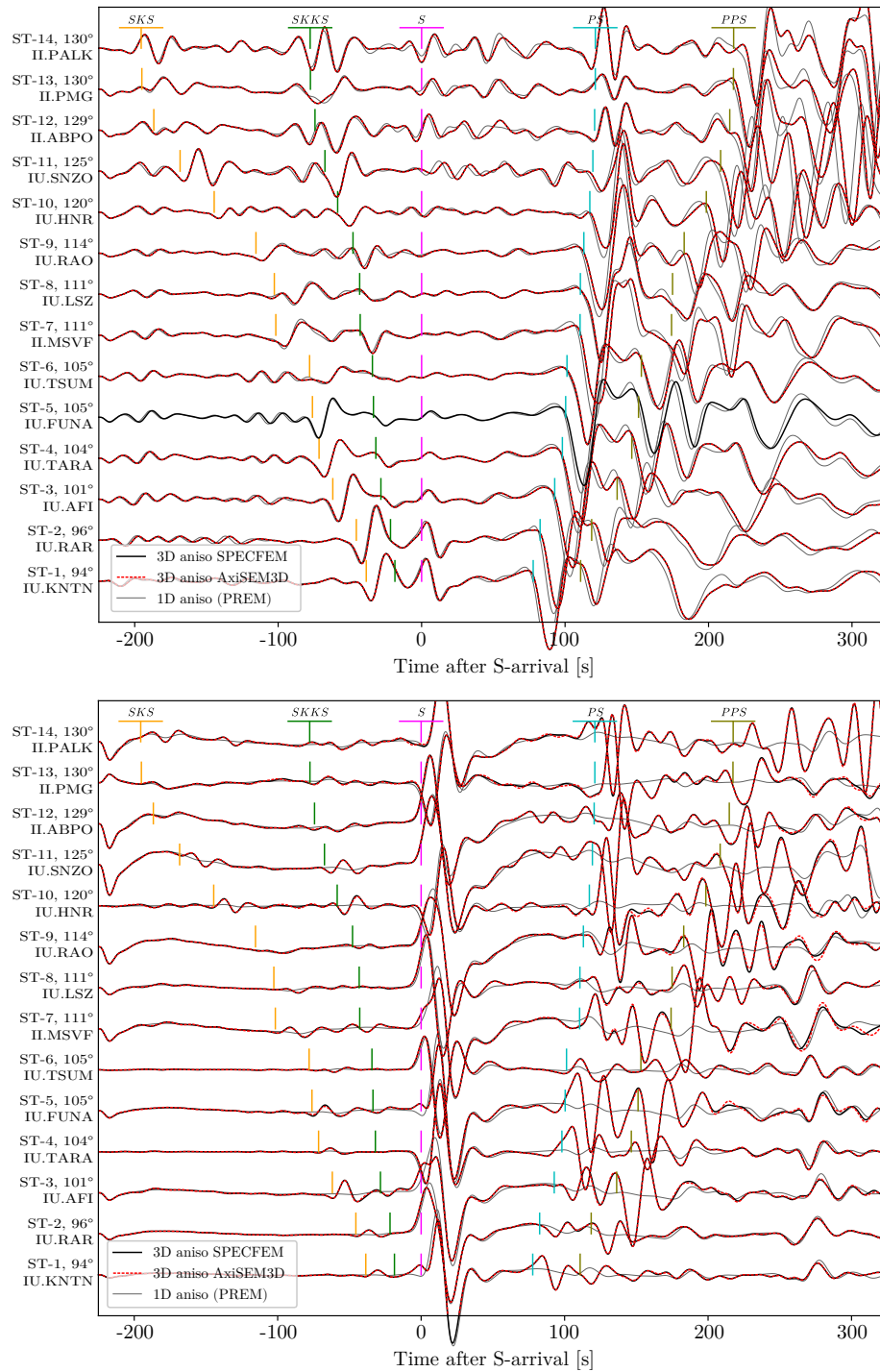


Figure 17. Details of the radial (top panel) and transverse (bottom panel) components of the synthetic waveforms presented in Figure 16. Both AxiSEM3D (dashed red) and SPECFEM3D_GLOBE (black) waveforms show clear evidence of splitting energy between radial and transverse components as a function of the back-azimuth.

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4 2 *Tesoniero et al., 2019*

5 1 **GENERAL INFORMATION**

6 In this supplementary material we present more record section plots showing synthetic seismo-
7 grams evaluated for the 4 different anisotropic models described in the main text at 3 different
8 epicentral distance (100°, 110° and 120°) and for the AxiSEM3D-SPECFEM3D_GLOBE bench-
9 mark test.

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SI: Full wave sensitivity of $SK(K)S$ phases to arbitrary anisotropy in the upper and lower mantle 3

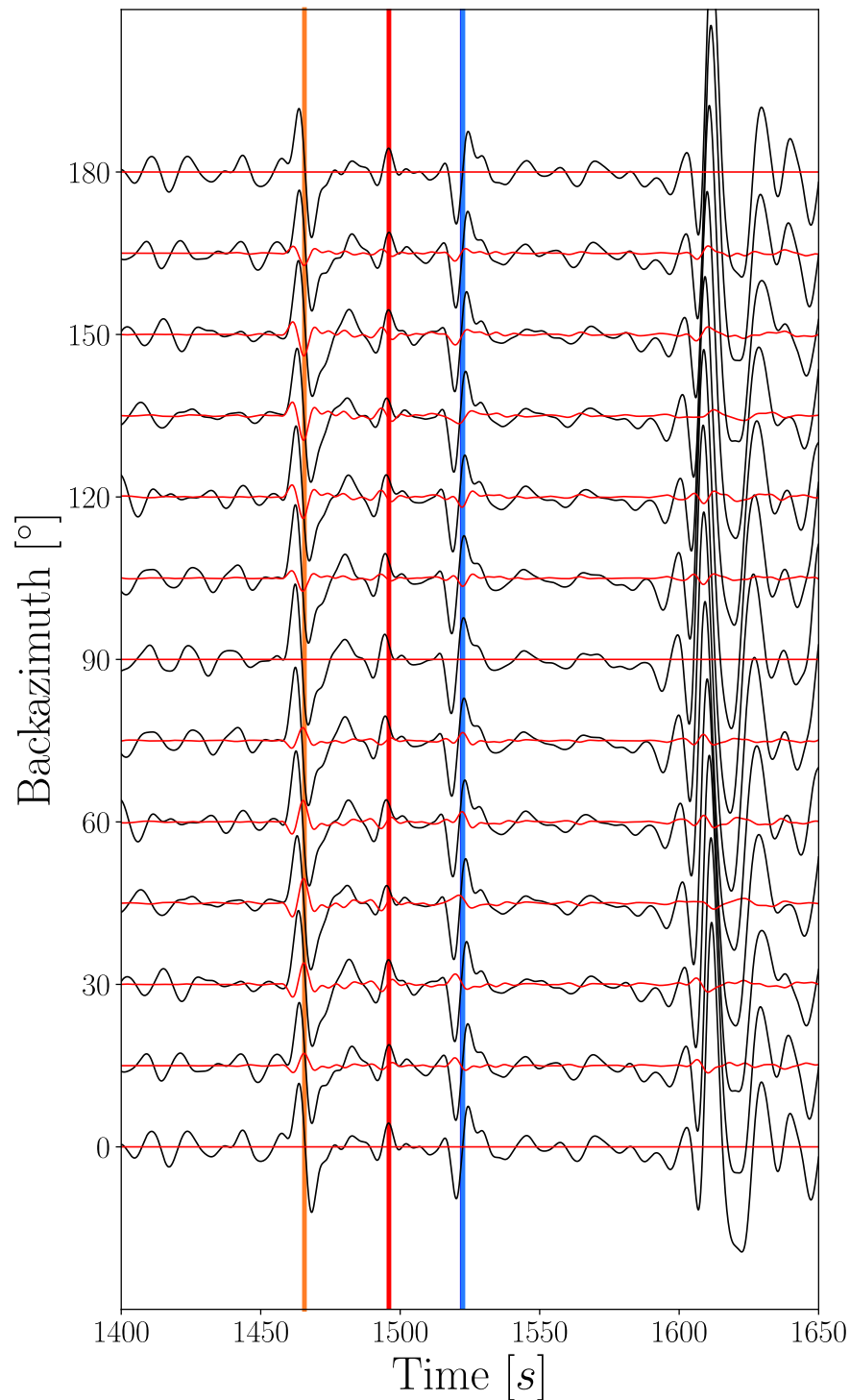


Figure S1. Synthetic seismograms computed for the PREM model with a HTI layer localized between 24 and 220km plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 100° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and ScS are shown as the vertical orange, red and blue bars respectively.

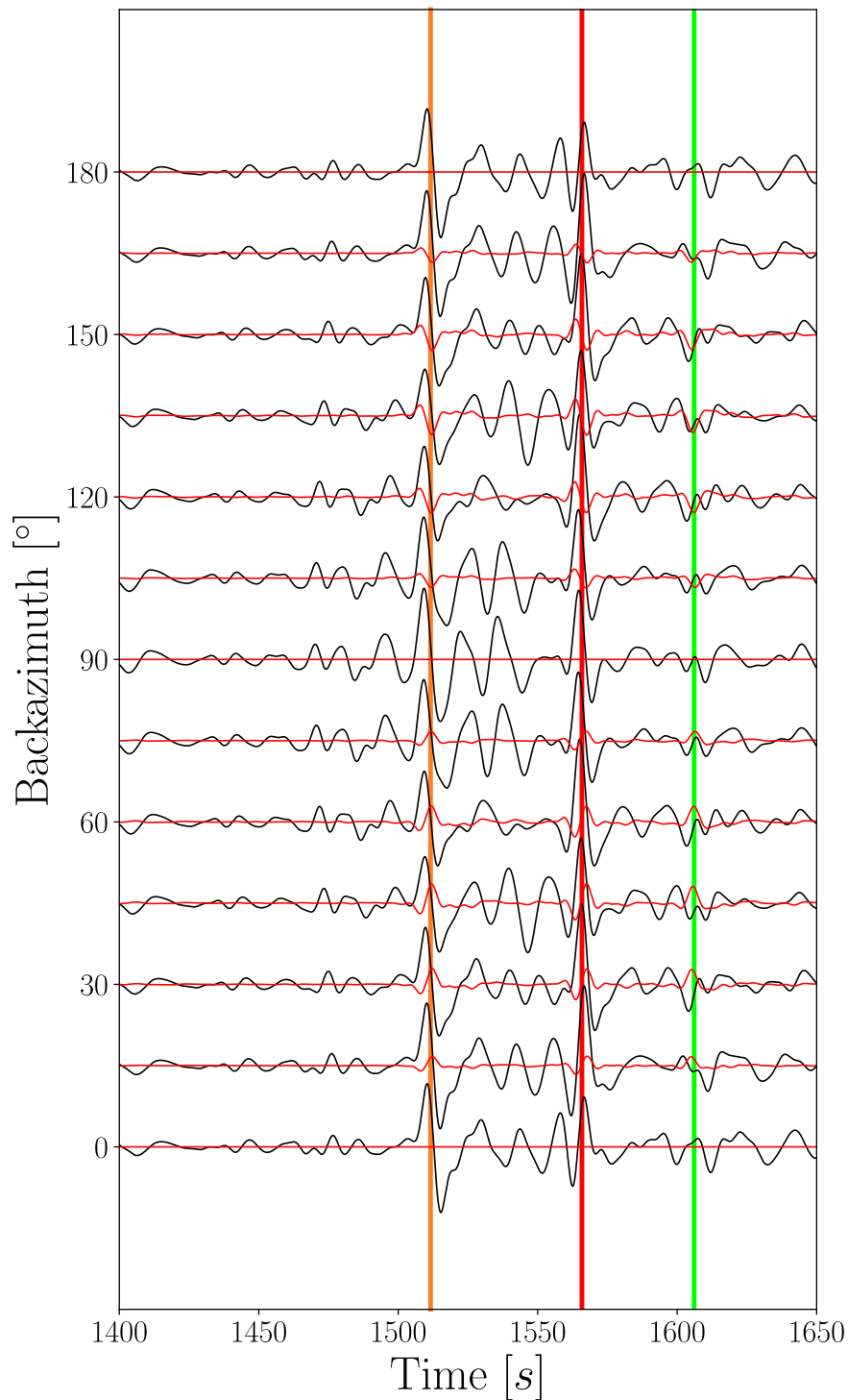
4 *Tesoniero et al., 2019*

Figure S2. Synthetic seismograms computed for the PREM model with a HTI layer localized between 24 and 220km plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 110° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for *SKS*, *SKKS* and *S_{dif}* are shown as the vertical orange, red and green bars respectively.

SI: Full wave sensitivity of $SK(K)S$ phases to arbitrary anisotropy in the upper and lower mantle 5

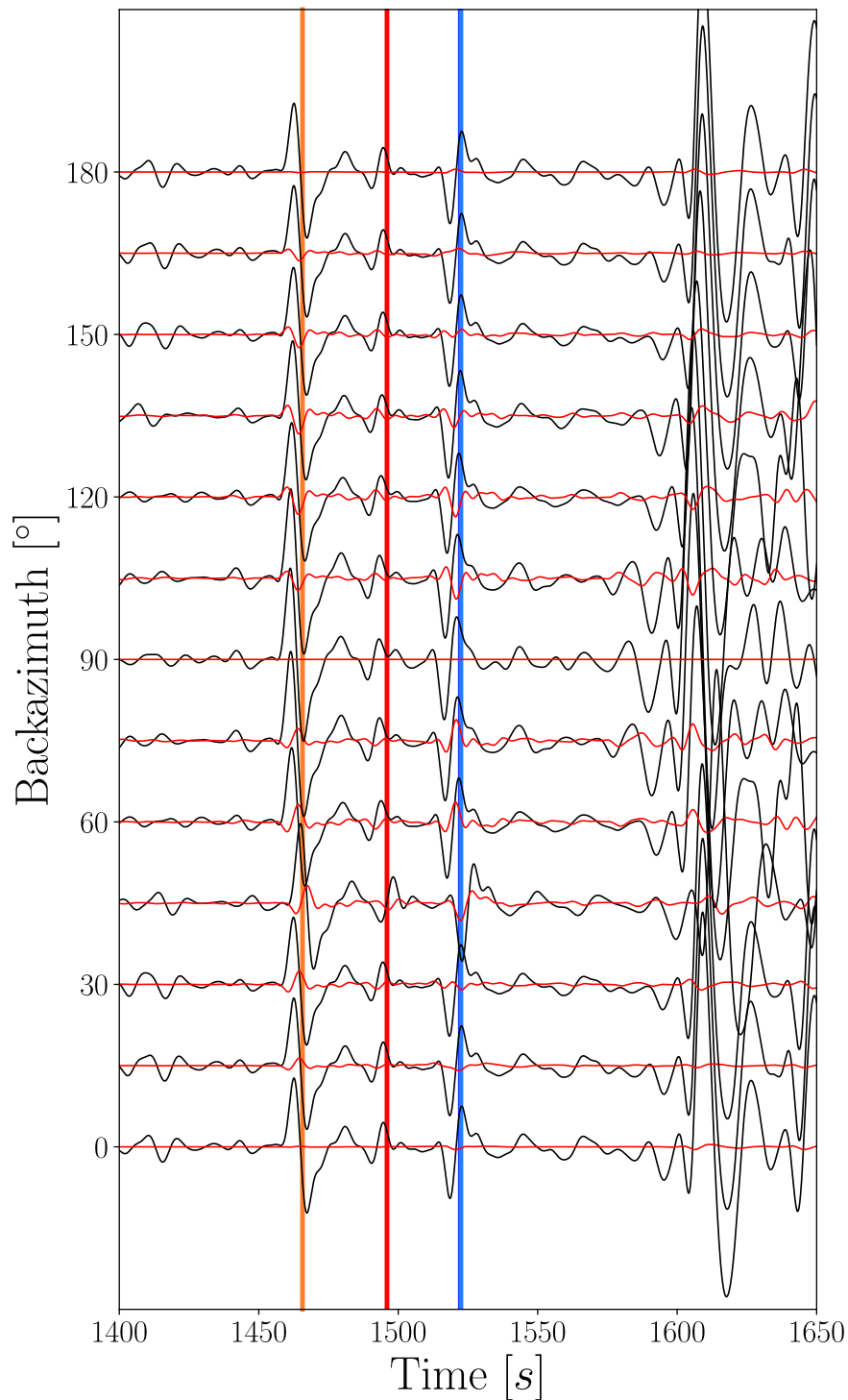


Figure S3. Synthetic seismograms computed for the PREM model with olivine style anisotropy localized between 24 and 220km plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 100° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and Scs are shown as the vertical orange, red and blue bars respectively.

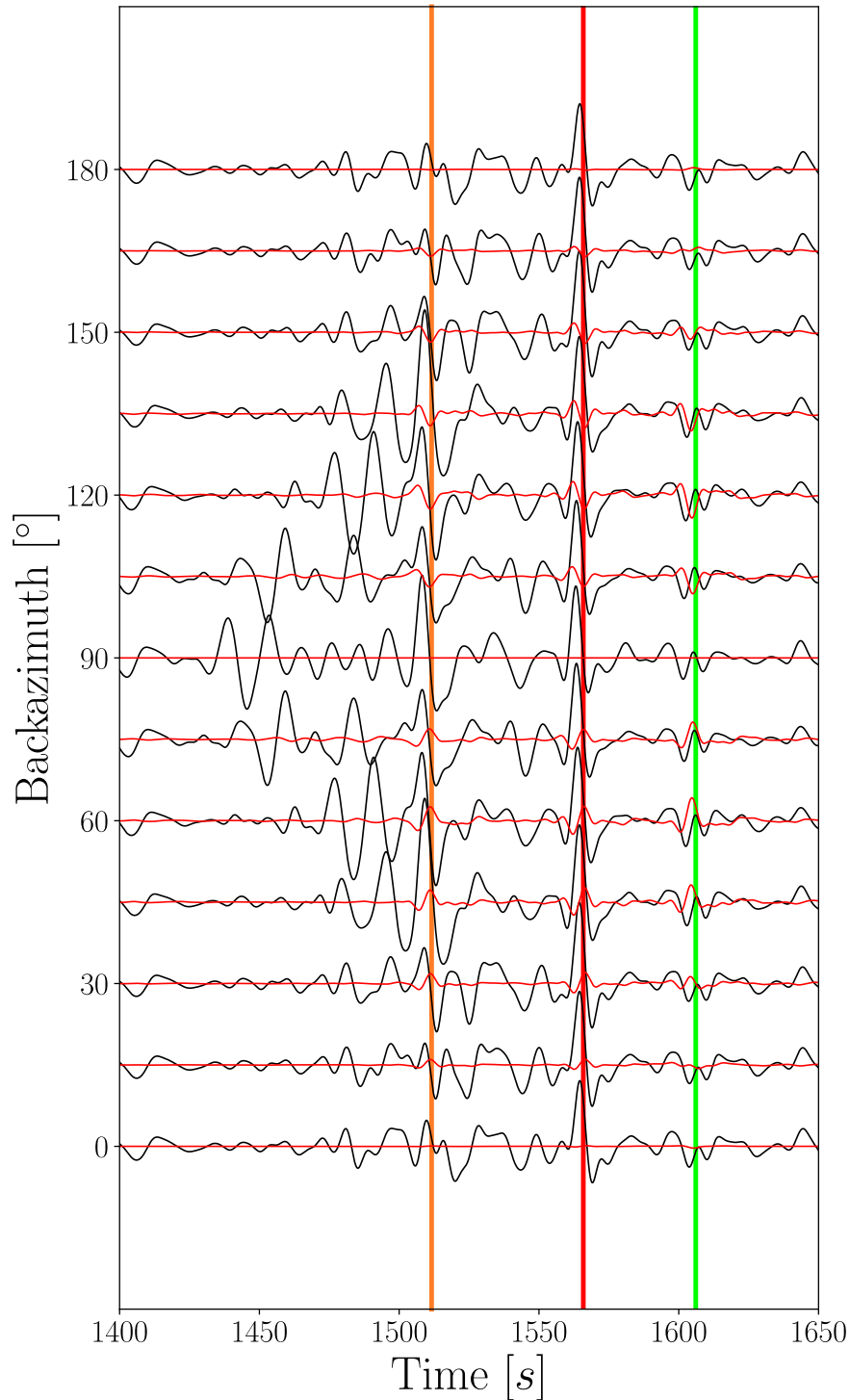
6 *Tesoniero et al., 2019*

Figure S4. Synthetic seismograms computed for the PREM model with olivine style anisotropy localized between 24 and 220km plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 110° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and S_{diff} are shown as the vertical orange, red and green bars respectively. Phases with large amplitude arriving in the 1450 – 1500s range correspond to $PPPP$. For this particular symmetry class they are strongly affected by the anisotropy in the backazimuthal range $45 - 135^\circ$.

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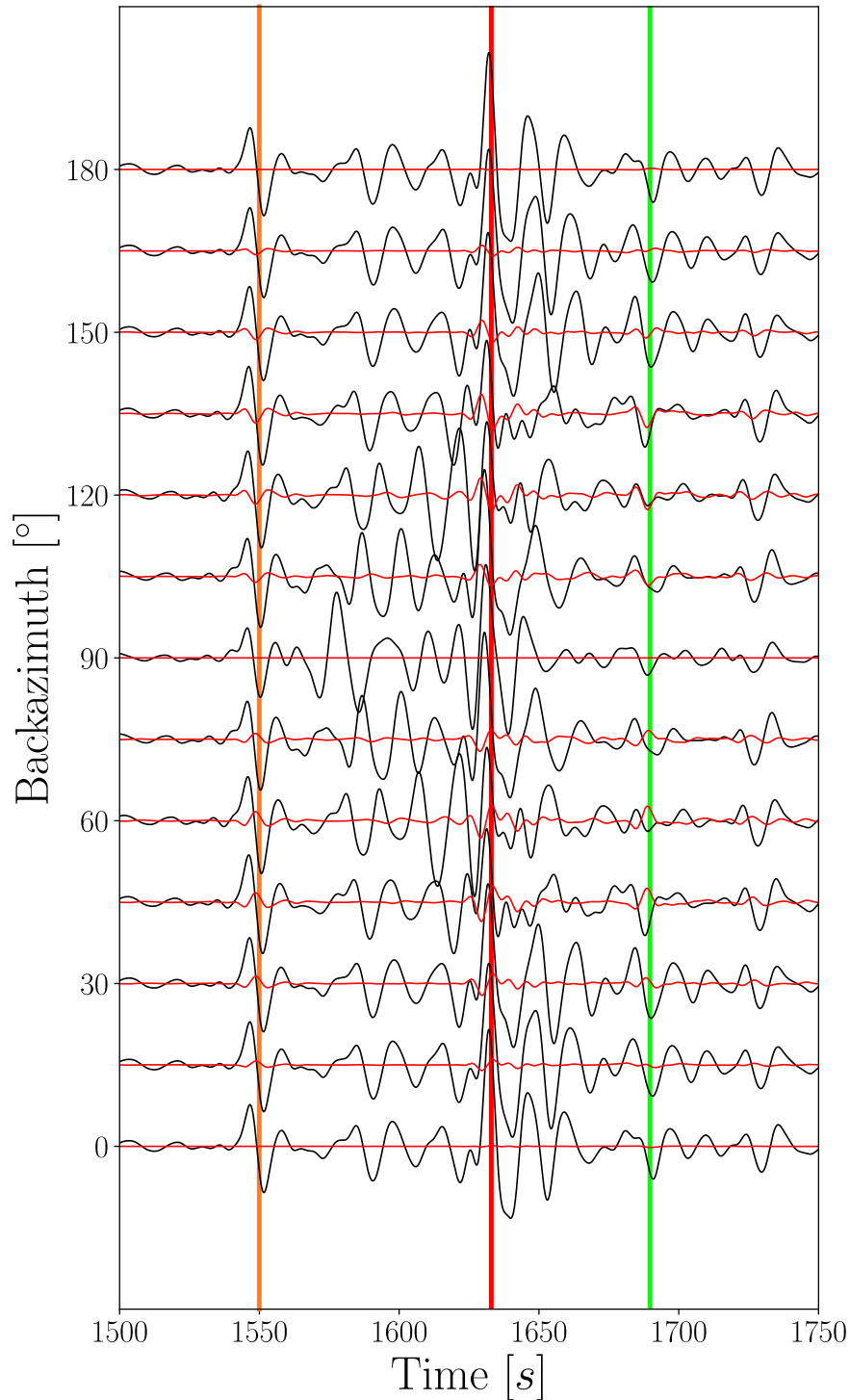


Figure S5. Synthetic seismograms computed for the PREM model with olivine style anisotropy localized between 24 and 220km plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 120° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and S_{diff} are shown as the vertical orange, red and green bars respectively. Phases with large amplitude arriving between SKS and $SKKS$ correspond to $PPPP$. For this particular symmetry class they are strongly affected by the anisotropy in the backazimuthal range $45 - 135^\circ$.

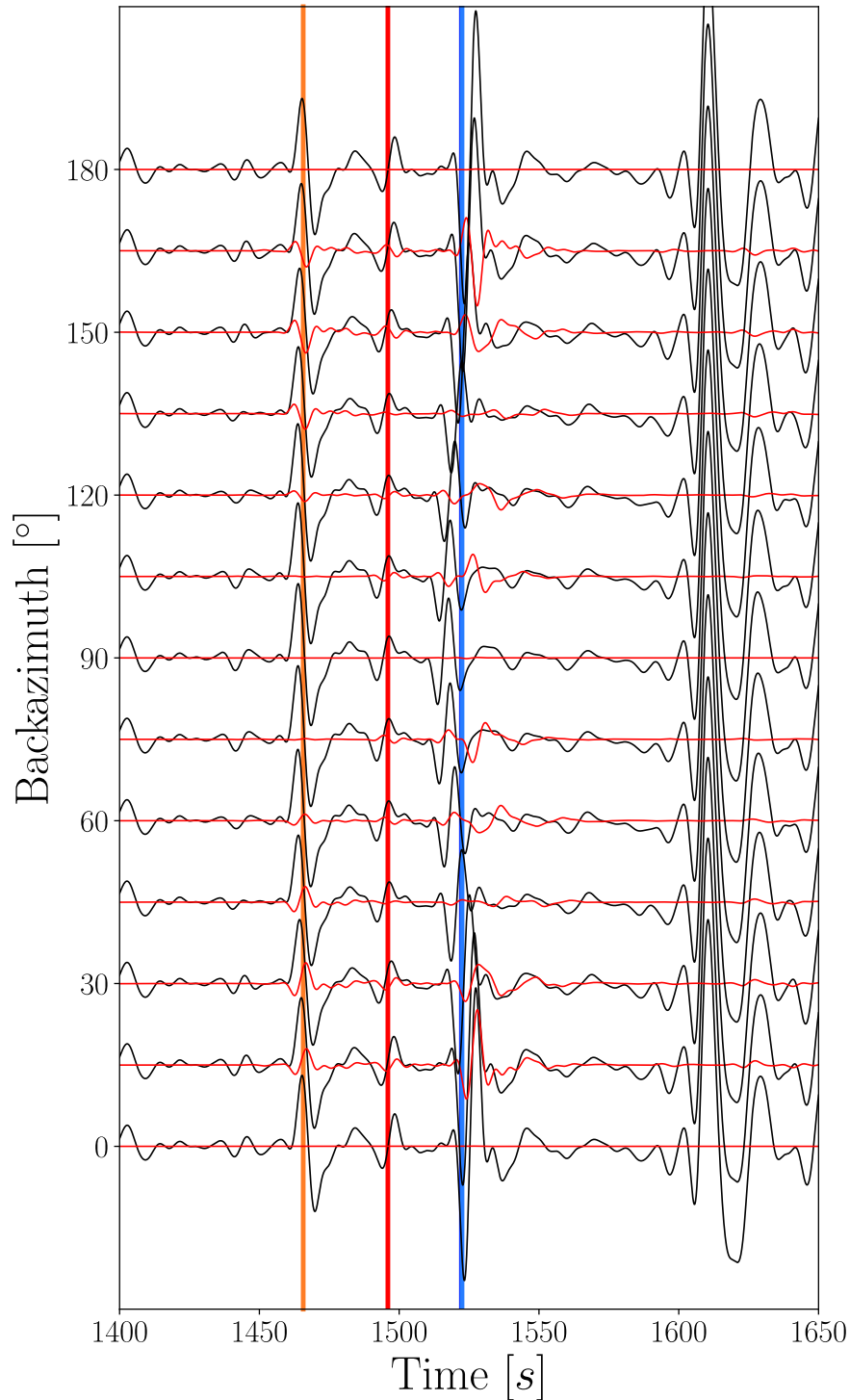
8 *Tesoniero et al., 2019*

Figure S6. Synthetic seismograms computed for the PREM model bridgmanite style anisotropy localized at the base of the base of the lower mantle in a 250km thick D'' plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 100° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and ScS are shown as the vertical orange, red and blue bars respectively.

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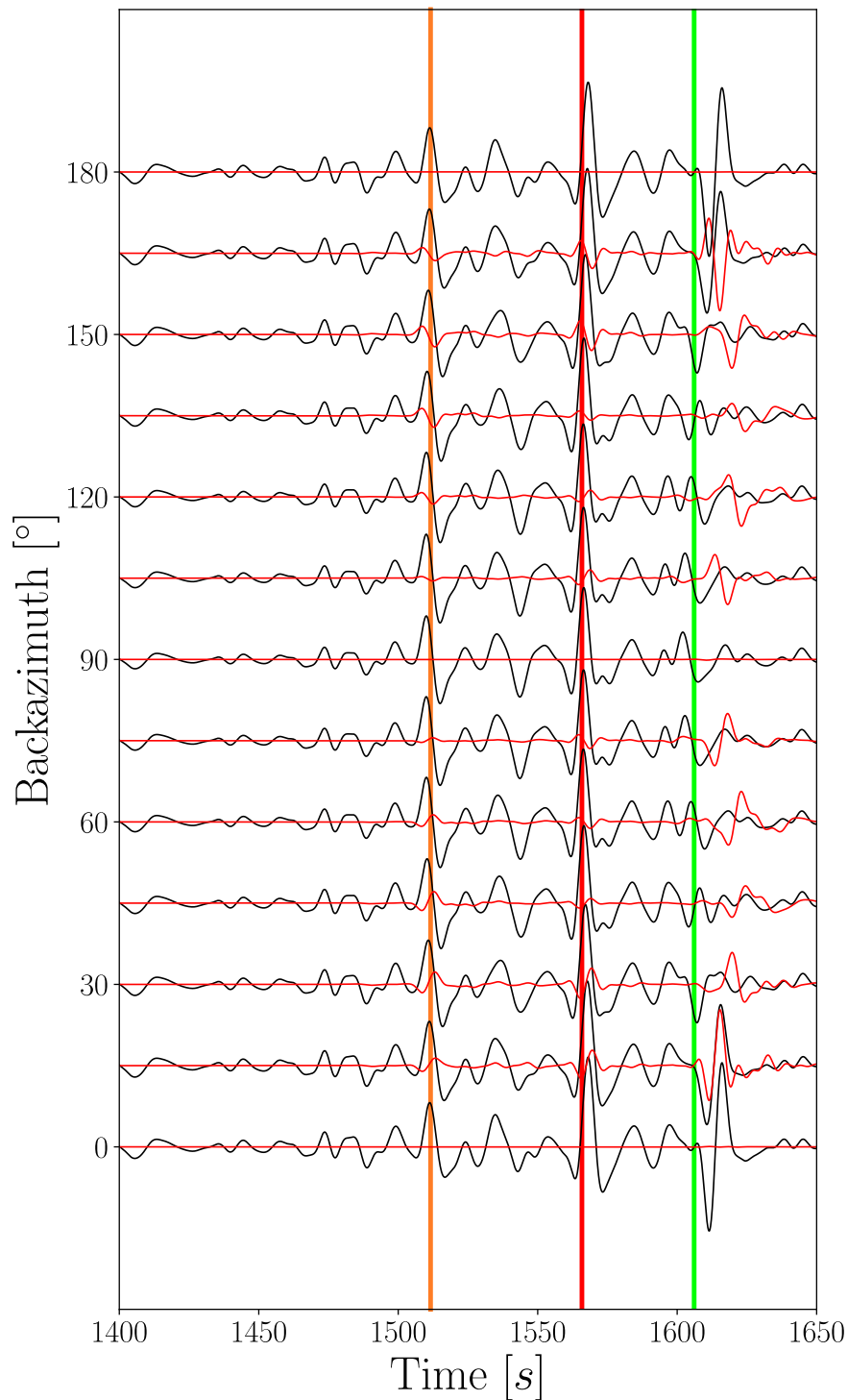


Figure S7. Synthetic seismograms computed for the PREM model a bridgmanite style anisotropy localized at the base of the base of the lower mantle in a $250km$ thick D'' plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 110° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and S_{diff} are shown as the vertical orange, red and green respectively.

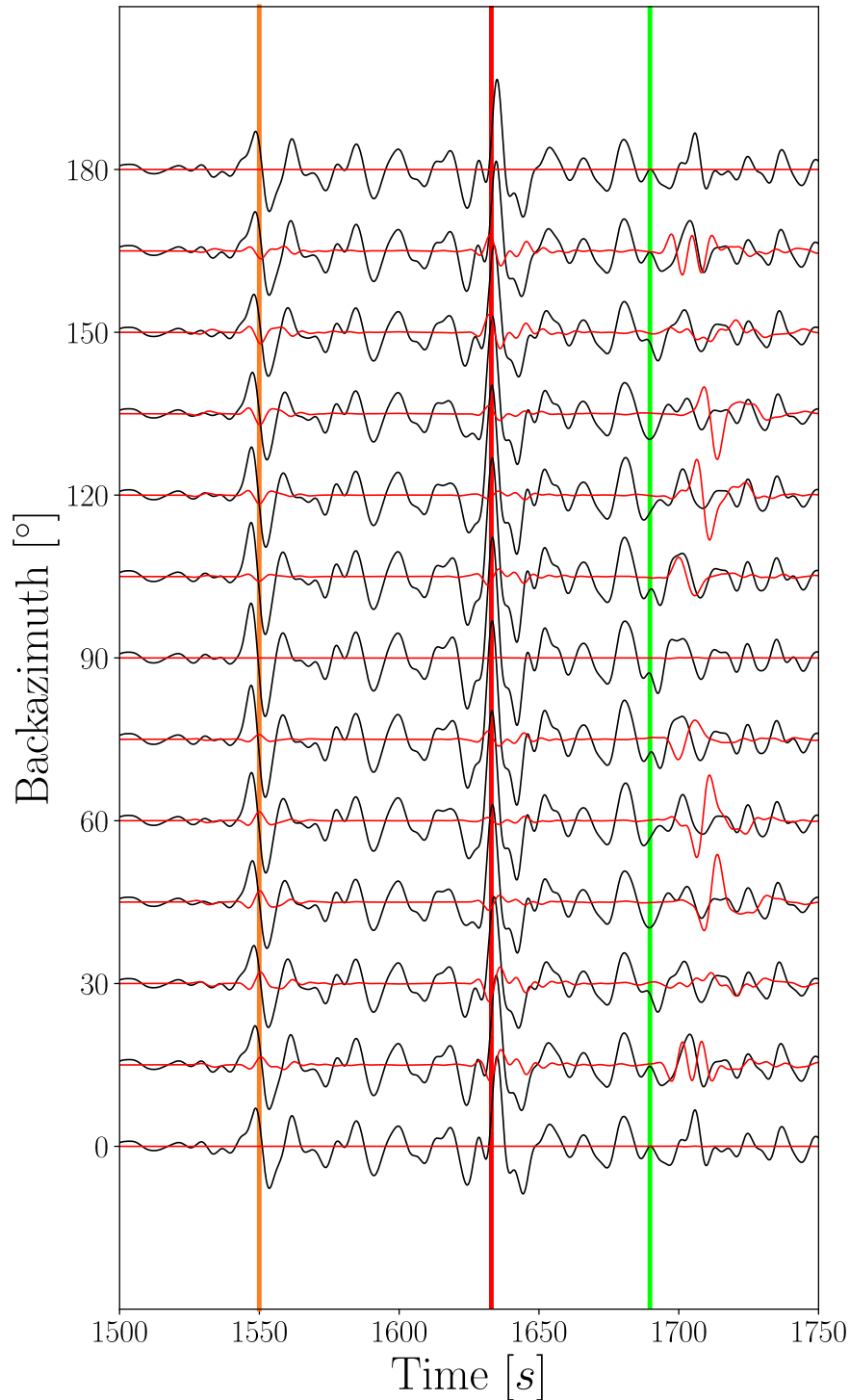
10 *Tesoniero et al., 2019*

Figure S8. Synthetic seismograms computed for the PREM model bridgmanite style anisotropy localized at the base of the base of the lower mantle in a 250km thick D'' plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 120° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and S_{diff} are shown as the vertical orange, red and green respectively.

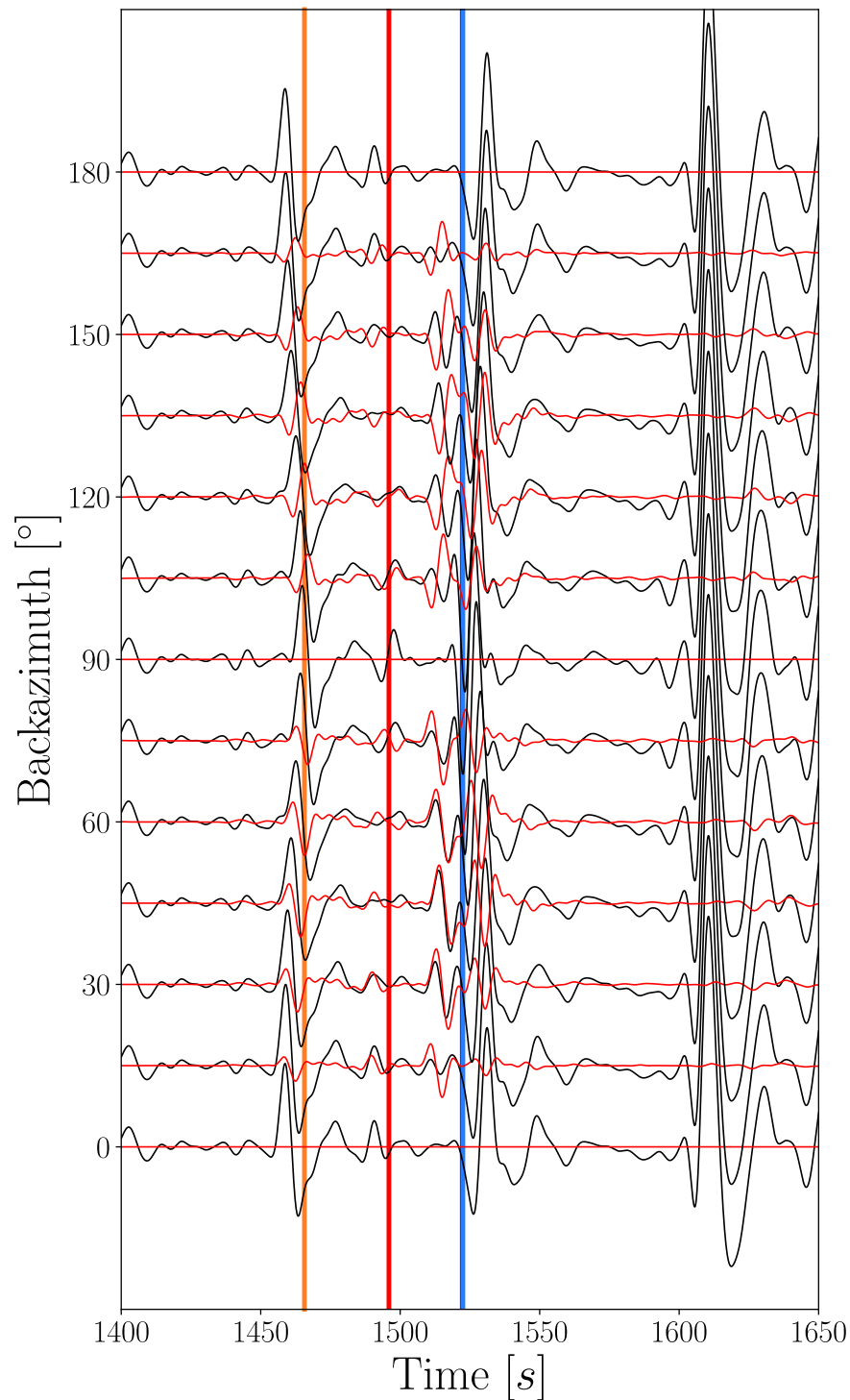


Figure S9. Synthetic seismograms computed for the PREM model with post-perovskite style anisotropy localized at the base of the base of the lower mantle in a 250km thick D'' plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 100° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and ScS are shown as the vertical orange, red and blue respectively.

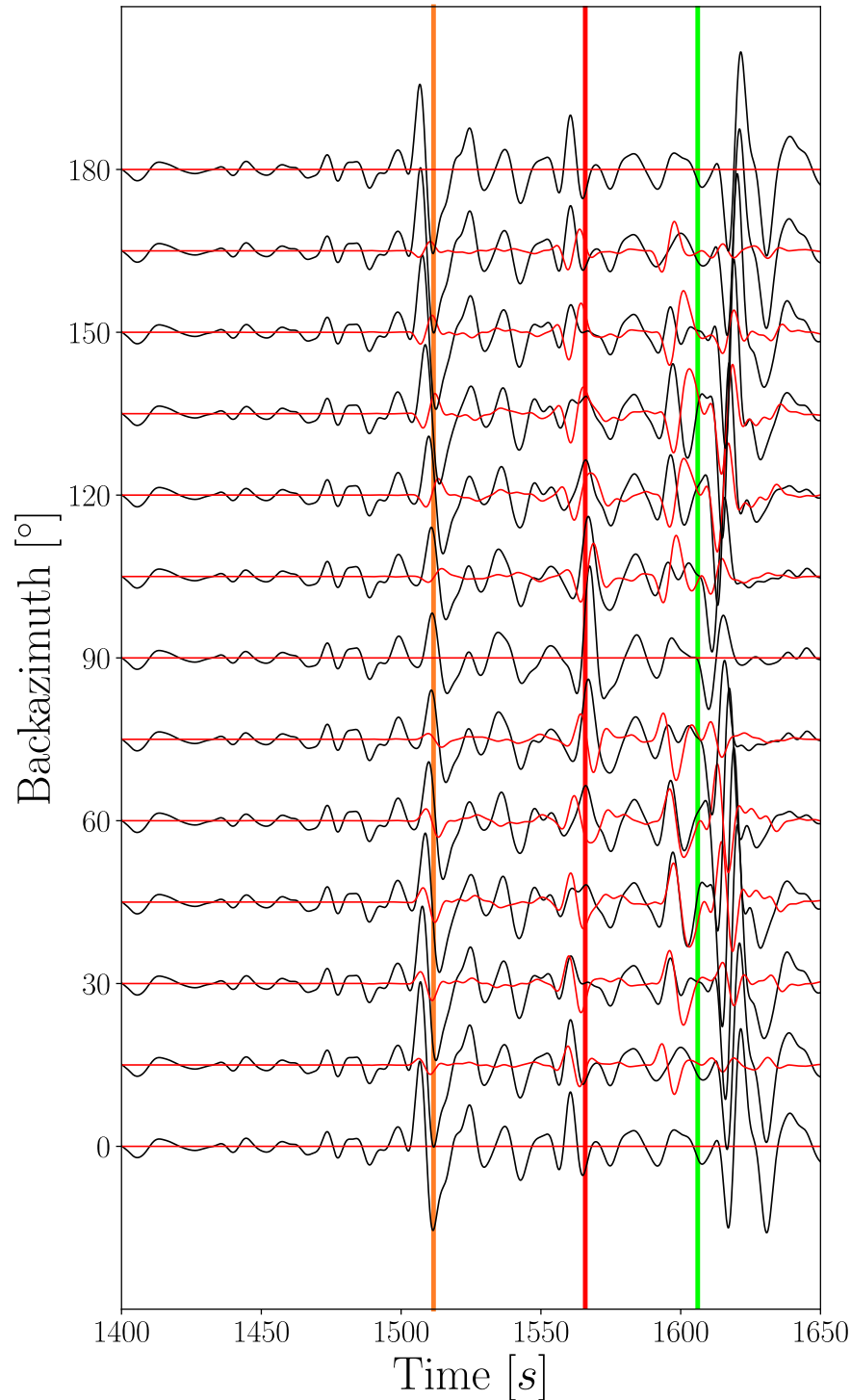
12 *Tesoniero et al., 2019*

Figure S10. Synthetic seismograms computed for the PREM model with post-perovskite style anisotropy localized at the base of the base of the lower mantle in a 250km thick D'' plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 110° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and S_{diff} are shown as the vertical orange, red and green respectively.

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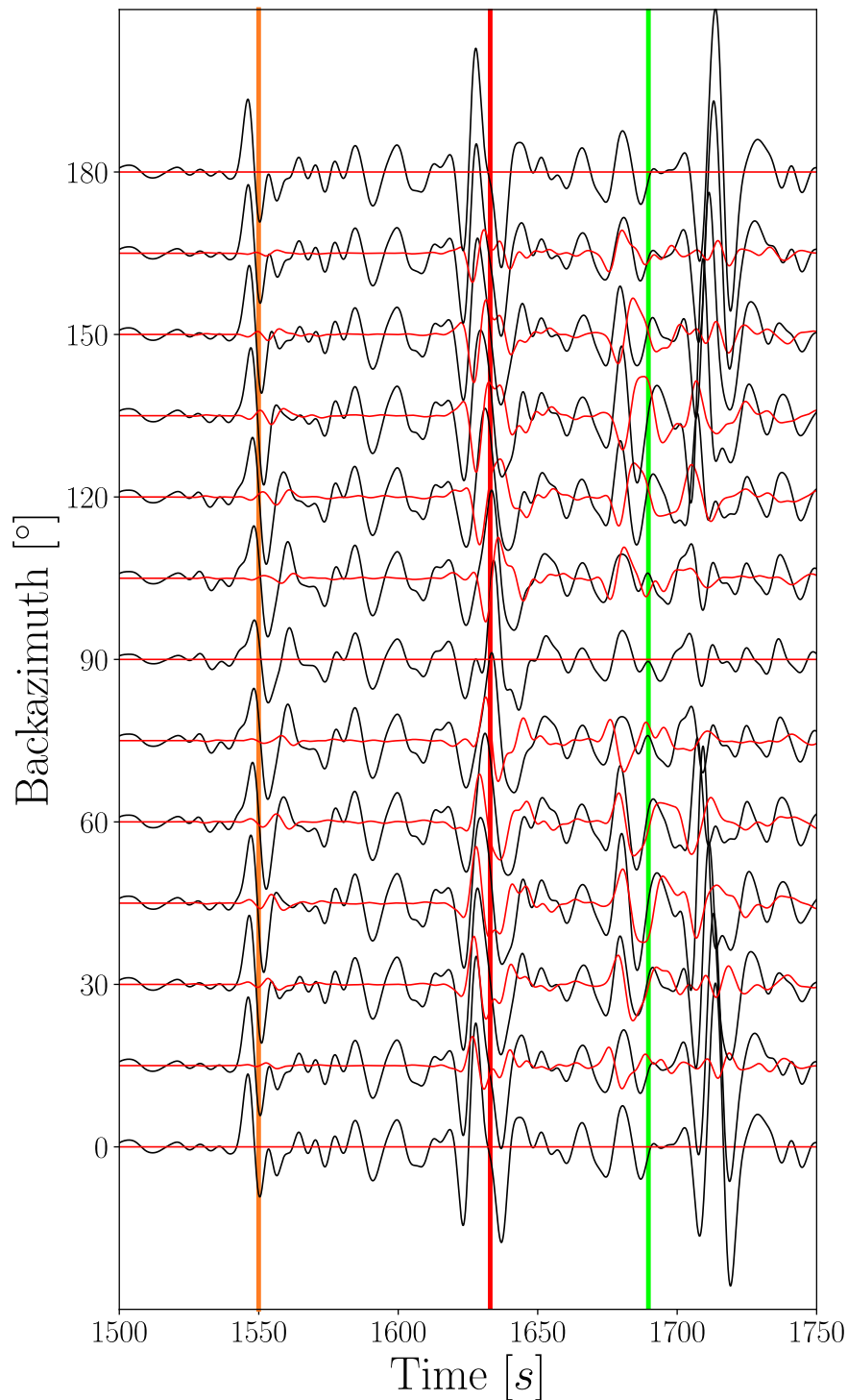


Figure S11. Synthetic seismograms computed for the PREM model with post-perovskite style anisotropy localized at the base of the base of the lower mantle in a 250km thick D'' plotted against the backazimuth of the incoming seismic energy. The epicentral distance is 120° . The radial component is shown in black and the transverse is shown in red. Predicted arrival time for SKS , $SKKS$ and S_{diff} are shown as the vertical orange, red and green respectively.

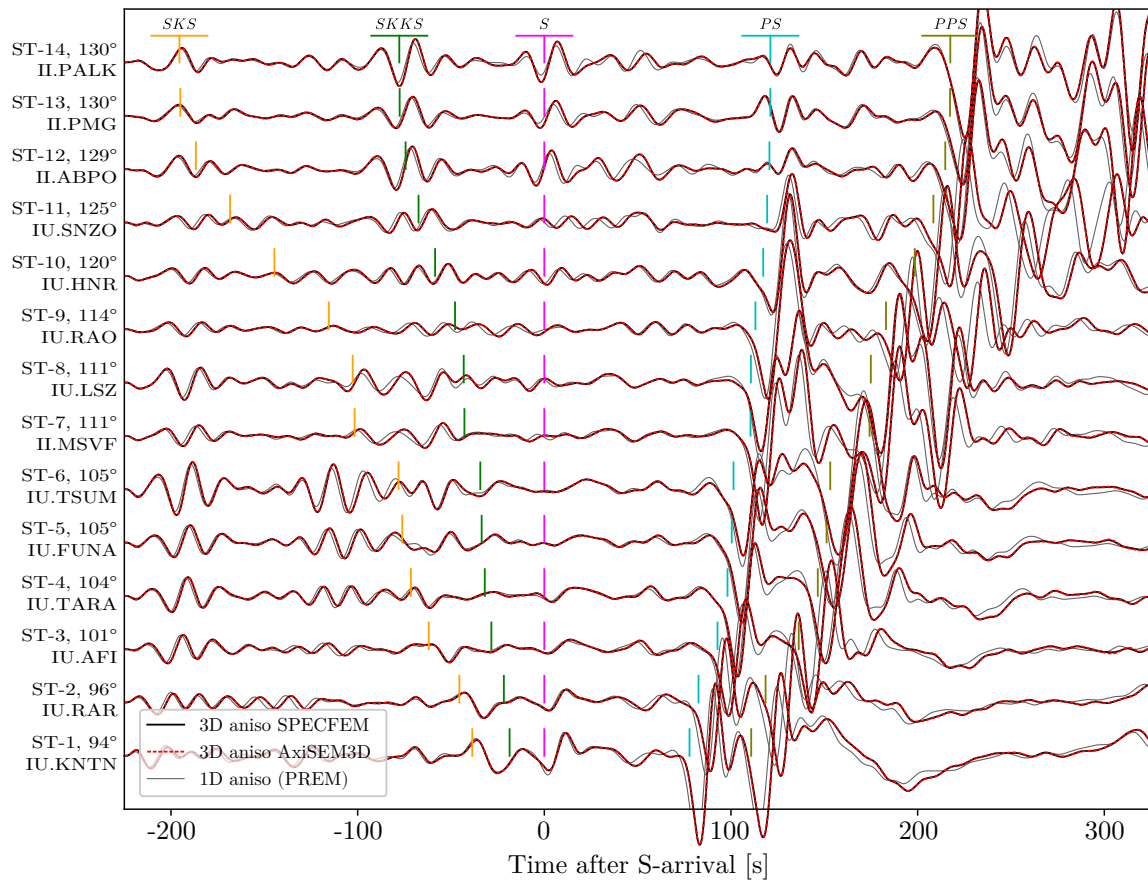
14 *Tesoniero et al., 2019*

Figure S12. Same as Figure 17 of the main manuscript but for the vertical component.

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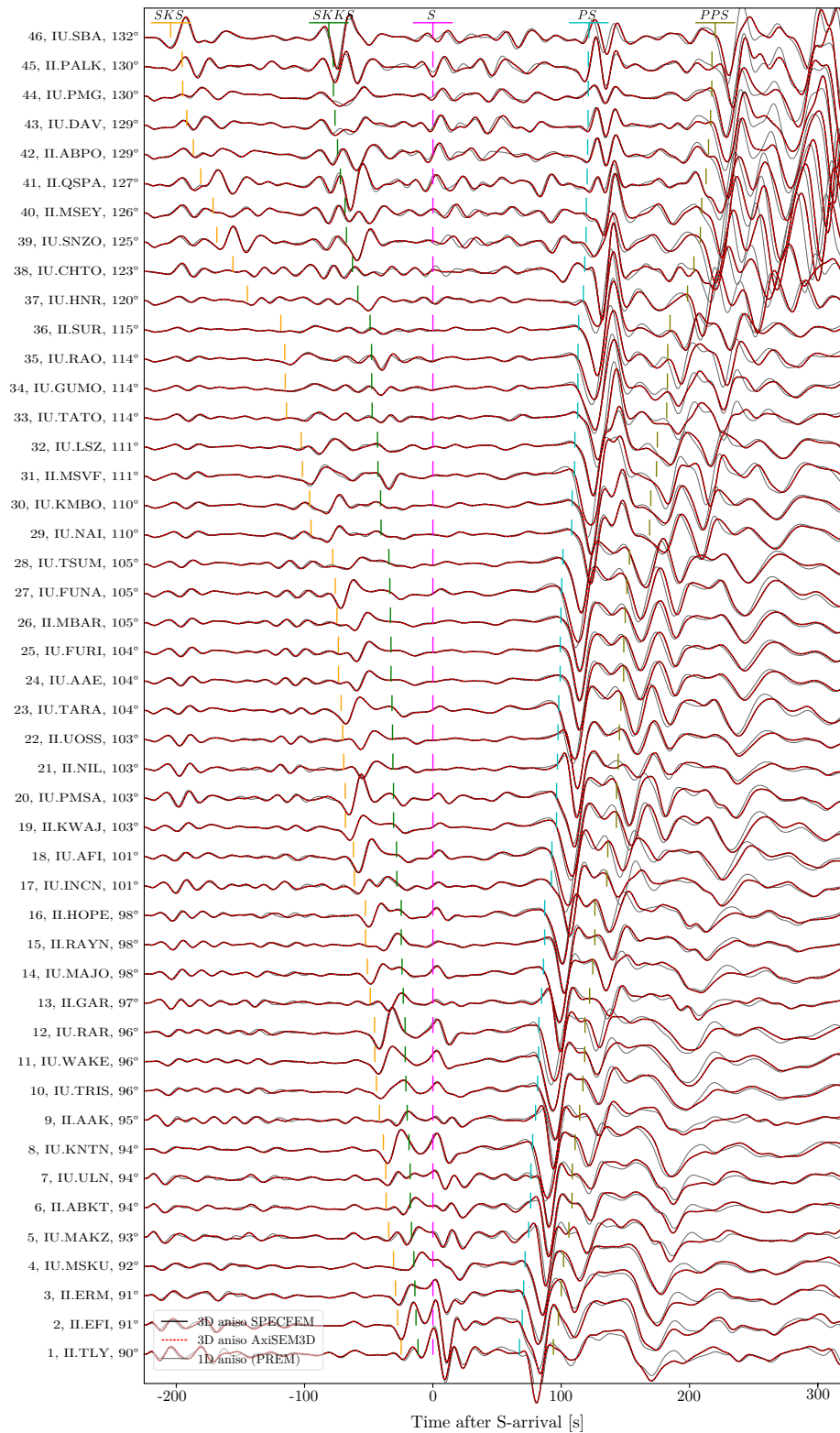


Figure S13. Comparison of synthetic waveforms computed by SPECFEM3D_GLOBE (black) and AxiSEM3D (dashed red) for the benchmark test presented in Section 6 of the main text. The record section shows the radial component for all the stations mapped in Figure 15 of the main text located in the $\sim 90^\circ - 130^\circ$ range of epicentral distance.

16 *Tesoniero et al., 2019*

Figure S14. Same as Figure S13 but for the transverse component.

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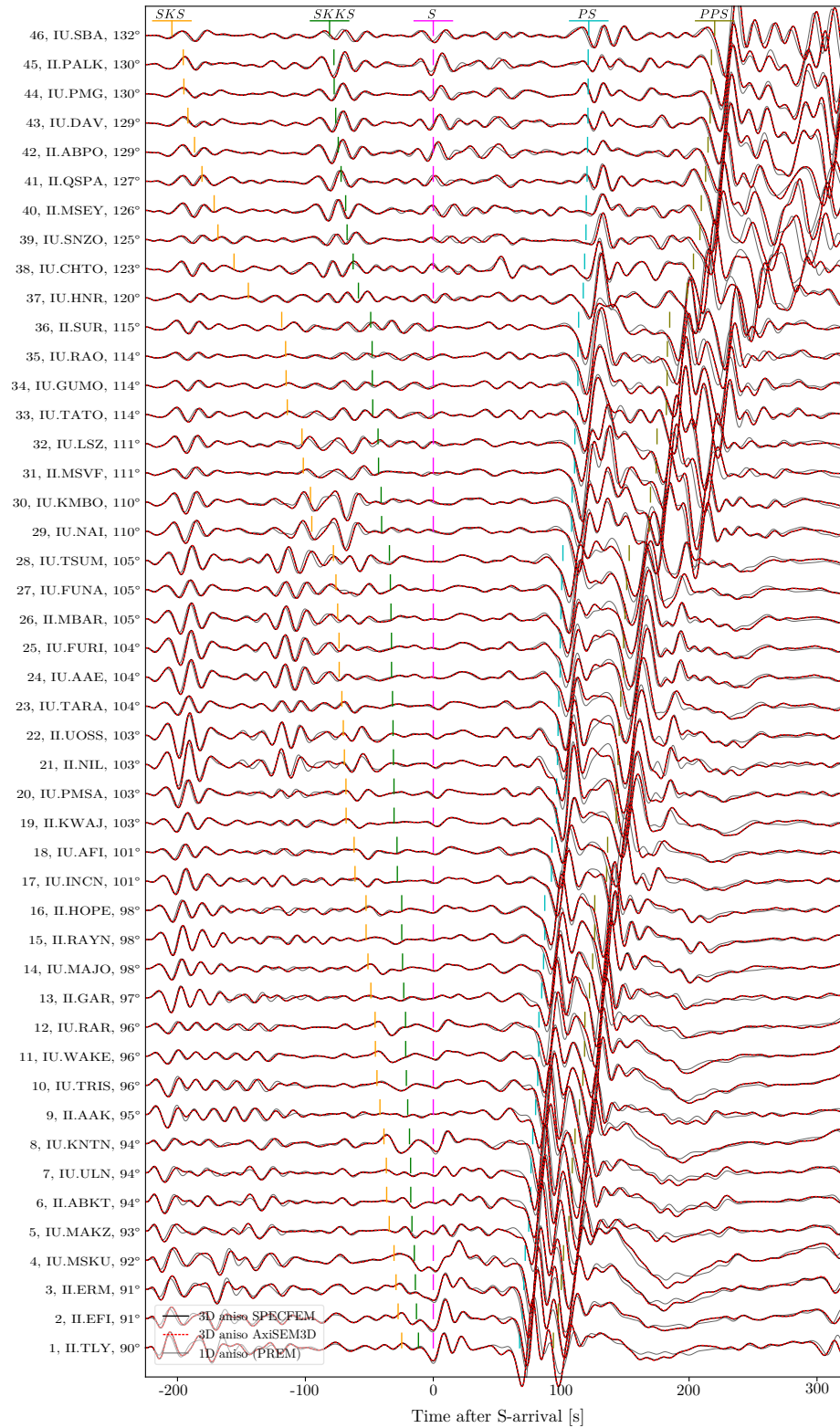


Figure S15. Same as Figure S13 but for the vertical component.