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The Paleocene–Eocene benthic foraminiferal extinction and stable isotope anomalies

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Abstract: In the late Paleocene to early Eocene, deep sea benthic foraminifera suffered their only global extinction of the last 75 million years and diversity decreased worldwide by 30–50% in a few thousand years. At Maud Rise (Weddell Sea, Antarctica; Sites 689 and 690, palaeodepths 1100 m and 1900 m) and Walvis Ridge (Southeastern Atlantic, Sites 525 and 527, palaeodepths 1600 m and 3400 m) post-extinction faunas were low-diversity and high-dominance, but the dominant species differed by geographical location. At Maud Rise, post-extinction faunas were dominated by small, biserial and triserial species, while the large, thick-walled, long-lived deep sea species *Nuttallides truempyi* was absent. At Walvis Ridge, by contrast, they were dominated by long-lived species such as *N. truempyi*, with common to abundant small abyssaminid species. The faunal dominance patterns at the two locations thus suggest different post-extinction seafloor environments: increased flux of organic matter and possibly decreased oxygen levels at Maud Rise, decreased flux at Walvis Ridge. The species-richness remained very low for about 50 000 years, then gradually increased.

The extinction was synchronous with a large, negative, short-term excursion of carbon and oxygen isotopes in planktonic and benthic foraminifera and bulk carbonate. The isotope excursions reached peak negative values in a few thousand years and values returned to pre-excursion levels in about 50 000 years. The carbon isotope excursion was about -2‰ for benthic foraminifera at Walvis Ridge and Maud Rise, and about -4‰ for planktonic foraminifera at Maud Rise. At the latter sites vertical gradients thus decreased, possibly at least partially as a result of upwelling. The oxygen isotope excursion was about -1.5‰ for benthic foraminifera at Walvis Ridge and Maud Rise, -1‰ for planktonic foraminifera at Maud Rise.

The rapid oxygen isotope excursion at a time when polar ice-sheets were absent or insignificant can be explained by an increase in temperature by 4–6°C of high latitude surface waters and deep waters world wide. The deep ocean temperature increase could have been caused by warming of surface waters at high latitudes and continued formation of the deep waters at these locations, or by a switch from dominant formation of deep waters at high latitudes to formation at lower latitudes. Benthic foraminiferal post-extinction biogeographical patterns favour the latter explanation.

The short-term carbon isotope excursion occurred in deep and surface waters, and in soil concretions and mammal teeth in the continental record. It is associated with increased CaCO_3 -dissolution over a wide depth range in the oceans, suggesting that a rapid transfer of isotopically light carbon from lithosphere or biosphere into the ocean–atmosphere system may have been involved. The rapidity of the initiation of the excursion (a few thousand years) and its short duration (50 000 years) suggest that such a transfer was probably not caused by changes in the ratio of organic carbon to carbonate deposition or erosion. Transfer of carbon from the terrestrial biosphere was probably not the cause, because it would require a much larger biosphere destruction than at the end of the Cretaceous, in conflict with the fossil record. It is difficult to explain the large shift by rapid emission into the atmosphere of volcanogenic CO_2 , although huge subaerial plateau basalt eruptions occurred at the time in the northern Atlantic. Probably a complex combination of processes and feedback was involved, including volcanogenic emission of CO_2 , changing circulation patterns, changing productivity in the oceans and possibly on land, and changes in the relative size of the oceanic and atmospheric carbon reservoirs.

During the late Paleocene and early Eocene important changes occurred in global climate, in plate tectonic processes and in the global carbon cycle. The discussion of these events has been complicated and confused by two facts: that there is no unequivocal definition of the 'Paleocene–Eocene boundary' because these words mean different things to different people (e.g. Berggren & Aubry 1996); and that different events occurred at different time scales, from millions of years to thousands of years. We will first discuss longer timescale events, occurring over several millions of years in palaeomagnetic chrons C25 and C24, at about 55–60 Ma in the geomagnetic polarity time scale of Berggren *et al.* (1985), and 52.5–57.5 Ma in that of Cande & Kent (1995). The beginning of this interval has generally been placed in the Paleocene and the end in the Eocene, but the location of the boundary has varied.

This transitional Paleocene–Eocene period witnessed a long-time warming of the deep oceans beginning in the earliest Paleocene, and a long-term decrease of carbon isotopic values in tests of benthic and planktonic foraminifera and bulk carbonate, beginning in early Chron C25 (Shackleton 1986, 1987; Zachos *et al.* 1993). High-latitude areas experienced the highest temperatures of the Cenozoic as indicated by the presence of warm-water pelagic marine organisms (e.g. Haq *et al.* 1977; Premoli-Silva & Boersma 1984; Boersma *et al.* 1987; Stott & Kennett 1990; Aubry 1992; Berggren 1992; Ottens & Nederbragt 1992). Crocodiles, turtles and other thermophilic biota occurred at high northern latitudes (Estes & Hutchison 1980; McKenna 1980; Gingerich 1983; Markwick 1994) in the presence of vegetation and soil-types indicating warm climates (Kemp 1978; Nilson & Kerr 1978; Wolfe 1978; Wolfe & Poore 1982; Schmidt 1991). Palynological data from the North Sea indicate peak warmth (Schroeder 1992). Clay mineral associations in oceanic sediments indicate high humidity and intense chemical weathering as indicated by high abundances of kaolinite (Robert & Maillot 1990; Robert & Chamley 1991; Robert & Kennett 1992, 1994); similar peaks in kaolinite abundance have been observed in sediments from the New Jersey margin (Gibson *et al.* 1993) and the North Sea region (Knox, written comm. 1993). Oxygen isotopic measurements show high temperatures and shallow temperature gradients from low to high latitudes (Shackleton & Boersma 1981; Oberhänsli & Hsü 1986; Stott *et al.* 1990; Barrera & Huber 1991; Seto *et al.* 1991; Stott 1992; Zachos *et al.* 1994; Bralower *et al.* 1995a, b). Dust concentrations in oceanic sediments reached very low levels in the upper part of Chron C24r (Janecek & Rea 1983;

Miller *et al.* 1987b; Rea *et al.* 1990; Hovan & Rea 1992; Rea 1994), suggesting very low wind strength.

During this warm period terrestrial and shallow-water marine organisms did not suffer major net extinctions (e.g. Raup & Sepkoski 1986), but it was a time of major origination of species and high diversity on land (European mammals and flora: Hooker 1991; Collinson 1983; North American mammals: Butler *et al.* 1981, 1987; Wing 1984; Rea *et al.* 1990; Wing *et al.* 1991) and in the surface oceans (e.g. planktonic foraminifera: Kennett 1978; Boersma *et al.* 1987; Boersma & Premoli-Silva 1991; Corfield & Shackleton 1988; Berggren 1992; Corfield 1993; calcareous nannofossils: Romein 1979; Aubry 1992; dinoflagellates, Oberhänsli & Hsü 1986; McGowran 1991).

An explanation of the several million year-long warm period has commonly been sought in elevated levels of atmospheric CO₂, caused by plate tectonic related processes (e.g. Williams 1986; McGowran 1991; Rea *et al.* 1990). During Chrons C24–C25 there was a worldwide plate-tectonic reorganization, involving the slow-down of the northward motion of the Indian subcontinent because of its collision with Asia (Klootwijk *et al.* 1991; Beck *et al.* 1995a). According to some researchers high-temperature metamorphism started in the Himalayas (Tonarini *et al.* 1993; Smith *et al.* 1994; Beck *et al.* 1995a) and delivered large amounts of CO₂ to the atmosphere by decarbonation (Touret 1992; Kerrick & Caldeira 1993, 1994), but this mechanism of CO₂ delivery has been doubted (Selverstone & Gutzler 1993). Increasing levels of CO₂ in the atmosphere resulting from the India–Asia collision could alternatively have been generated by erosion of sediments rich in organic matter (Beck *et al.* 1995b). In addition, subduction changed in direction in the North Pacific (Goldfarb *et al.* 1991). Continental break-up started in the North Atlantic (Roberts *et al.* 1984; Eldholm 1990; Larsen *et al.* 1992), accompanied by violent, partially subaerial plateau basalt eruptive activity (White 1989; White & MacKenzie 1989; Eldholm & Thomas 1993; Kaiho & Saito 1994). High hydrothermal activity along Pacific oceanic ridges may have contributed to increased atmospheric CO₂ levels (Owen & Rea 1985; Olivarez & Owen 1989; Kyte *et al.* 1993), although it is doubted whether increased spreading activity will cause a net increase in atmospheric pCO₂ (Staudigel *et al.* 1989; Varekamp *et al.* 1992). High concentrations of other greenhouse gases such as methane have also been invoked (Sloan *et al.* 1992, 1995).

High atmospheric pCO₂ can not fully explain the global climate, however: how could high latitudes have warmed while the equatorial regions were not

warmer than today (Shackleton & Boersma 1981; Zachos *et al.* 1994; Sloan *et al.* 1995)? The oceans could have transported more heat, involving a larger part of the ocean waters (Barron 1987; Sloan & Barron 1992; Barron & Peterson 1991), but it is not clear that even involvement of the whole ocean in heat transport can succeed in keeping the high latitudes at the temperatures of 15–18°C suggested by oxygen isotope studies (Crowley 1991; Walker & Sloan 1992; Sloan *et al.* 1995). The theory of increased oceanic heat transport appeared attractive because it had been long theorized that deep and intermediate waters in the oceans could have formed as high salinity, high temperature waters by evaporation in subtropical latitudes in the absence of very cold polar areas (Chamberlin 1906; Brass *et al.* 1982; Hay 1989). Surface oceans at low latitudes show the cooling expected if heat transport was more active than today (Zachos *et al.* 1993, 1994; Bralower *et al.* 1995b). The mechanisms for enhanced oceanic heat transport, however, remain undefined (Sloan *et al.* 1995).

Evidence for the presence of such warm saline bottom water (WSBW) is not unequivocal. Some investigators suggest that WSBW was the dominant water mass over much of the Cenozoic (Matthews & Poore 1980; Prentice & Matthews 1988), or during at least the early Paleogene (Kennett & Stott 1990). Some models, however, predict that the deep oceans will turn anoxic over large regions when deep waters form largely at subtropical latitudes (Herbert & Sarmiento 1991), and this did not happen during the Cenozoic (e.g. Thomas 1992). Some investigators concluded that much of the oceans' intermediate and deep waters was formed at high southern latitudes during most of the Maastrichtian and Cenozoic (e.g. Barrera *et al.* 1987; Miller *et al.* 1987a; Katz & Miller 1991; Thomas 1992; Zachos *et al.* 1992, 1993).

Major changes in the deep-water environments of the oceans would be expected to result from a reversal in deep water circulation (Kennett & Stott 1991), and these should be reflected in the composition of deep-sea benthic foraminiferal faunas. Paleocene deep-sea benthic foraminiferal faunas closely resemble Late Cretaceous faunas (Cushman 1946), and the major break in deep-sea benthic foraminiferal faunas occurred somewhere in Chron 24, i.e. between Paleocene and Eocene if seen at coarse time-scales (e.g. Beckmann 1960; von Hillebrandt 1962; Braga *et al.* 1975; Schnitker 1979; Tjalsma & Lohmann 1983; Boersma 1984b; Thomas 1990b; Bolli *et al.* 1994). In Chron 24r (corresponding to planktonic foraminiferal zones P5–P6, and calcareous nannofossil zone NP9) major faunal change occurred in bathyal and abyssal faunas in all the world's oceans (Miller *et al.* 1987b; Boltovskoy & Boltovskoy 1988, 1989;

Berggren & Miller 1989; Thomas 1989, 1990a, b, 1992; Katz & Miller 1991; Mackensen & Berggren 1992; Reynolds 1992, unpublished MSc thesis, Univ. of Maine; Pak & Miller 1992, 1995; Nomura 1991; Kaiho 1988, 1991, 1994a, b; Miller *et al.* 1992; Kaiho *et al.* 1993; Bolli *et al.* 1994.

Benthic foraminifera underwent coeval extinction in neritic to upper–middle bathyal environments in land sections from Israel to Egypt, North Africa and Spain (Molina *et al.* 1992; Speijer 1994), in the North Sea shelf seas (King 1989; Charnock & Jones 1990), in New Jersey (USA, Gibson *et al.* 1993), and along the western Pacific margin from Japan (Kaiho 1988) to New Zealand (Hornibrook *et al.* 1989; Kaiho *et al.* 1993). At these depths the faunal change was associated with low oxygen conditions, as indicated by the presence of black or dark grey, commonly laminated sediments (North Sea, Japan, New Zealand, New Jersey, Spain, Israel, Egypt). Many authors have explained these conditions as resulting from local lack of circulation (e.g. for the North Sea Basin; Charnock & Jones 1990), but such local effects occurred during worldwide low oxygen conditions, which may have been at least partially responsible.

Some authors suggested that relatively low oxygen conditions also occurred, at least locally, in the world's deep oceans (Thomas 1990b, 1992; Kaiho 1991). Most authors implied at least some form of change in deep-water circulation in the benthic foraminiferal extinction, which was seen as caused by changes in temperature as well as nutrient and oxygen content of the deep waters (Miller *et al.* 1987b; Thomas 1989; Katz & Miller 1991; Nomura 1991). Others suggested that oceanic productivity decreased drastically (e.g. Moore *et al.* 1984; Shackleton 1987; Shackleton *et al.* 1985; Corfield & Shackleton 1988; Rea *et al.* 1990; Corfield & Cartlidge 1992a, b; Corfield 1993), with possible effects on the benthic faunas. Yet others suggested that at least in some areas productivity increased (Thomas 1992; Speijer 1994).

Only recently, however, has it become clear that the deep sea benthic extinction in the latest Paleocene was a unique event in its global extent and rapidity (Thomas 1989; Kennett & Stott 1991; Pak & Miller 1992; Kaiho 1994b; Robert & Kennett 1994). In the next paragraphs we will discuss events which happened on timescales of thousands to ten thousands and not millions of years, at some time early in the reversed part of chron C24, in nannofossil zone NP9, and in planktonic foraminiferal zone P5 (Berggren & Aubry 1996; Aubry *et al.* 1996). It had seemed reasonable to suppose that the benthic foraminiferal extinction was not synchronous world wide (Miller *et al.*

1987*b*) because a mechanism for globally synchronous extinction in such a large part of the earth's environment as the deep ocean is difficult to envisage. The observation, however, that the extinction was coeval with a very large, short-term negative excursion in the benthic as well as planktonic $d^{13}C$ and $d^{18}O$ records at ODP Site 690 (Kennett & Stott 1991) suggested that the extinction might have been caused by rapid warming at high latitudes, causing large-scale changes in oceanic circulation and upheaval in the global carbon cycle. Additional research demonstrated that the short-term, extremely negative shift in the carbon isotope values was, as expected by Kennett & Stott (1991), a more than local phenomenon: the event has been recognized in the Bay of Biscay (Pak & Miller 1992; Stott *et al.* 1996), Pacific Ocean (Pak & Miller 1992, 1995; Bralower *et al.* 1995*a, b*), and Indian Ocean (Thomas *et al.* 1992; Lu & Keller 1993). A large negative anomaly in carbon isotopes was also recognized in enamel of land-herbivore teeth and in carbonate concretions, clearly demonstrating that the atmosphere as well as the ocean was involved (Koch *et al.* 1992; Stott *et al.* 1996).

The globally averaged magnitude of the short-term shift is in question: it is about -2‰ in the deep Antarctic (Kennett & Stott 1991; this paper) and the southern Indian Ocean (Lu & Keller 1993), but only about -1‰ in the Pacific and the Bay of Biscay (Pak & Miller 1992). The smaller values may not reflect a global average because the iso-

tope excursion was very short-lived, there may have been hiatuses at Site 577 (Berggren *et al.* 1995), and sampling by Pak & Miller (1992) was not so detailed as to ensure capture of the most extreme values.

The explanation of these short-term changes in climate and in the carbon cycle is being actively debated. Oxygen isotope data suggest rapid warming (over less than 5000 years) of the deep ocean waters at high and low latitudes by $4\text{--}6^{\circ}C$, but essentially no warming of surface waters in the tropical Pacific (Zachos *et al.* 1994; Bralower *et al.* 1995*a, b*), so that latitudinal temperature gradients of surface water were very low for about 50 000 years. Such a rapid, deep ocean-wide temperature change has been explained by a change in dominant oceanic circulation pattern, from dominant production of deep to intermediate waters at high latitudes to dominant production in subtropical regions (e.g. Kennett & Stott 1991; Thomas 1992). Major questions remain as to the exact nature of the upheaval in the carbon cycle, the feedback processes resulting from atmosphere–ocean interactions, and whether deep to intermediate waters did indeed form dominantly at subtropical latitudes, and for how long? In this paper, we present new benthic faunal and isotope data at high resolution from four sites at different depths in the south-eastern Atlantic Ocean and the Weddell Sea (Fig. 1), to investigate the magnitude of the isotopic excursions in different areas and at different water depths. We use the new, detailed data to discuss

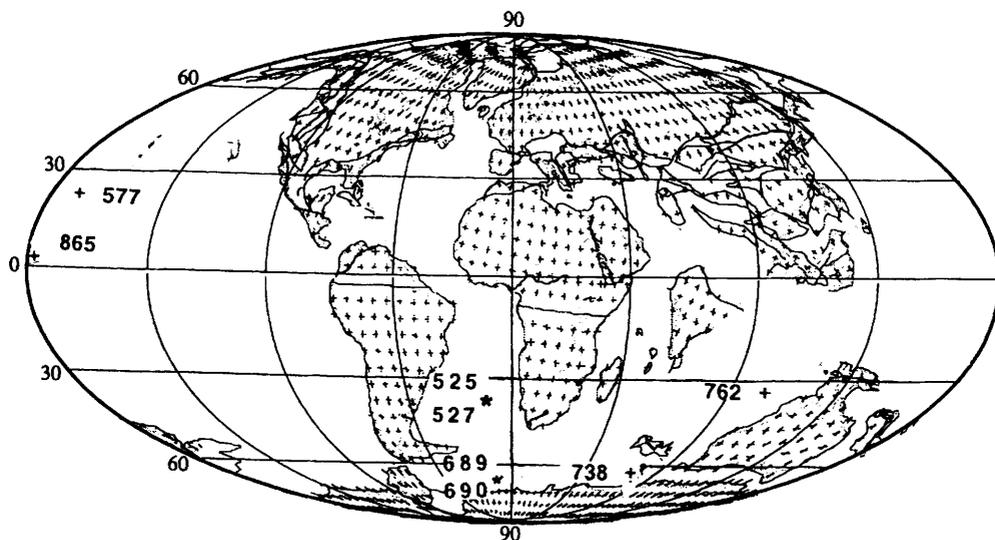


Fig. 1. Palaeogeographic map of the continents in the late Paleocene, after Zachos *et al.* (1994). Sites from which data are presented in this study are indicated by *; other sites mentioned in the text by +.

the bathymetric and biogeographical patterns of the benthic foraminiferal extinction, and re-evaluate the existing database.

Material and methods

Sites and stratigraphy: Maud Rise

Sites 689 (64°31.009'S, 03°05.996'E, present water depth 2080 m) and 690 (65°9.629'S, 1°12.296'E, present water depth 2914 m) were drilled on Maud Rise at the eastern end of the Weddell Sea (Barker *et al.* 1988, fig. 1). Site 689 is on the northeastern side of the ridge near its crest, Site 690 is 116 km to the southwest on its southwestern flank. Depth estimates for the sites at the end of the Paleocene, based on faunal contents, agreed well with backtracking estimates, giving about 1100 m for Site 689, about 1900 m for Site 690 (Thomas 1990, 1992). At both sites sediments of Maastrichtian to Pleistocene age were recovered. Paleogene bio-magnetostratigraphy was reviewed by Thomas *et al.* (1990). Upper Paleocene to lower Eocene sediments consist of chalks and calcareous oozes, with admixture of fine-grained terrigenous matter at Site 690. Recovery was good at Site 690, less so at 689. Core deformation was minimal, but the biostratigraphy was difficult to interpret because the high latitude of the sites caused the absence of many marker species of planktonic foraminifera, and opened the possibility of diachroneity for nannofossil markers (Pospichal & Wise 1990; Aubry *et al.* 1996). Major differences of opinion in the interpretation of the records centred on the lowermost Eocene. Stott & Kennett (1990) maintained that the section was basically complete over the upper Paleocene–lowermost Eocene, whereas Spieß (1990), Thomas *et al.* (1990) and Pospichal & Wise (1990) thought there was at least one, about 2 million year-long hiatus in Chrons 22 and 23. Aubry *et al.* (1996) concluded that an additional hiatus occurs in Chron C24n.

We do not agree with Kennett & Stott (1991) that the records at Site 690 were undisturbed by bioturbation. In our opinion, the sediments show a different ichnofossil assemblage in the interval just after the extinction event, but no lamination. A few corroded specimens of *Gavelinella beccariiiformis* are present two samples above the extinction, and presumably reworked. Therefore we think that the exact sequence of events in samples spaced only by a few centimetres is difficult to determine.

At Site 689 the upper Paleocene–lower Eocene succession was very incomplete, but a section of sediment extending from in the lower part of Core 689B-22X to the poorly recovered Core 689B-24X contains the benthic foraminiferal extinction and the CP7/CP8 nannofossil zonal boundary (Thomas

1990; Pospichal & Wise 1990; Thomas *et al.* 1990). The section is bounded by hiatuses and sedimentation rates can not be determined with precision.

Benthic foraminifera were studied by Thomas (1990a); in this paper we present additional data at higher resolution for Sites 689 and 690 (Appendices 1, 2). For Site 690 our data are from Cores 690B-16 to 690B-22, encompassing sediment deposited in Chron C24r according to Aubry *et al.* (1996). Our data from Hole 689B are from the interval described above (Fig. 2). Carbon and oxygen isotope data of benthic foraminifera were collected by Kennett & Stott (1990, 1991), for planktonic foraminifera by Stott *et al.* (1990) and Corfield & Cartlidge (1992b), and for bulk carbonate by Shackleton & Hall (1990). We present additional data at high resolution for both sites (Appendix 5).

Sites and stratigraphy: Walvis Ridge

Sites 525 (29°04.24'S, 02°59.12'E, present water depth 2467 m) and 527 (28°02'49'S, 01°45.80'E, present water depth 4428 m) were drilled on the summit area and the lowermost western slopes of the Walvis Ridge, respectively (Moore *et al.* 1984, fig. 1). At both sites, Paleocene to Eocene calcareous oozes and chalks were recovered (Shackleton *et al.* 1984a). Palaeodepths for the end of the Paleocene as derived by backtracking (1600 m for Site 525, 3400 m for Site 527; Moore *et al.* 1984) are in agreement with benthic faunal data (Boersma 1984b; this paper). Carbon and oxygen isotope data from foraminifera as well as bulk carbonate were determined (Shackleton *et al.* 1984b; Shackleton & Hall 1984). Nannofossil stratigraphy was described at fairly low resolution by Manivit (1984) and in more detail by Backman (1986a, b). Planktonic and benthic foraminiferal biostratigraphy at low resolution was described by Boersma (1984a, b).

In this paper, we present additional benthic faunal data and isotope data on benthic foraminifera and bulk carbonate over the interval of the benthic extinction. At both sites, despite the difference in palaeodepths, the extinction occurred within the lower few centimetres, but above the base of, an interval of dissolution of 40–50 cm thick, with a sharp lower boundary and a gradual upper boundary (Table 1). In this interval planktonic foraminifera were fragmented, but benthics showed good preservation. We took samples from the cores in which the extinction occurred only.

Numerical ages

Numerical ages for Paleocene sediments are in the process of being revised (e.g. Odin & Luterbacher

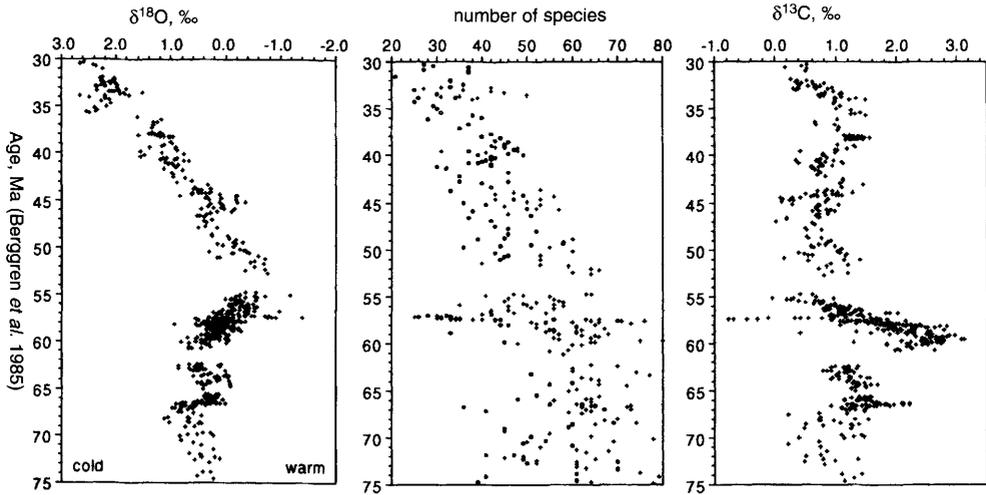


Fig. 2. Benthic foraminiferal carbon and oxygen isotope data (Kennett & Stott 1990; this paper) and species richness of benthic foraminifera (Thomas 1990; this paper) over the Upper Cretaceous through middle Oligocene at Maud Rise, plotted versus numerical age on the Berggren *et al.* (1985) and Aubry *et al.* (1988) timescale. Note the extreme nature of the carbon isotope event as compared to the events at the K/T boundary (66.4 Ma), and its location toward the end of the long-term shift to more negative values, from about 61–58 Ma.

1992) and major changes are proposed from the Berggren *et al.* (1985) and Aubry *et al.* (1988) geomagnetic polarity timescale. The most recent compilations show Paleocene–Eocene boundary ages several million years less than in these two papers (Wing *et al.* 1991; Cande & Kent 1992, 1995; Jenkins & Luterbacher 1992; Odin & Luterbacher 1992; Berggren *et al.* 1995; Berggren & Aubry, 1996; Aubry *et al.* 1996). We are concerned with events during Chron C24r, which has changed little in length between different time scales (around 2.5 million years), although its numerical age has changed from 56.14–58.64 Ma in Berggren *et al.* (1985) to 53.250–55.981 Ma (Cande & Kent 1992) to 53.347–55.904 Ma (Cande & Kent 1994; see also Berggren & Aubry 1996).

Modifications have also been proposed for the planktonic foraminiferal zonation in the Paleocene–Eocene boundary interval (Berggren *et al.* 1995). In the new zonation Zones P5 and P6a of Berggren & Miller (1988) are combined in Zone P5 (the *Morozovella velascoensis* Partial Range Zone; Berggren *et al.* 1995). This revised zonation supports our statement that the benthic foraminiferal extinction was synchronous worldwide. At Sites 525 and 527 the benthic extinction occurred before the last appearance of *Morozovella velascoensis* and before the first appearance of *Morozovella subbotinae* (Boersma 1984a), i.e. in planktonic foraminiferal Zone P5 as modified by Berggren *et al.* (1995). In the planktonic

Table 1. Dissolution intervals and benthic foraminiferal extinction, Walvis Ridge sites

	Depth (mbsf)	
	Hole 525A	Hole 527
Top dark layer	391.87–392.01	200.37–200.56
Benthic extinction	392.41–392.57	200.89–201.07
Bottom dark layer	392.57	201.08

foraminiferal zonation of Berggren & Miller (1988) the extinction occurred in Zone P5 at Sites 525 and 527, but in P6a at equatorial Pacific Site 577 (Miller *et al.* 1987b; Pak & Miller 1992) and in Mediterranean land sections (Speijer 1994).

We agree with Pak & Miller (1992) and Aubry *et al.* (1996) that the benthic extinction occurs lower in palaeomagnetic Chron C24r than estimated by assuming constant sedimentation rates in Chron 24r (Kennett & Stott 1991; Thomas 1992), and that its age was close to 58 Ma in the Berggren *et al.* (1985) time scale (Eldholm & Thomas 1993). In the time scale of Berggren *et al.* (1995) the extinction is close to 55.5 Ma as proposed by Aubry *et al.* (1996).

The isotope excursions and the benthic faunal extinction were very short events (less than 10^5 years; Kennett & Stott 1991; Robert & Kennett 1994; Aubry *et al.* 1996). We cannot assume that sedimentation rates were constant over the interval between age markers where the event occurred at

Sites 525 and 527, because of the presence of a dissolution interval, suggesting decreased sedimentation rates for a short interval starting just before the extinction. Our measurements of CaCO_3 content gave minimum values of 35% at Site 525, 20% at Site 527. At Site 690, where sedimentation rates over the interval between the top of Chron 25 and the nannofossil zone NP9/NP10 boundary were the highest of all sites, CaCO_3 -values also fluctuate, with minimum values of about 65% just below the benthic extinction (O'Connell 1990; fig. 6 in Thomas 1992).

We used the geomagnetic polarity timescale of Berggren *et al.* (1995) and the biostratigraphic interpretation of Site 690 of Aubry *et al.* (1996), giving ages of 55.5 Ma to the benthic extinction event, and 55 Ma to the NP9/NP10 boundary (Berggren & Aubry 1996). We extrapolated age at constant sedimentation rates between 55 and 55.5 Ma, and between 55.5 Ma and the top of Chron 25, using numerical ages as derived from the Berggren *et al.* (1995) timescale for all age marker levels given in Thomas *et al.* (1990) for the Maud Rise sites, in Shackleton *et al.* (1984b) for the Walvis Ridge sites. Then we revised the numerical ages for Site 689 and the Walvis Ridge samples close to the extinction to reflect the supposedly lower sedimentation rates of the intervals with low values of CaCO_3 . To do this we used the carbon isotope record of *Nuttallides truempyi* at Site 690 as standard, because sedimentation rates were highest at this site. For the interval between the boundary between NP9 and NP10 and the benthic foraminiferal extinction we revised ages until the carbon isotope curves at Sites 525, 527 and 690 agreed best in general shape (Fig. 3); we then did the same for the *Lenticulina* spp. isotope record at Site 689. The re-calculation of the ages using the shape of the carbon isotope curve resulted in lower calculated sedimentation rates for the CaCO_3 -poor sediment interval at the sites. All numerical ages are listed in the appendices.

Sample preparation

Samples for benthic foraminiferal faunal and isotope analysis were dried overnight at 50°C and weighed, then soaked overnight in distilled water. Most samples disaggregated readily and could be washed over a 63 µm screen. For isotope analysis of planktonic foraminifera, samples were split and foraminifera picked from two size fractions: size e (212–250 µm) and size d (300–250 µm) as is standard use in the Cambridge and Oxford isotope laboratories; there were no specimens larger than 300 µm in the Maud Rise samples. We analysed 20–40 specimens of *Acarinina mckennai*. In a few

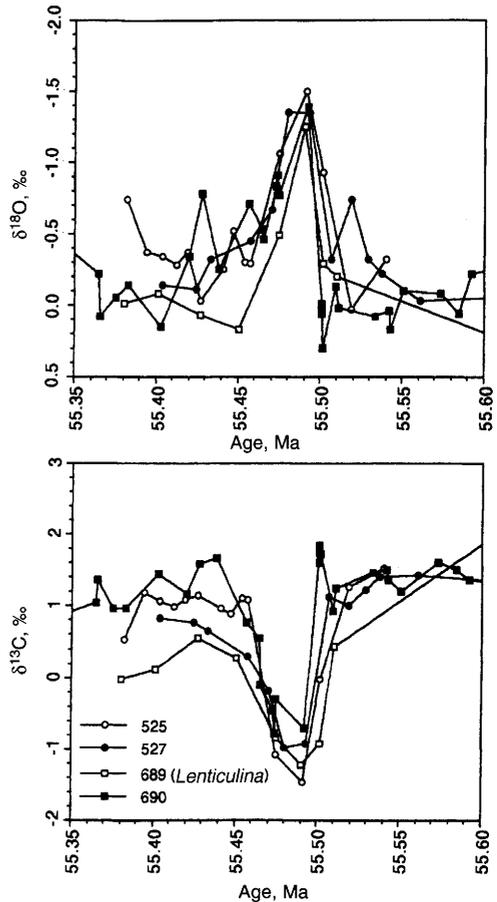


Fig. 3. Benthic foraminiferal carbon and oxygen isotope excursions at Sites 689, 690, 525 and 527; plotted are uncorrected values for *N. truempyi* for Sites 690, 525 and 527, for *Lenticulina* spp. for Site 689, using the geomagnetic polarity timescale of Cande & Kent (1995) and Berggren *et al.* (1996). Ages of Site 690 after Aubry *et al.* (1996); ages for the other sites derived by fitting the carbon isotope curves as described in the text.

samples '*Morozovella*' *convexa* was present in large enough numbers for analysis. Benthic foraminifera for isotope analysis were picked from the > 125 µm size fraction, and we analysed 12–20 specimens. Samples for isotope analysis of the small species *Tappanina selmensis* consisted of 60–80 specimens. *Lenticulina* spp. commonly show large fluctuations in isotope values; we excluded from analysis specimens larger than 500 µm, which showed erroneous values, possibly because they resulted from reworking. All specimens were ultrasonicated to remove adhering

microfossils, then dried at 50°C. Foraminifera were transferred to reaction vessels and roasted at 400°C. Bulk samples, weighing a few milligrams, were taken from samples from Cores 525A-32 and 527-24, dried and vacuum roasted at 400°C to remove organic contaminants. The samples were then reacted with 100% orthophosphoric acid at 90°C using a VG Isotech Isocarb common acid bath system. The evolved carbon dioxide was analyzed in a VG isotech SIRA Series II mass spectrometer. The results were calibrated to PDB by repeated analysis of a carbonate standard. Analytical accuracy is better than 0.08‰ for both $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$.

Benthic foraminifera for faunal analysis were picked from the > 63 μm size fraction, following Thomas (1990a). All specimens counted were picked and mounted in cardboard slides. All samples contained sufficient specimens for analysis (> 250), and counts are shown in the appendices. Taxonomy is as in Thomas (1990a) and largely follows Van Morkhoven *et al.* (1986). Thomas (1990a) misidentified *Neoepionides hillebrandti* as *Neoepionides lunata* and the reverse; that mistake has been rectified in the appendices of this paper. All species richness numbers were recalculated to 100 specimens using rarefaction (Sanders 1968).

Results

Faunas

We define the level of extinction as the interval where species richness declined most rapidly. Pre-extinction faunas at Walvis Ridge and Maud Rise are similar, with almost all species present at all 4 sites (Figs 4, 5, 6, 7), although at varying relative abundances. These faunas contain many cosmopolitan, Late Cretaceous through Paleocene taxa with a large depth range, such as *Gavelinella beccariiiformis*, *Gavelinella hyphalus*, *Neoepionides hillebrandti*, *Neoepionides lunata*, *Pullenia coryelli*, *Bolivinooides delicatulus*, *Neoflabellina semireticulata* and agglutinated taxa such as *Tritaxia paleocenica*, *Tritaxia havanensis* and *Dorothyia oxycona*. The agglutinated taxa are more common at Site 527, the deepest site studied. *Aragonia velascoensis* is rare or absent at Maud Rise, and rare at Site 525 (Fig. 6). *Stilostomella* spp. are more common at Maud Rise than at Walvis Ridge. Pre-extinction faunas have a very high species richness, with many rare uniserial lagenid species and unilocular taxa. Many common, long-lived species occur, such as *Cibicidoides pseudoperlucidus*, *Oridorsalis umbonatus*, *Nonion havanense*, *Nonionella robusta*, *Anomalina spissiformis* and *Anomalinooides semicibrata*. Overall, bi- and tri-serial species were less common at Walvis Ridge

than at Maud Rise. Pre-extinction, large peaks in relative abundance of species such as *Bulimina thanetensis* and *S. brevispinosa* are not observed at Walvis Ridge. *Rectobulimina carpentierae* occurs only at the two deeper sites, 527 and 690, with strongly fluctuating relative abundances at the latter site, and *Bulimina thanetensis* was more common at these sites. *Gavelinella beccariiiformis* was more common at the shallower sites with highest relative abundances at Site 689, in agreement with Katz & Miller (1991) who consider this species typical for bathyal sites at high latitudes. Overall, the pre-extinction faunas were remarkably similar in species composition given the large differences in depth and geographical location, in agreement with Tjalsma & Lohmann (1983) and Kaiho (1988, 1991).

This uniformity ended with the extinction. At Site 690, many of the typical Paleocene species had their last appearance in the highest sample with high diversity, 690B-19H-3, 72–74 cm. In sample 690B-19H-3, 66–68 cm, the diversity decreased from 49 to 26 species per 100 specimens, and bi-triserial species were much more abundant, although in our age model these samples differ in age by only 1000 years. Faunal abundance patterns thus changed at the same level where a high number of last appearances occurred. At Site 689, however, the last appearance of *G. beccariiiformis* and other Paleocene species occurred in sample 689B-23X-1, 80–82 cm, whereas the major drop in species richness from 60 to 42 species per 100 specimens occurred between 23X-1, 80–82 cm and 23X-1, 87–89 cm. Samples 23X-1, 80–82 and 23X-1, 87–89 both contain common *T. selmensis* and *A. aragonensis*, typical species for the post-extinction period, but the species richness is high in these samples and *G. beccariiiformis* is present. These appearances might result from bioturbation or core disturbance. At both Maud Rise sites the extinction occurred after a decrease in carbonate content of the sediments to about 65% (O'Connell 1990; Thomas 1992), and after the first appearance of the keeled, warm-water planktonic species '*Morozovella*' *convexa* at high southern latitudes (Stott & Kennett 1990; this paper).

At Walvis Ridge a sharp drop in benthic foraminiferal species richness from 52 to 34 species per 100 specimens occurred between samples 525A-32-6, 145–147 cm and 525A-32-6, 130–132 cm. At Site 527 species richness decreased from 53 to 24 species per 100 specimens between 527-24-2, 56–58 cm and 527-24-2, 38–40 cm. The level of most rapid decrease in species richness is within the lower part of a dark, low-carbonate layer (Table 1): the benthic extinction thus occurred after the onset of increased CaCO_3 dissolution. The preservation of benthic foraminifera in this layer of

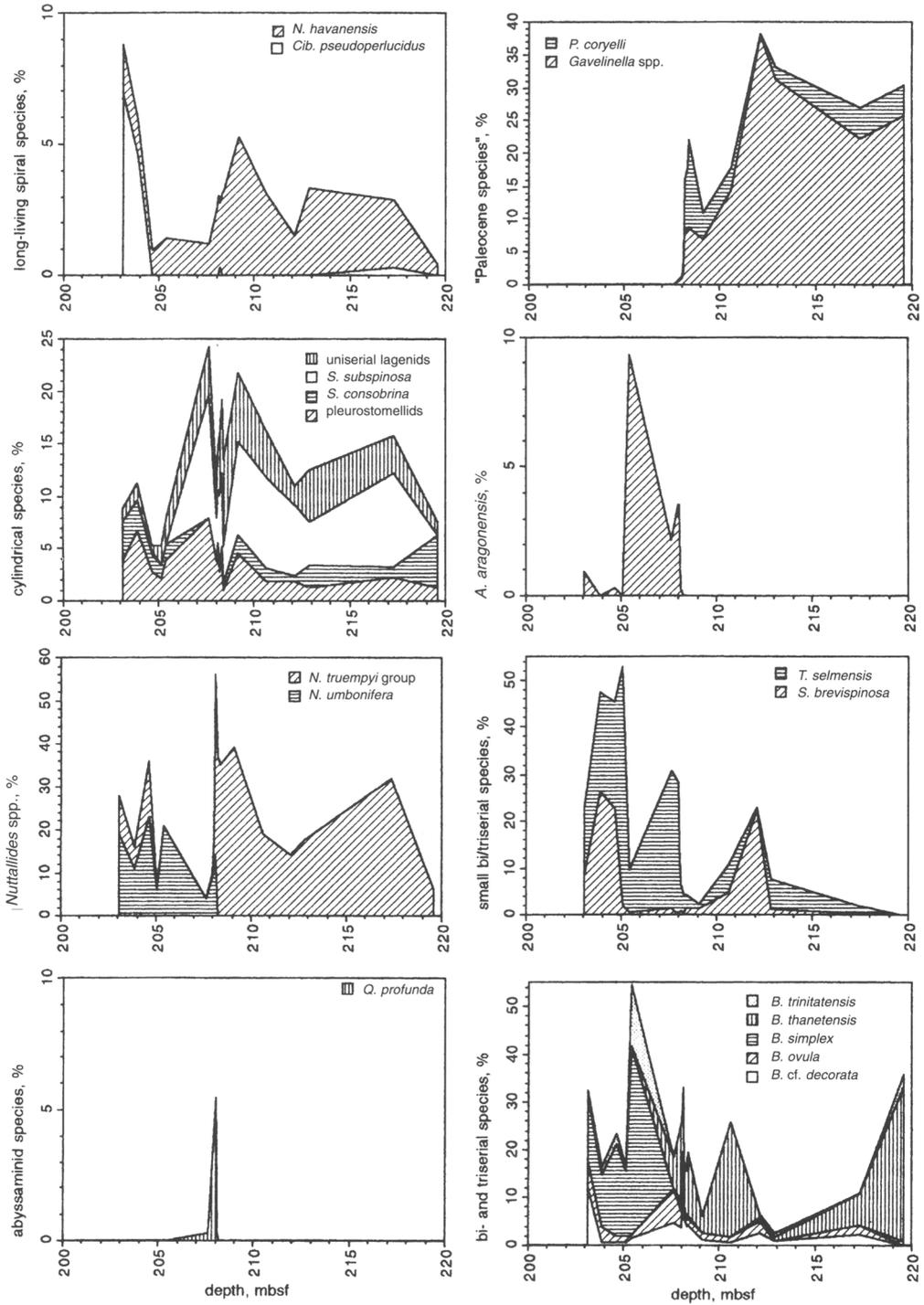


Fig. 4. Benthic foraminiferal relative abundances of most taxa at Site 689; see Appendix 1 for counts.

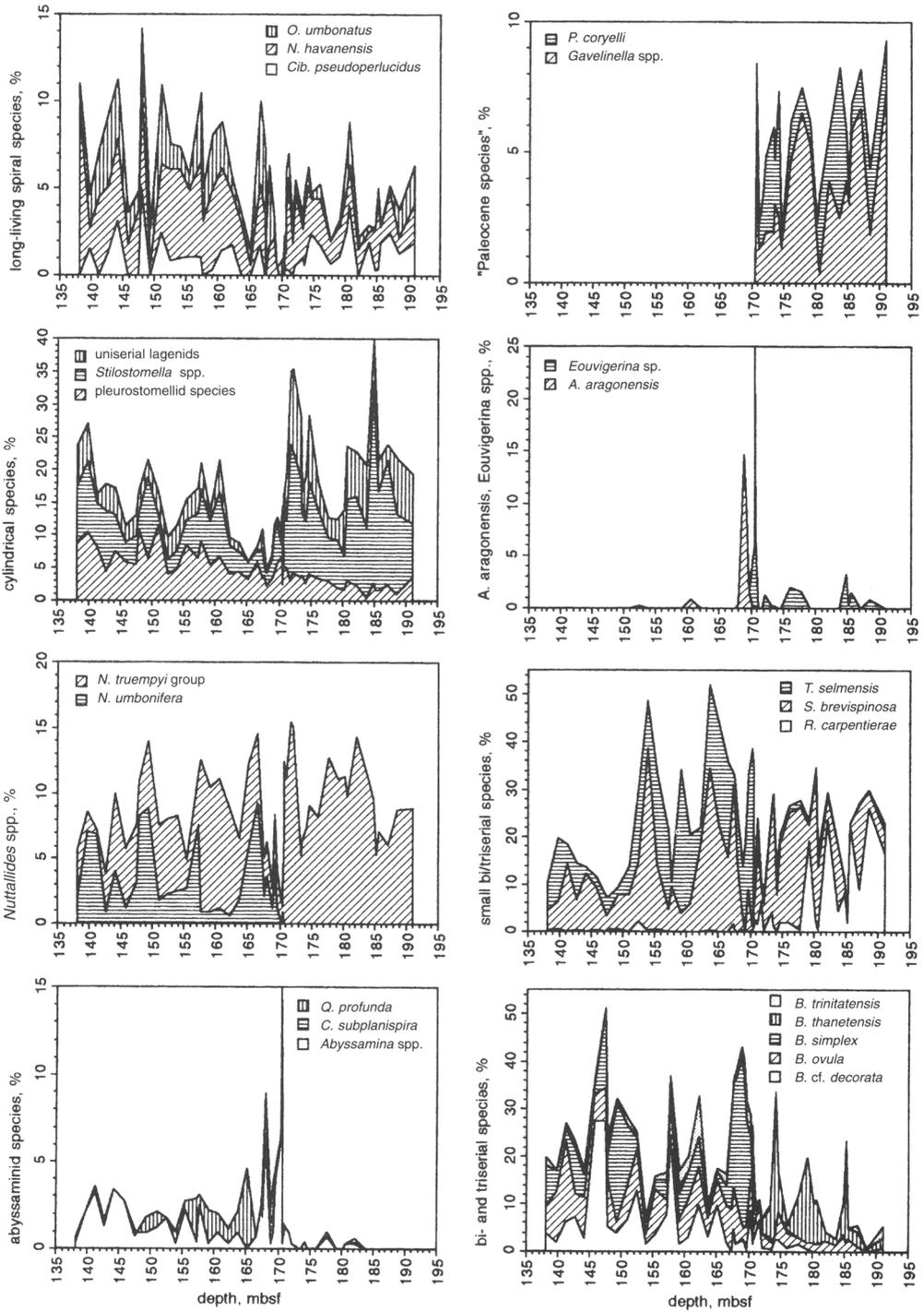


Fig. 5. Benthic foraminiferal relative abundances of the taxa, Site 690; see Appendix for counts.

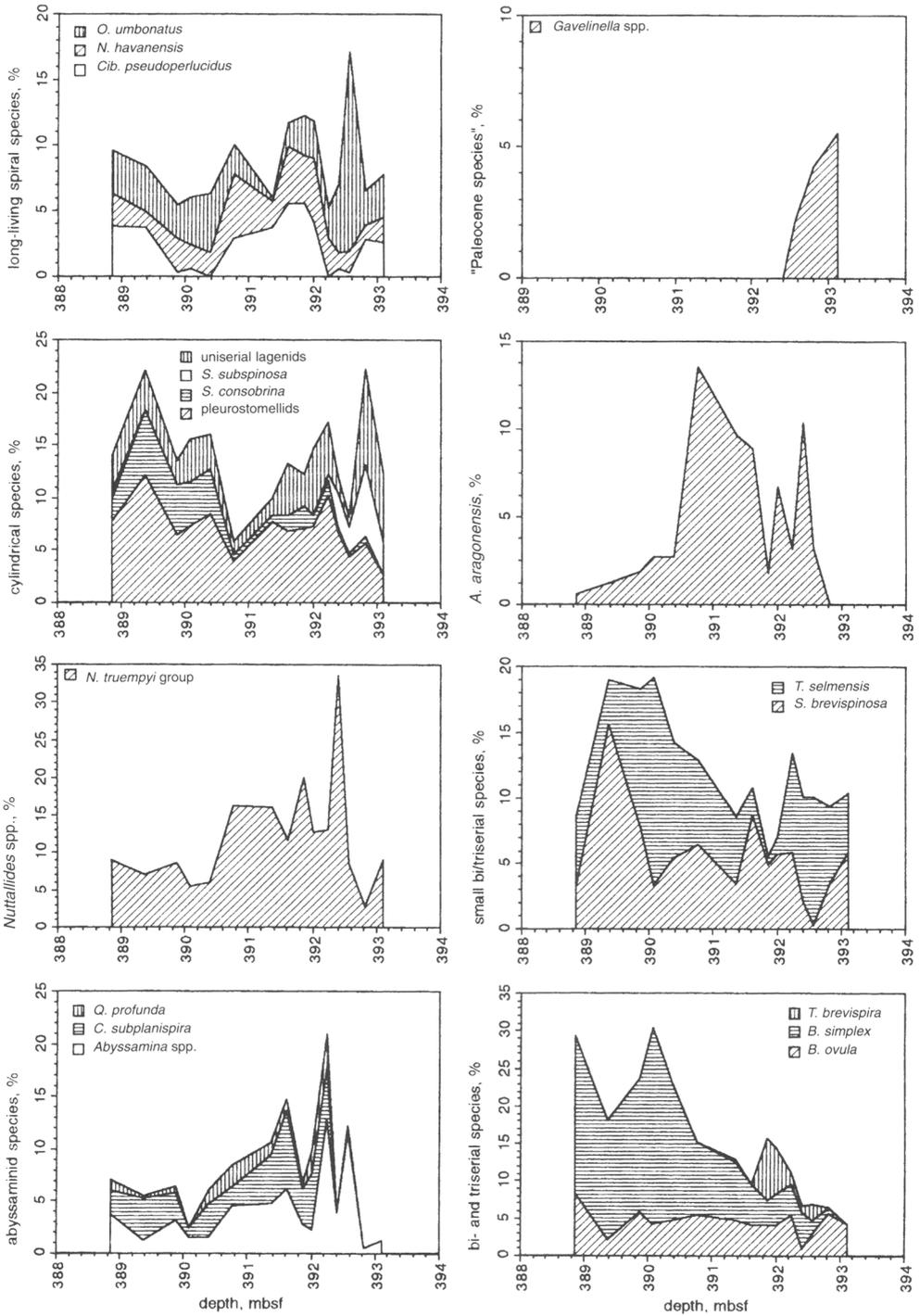


Fig. 6. Benthic foraminiferal relative abundances of the most common taxa at Site 525; see Appendix 2 for counts.

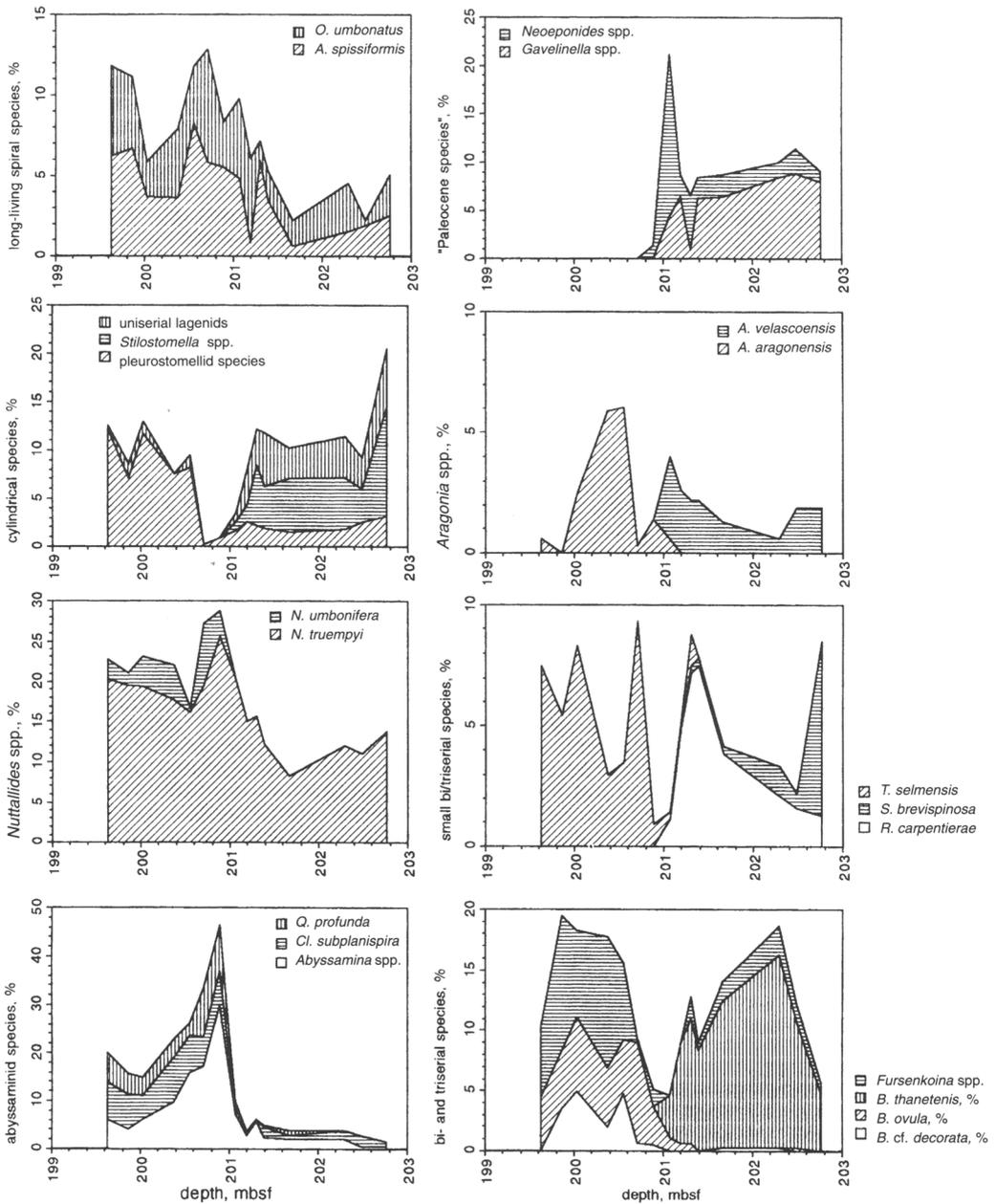


Fig. 7. Benthic foraminiferal relative abundances of the most taxa, Site 527; see Appendix 4 for counts.

low CaCO_3 -content and increased fragmentation of planktonic foraminifera is good.

The samples in the interval with low diversity in the first 50 000 years after the extinction differ in faunal composition at the sites, with most differences between the Walvis Ridge and the Maud Rise

faunas, lesser differences between the sites at different depths in each area. These faunas occur during the interval with very low $\delta^{13}\text{C}$ isotopic values at all 4 sites (see below), and thus can be considered coeval. At Site 690, the post-extinction faunas had high relative abundances of *Eouvigerina*

sp., and at both Maud Rise sites *Tappanina selmensis*, *Bolivinoidea* cf. *decorata*, *Bulimina ovula*, *Bulimina simplex* and *Bulimina trinitatensis* had high, but strongly fluctuating relative abundances. *N. truempyi* was rare or absent just after the extinction. *Nuttallides umbonifera* first occurred just after the extinction at Maud Rise, but is not present in the samples studied from Walvis Ridge.

At the Maud Rise sites abyssaminid species (small, thin-walled, spiral species; Schnitker 1979; Tjalsma & Lohmann 1983) were more common after the extinction. At Walvis Ridge there was a much stronger increase in the relative abundance of these species, with the highest at the deeper Site 527, in agreement with Tjalsma & Lohmann (1983) and Katz & Miller (1991). In contrast to the patterns at Maud Rise, the post-extinction faunas at Walvis Ridge had very high relative abundances of *Nuttallides truempyi*, as described by Tjalsma & Lohmann (1983) and Katz & Miller (1991) for other Atlantic sites. A distinct drop in relative abundance of biserial and triserial species occurred at Site 527. Slightly higher in the section, however, the biserial species *Tappanina selmensis* and *Aragonia aragonensis* increased in relative abundance, with highest abundances at the shallower site. *Tappanina selmensis* never reached such a high abundance as at the Maud Rise sites (Figs 4, 5, 6, 7). Many rare species (such as lenticulinids, uniserial lagenids, unilocular taxa) were absent at all sites in the post-extinction faunas, at least partially causing the low species richness. Many of these species, however, reappeared higher in the section, and their absence may thus be apparent because of the decreased evenness of the faunas (Signor & Lipps 1982). At counts of about 300 specimens a lower number of species would be observed in an assemblage with higher dominance (Gage & Tyler 1991).

Comparison of the faunas (Fig. 8) showed that at all sites the observed species richness dropped precipitously, with lowest values reached at the deepest Maud Rise site. The species richness appears to recover more quickly at the Walvis Ridge sites, especially at Site 525, if ages are determined according to Fig. 3 (fitting of the $\delta^{13}\text{C}$ curve of *N. truempyi*). Interesting is the strong increase in relative abundance of *N. truempyi* at the Walvis Ridge Sites, coeval with its strong decline at both Maud Rise sites. Abyssaminid species increased in relative abundance after the extinction at all sites, but much more so at the Walvis Ridge sites, and the shallower Walvis Ridge site has higher relative abundances of these species than the deeper Maud Rise site. Biserial and triserial species increased in relative abundance just after the extinction at both Maud Rise sites, but decreased

at the deeper Walvis Ridge site and showed little change at Site 525.

A typical component of post-extinction faunas in all oceans are the species *T. selmensis* and *A. aragonensis* (Boersma 1984b), although their relative abundances vary from site to site. *T. selmensis* first appeared in the Late Cretaceous, but occurred at low abundances until the extinction event. *A. aragonensis* had its first appearance at the time of the extinction (Tjalsma & Lohmann 1983; Van Morkhoven *et al.* 1986; Bolli *et al.* 1994). *G. beccariiiformis* last appeared at all sites at about the same time, although the species started a decline in relative abundance at Site 689 several hundred thousand years earlier (Fig. 4). Overall, however, this pattern does not agree with the statement by Tjalsma & Lohmann (1983) that the *G. beccariiiformis* faunas became gradually restricted to shallower depths before becoming extinct. *B. thanetensis* became extinct in the late Paleocene at all sites, but always a few samples higher than *G. beccariiiformis*.

Benthic foraminiferal isotopes

We measured stable isotopes in several species of deep sea benthic foraminifera from the Maud Rise sites (Figs 9, 10). We wanted to derive an isotopic record of the extinction event that is as complete as possible, and *N. truempyi* and *Cibicidoides* spp. are not present in large enough numbers for isotope analysis just after the extinction. We also wanted to determine whether species that are presumed to be infaunal or epifaunal because of their morphology (Corliss & Chen 1988; Rosoff & Corliss 1992; Thomas 1990) show a different carbon isotopic signature, as observed in Recent and Neogene taxa (Woodruff & Savin 1985; Zahn *et al.* 1986; Altenbach & Sarnthein 1989; McCorkle *et al.* 1990). Infaunal species precipitate their tests in contact with pore waters, not sea water, and can thus be expected to have lower $\delta^{13}\text{C}$ values.

Most species thought to be infaunal because of their morphology had indeed lower $\delta^{13}\text{C}$ values than species thought to be epifaunal, with the exception of *Oridorsalis umbonatus*. This long-lived, cosmopolitan, trochospiral species plots in the field of infaunal species (Fig. 10). Its carbon isotope data should thus be interpreted carefully, especially because Rathburn & Corliss (1994) reported that Recent representatives of this taxon live infaunally.

The benthic species show considerable scatter in their isotopic values, and there are differences between the patterns at Sites 689 and 690. At Site 690, *Lenticulina* spp., a presumed infaunal group, shows clear separation in $\delta^{13}\text{C}$ values from the epifaunal species, and has consistently lower values

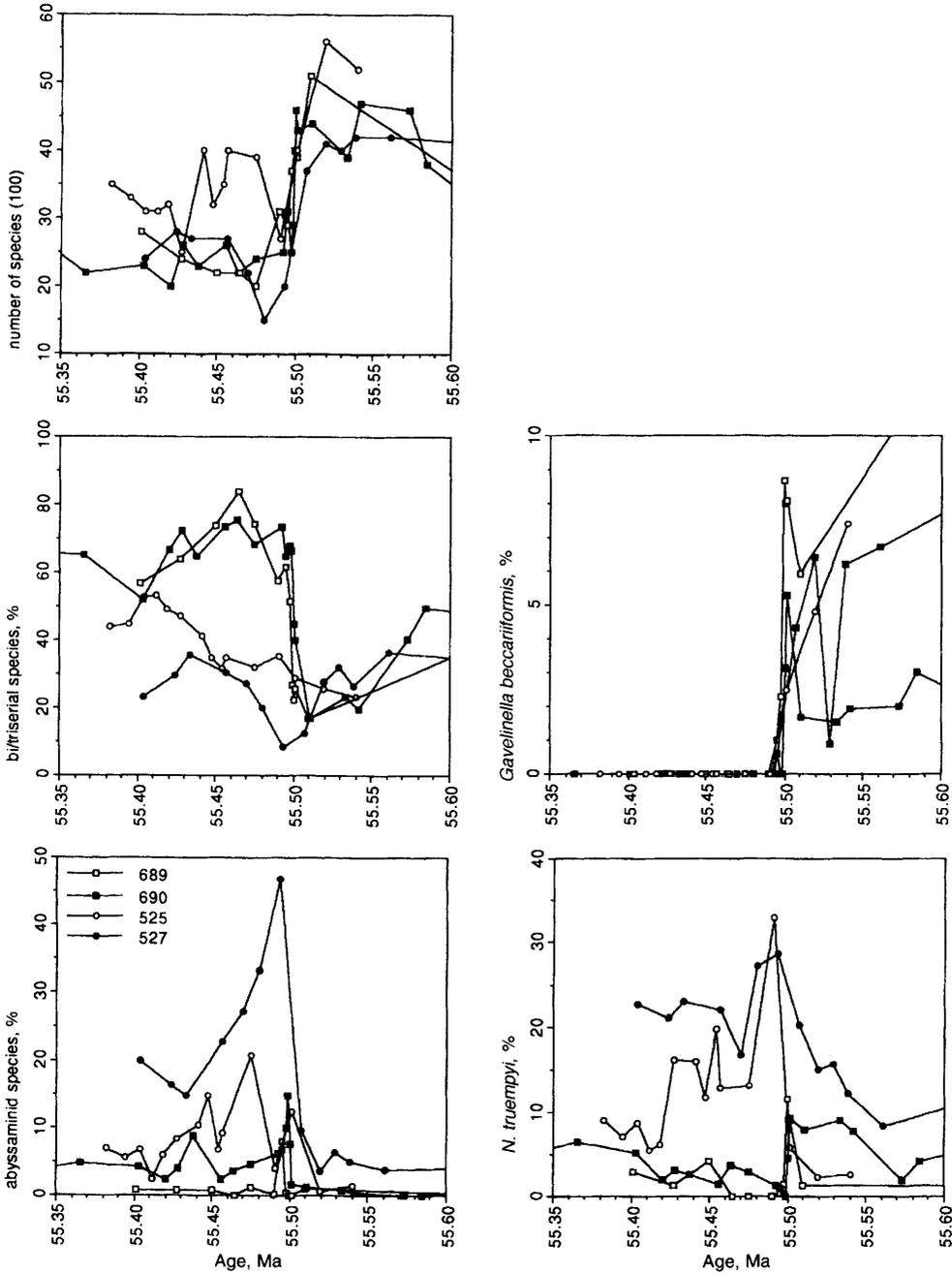


Fig. 8. Number of species (per 100 specimens, calculated by rarefaction, Sanders 1968) and relative abundances of some common species groups of benthic foraminifera, compared for the Walvis Ridge and Maud Rise sites. Closed symbols indicate the deeper site at each location, the open symbols the shallower site.

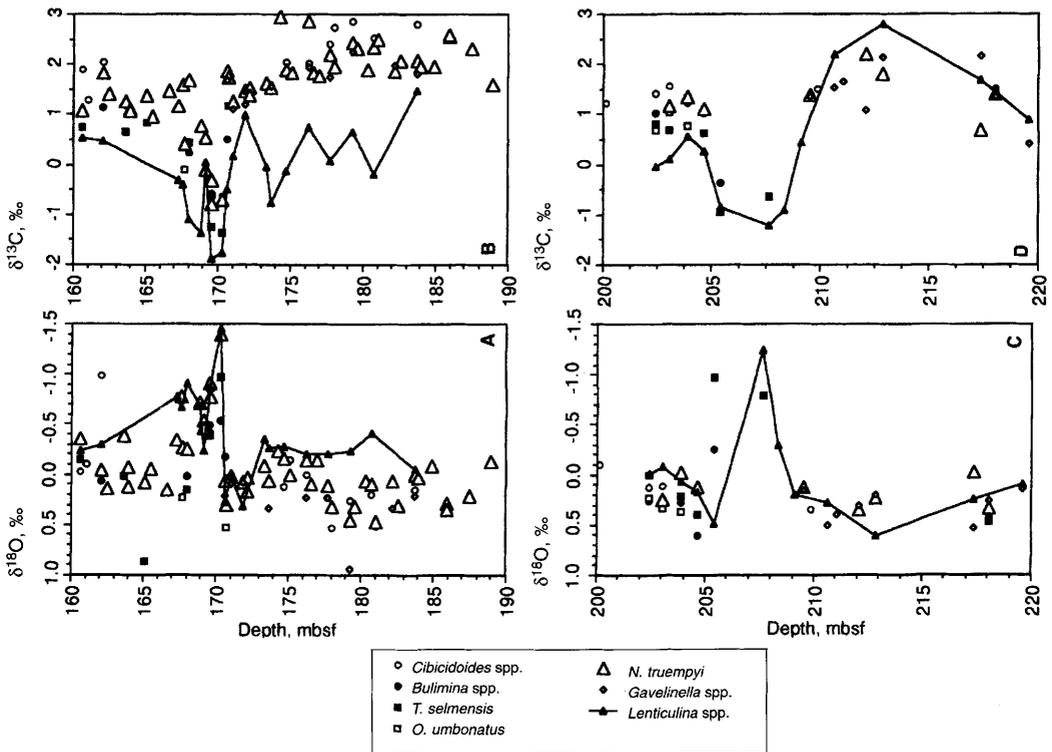


Fig. 9. Oxygen and carbon isotope value for selected benthic foraminiferal species, Sites 689 and 690. Open symbols indicate species considered infaunal because of their morphological resemblance to Recent infaunal taxa (Corliss & Chen 1988; Kaiho 1988, 1991, 1994a; Thomas 1990a); closed symbols show epifaunal species.

than *N. truempyi* (Fig. 9). This is not so at Site 689, and *Lenticulina* spp. values are commonly higher than those of *N. truempyi*. None of the other species in the same samples shows a similar pattern, so that this difference between sites is probably not due to difference in diagenesis. It has been proposed that greater separation in carbon isotopic values of infaunal and epifaunal species results from a higher flux of organic matter to the sea floor (Woodruff & Savin 1985; Zahn *et al.* 1986). With the exception of anomalously low $\delta^{13}\text{C}$ values, the carbon isotopic values of *Lenticulina* spp. are overall lower at Site 690 than at Site 689, thus suggesting a lesser flux of organic matter at the latter site. This is unexpected, because at shallower sites a greater flux of organic matter arrives at the seafloor than at deeper sites at equal productivity (e.g. Berger *et al.* 1994). Higher fluxes of organic matter to the deeper site thus suggest higher productivity at that site. Higher sedimentation rates at Site 690 might be seen as supportive of higher productivity at that site, but

might as well have been caused by admixture of fine terrigenous material (Barker *et al.* 1988).

We can not at present explain why there should have been higher productivity at Site 690, because the sites are very close. Possibly, the pattern of deep-water circulation around Maud Rise caused upwelling at Site 690, on the southwestern flank, but not at Site 689, closer to the top. At present, waters well up along the southwestern flank of the rise (Comiso & Gordon 1982). In the presence of large stratigraphic variations in isotopic values, bioturbation may disturb the signal easily, and we need better age control at Site 689 before we can decide whether the difference in benthic infaunal carbon isotope values at the sites reflects a real environmental difference.

Lenticulina spp. constitute the only group present at both Maud Rise sites over the extinction interval. Comparison of the isotope data of this taxon at both sites (Fig. 9) shows that the excursion had a similar magnitude in the bottom waters of the two sites. Oxygen isotope values at both sites

were not significantly different at both sites before, during or after the excursion.

Isotope excursions

Extreme excursions in carbon and oxygen isotopic values are present in benthic and planktonic foraminiferal records from the Maud Rise sites, and in benthic and bulk records from the Walvis Ridge sites (Figs 3, 11,12,13). Our data for Site 690 agree with those of Kennett & Stott (1990, 1991) and Stott *et al.* (1996). The benthic foraminiferal extinction was synchronous with a very large, negative, short-term excursion of carbon and oxygen isotopes in planktonic and benthic foraminifera and bulk carbonate (Shackleton & Hall 1990). The isotope excursions reached peak negative values in a few thousand years, and values returned to pre-excursion levels more gradually, in about 50 000 years as estimated at Site 690. We used the isotope excursion to construct the time-scale (see above; Fig. 3), and therefore we assumed that the duration was the same at all 4 sites. The carbon isotope excursion was about -2‰ for benthic foraminifera at Walvis Ridge and Maud Rise, about -4‰ for planktonic foraminifera at Maud Rise, leading to decreased vertical gradients.

Bulk carbon isotope values decreased by about 2‰ at the Walvis Ridge sites. The oxygen isotope excursion was about -1.5‰ for benthic foraminifera at Walvis Ridge (1600–3400 m palaeodepth) and Maud Rise (1100–1900 m palaeodepth), about -1‰ for planktonic foraminifera at Maud Rise. There are no significant differences in the oxygen isotope values of benthic foraminifera at Sites 689, 690, 525 and 527 over the full interval studied, suggesting little stratification of the water masses at the sites before, during and after the isotope excursions. This observation suggests that intermediate to deep water masses were poorly stratified, from whatever source they were derived.

Comparison of the planktonic foraminiferal data from Sites 689 and 690 (Fig. 13) shows that oxygen isotopic values were very close throughout the studied sequence, as expected for sites in such close proximity. Possibly the carbon isotope values of planktonic foraminifera were more positive at Site 690, in support of the hypothesis of higher productivity at that site, as outlined above.

Discussion of faunas, isotopes and palaeoenvironments

An investigation of the biogeography of the benthic foraminiferal faunas after the extinction is difficult because many stratigraphical sections have

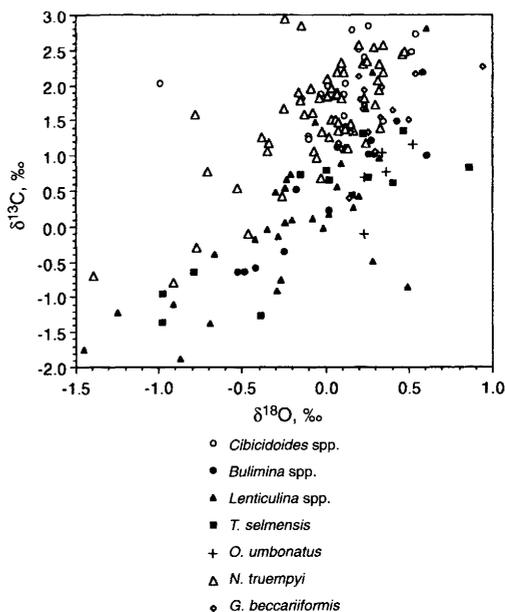


Fig. 10. Scatter plot of $\delta^{13}\text{C}$ versus $\delta^{18}\text{O}$ values of all benthic foraminiferal data for Sites 689 and 690. Presumed infaunal species (*Bulimina* spp., *T. selmensis*, *Lenticulina* spp.) have closed symbols, presumed epifaunal species (*Cibicidoides* spp., *N. truempyi*, *Gavelinella* spp.) have open symbols. *O. umbonatus*, which has an epifaunal morphology according to Corliss & Chen (1988) and Kaiho (1988, 1991) is indicated by +, to show its presence within the field of infaunal species.

hiatuses (Aubry *et al.* 1996). We can therefore only be confident about our time correlations if we compare faunas from the interval of the large, short-lived, negative $\delta^{13}\text{C}$ -excursion. The benthic foraminiferal extinction shows biogeographical variation between the Walvis Ridge and the Maud Rise sites, which were in the same ocean basin and at similar longitude, separated by about 30° of latitude. The pre-extinction faunas were largely cosmopolitan, with relatively minor differences with location as well as depth (e.g. Kaiho 1988, 1991), but post-extinction faunas show very wide differences in composition. At our sites biogeographical differences were more important than depth differences.

At Maud Rise, the extinction was followed by an increase in relative abundance of species presumed to live infaunally (Corliss & Chen 1988). Such an increase most probably reflects an increase of the flux of particulate organic matter to the sea floor (e.g. Lutze & Coulbourn 1984; Pedersen *et al.* 1988; Hermelin & Shimmield 1990; Hermelin

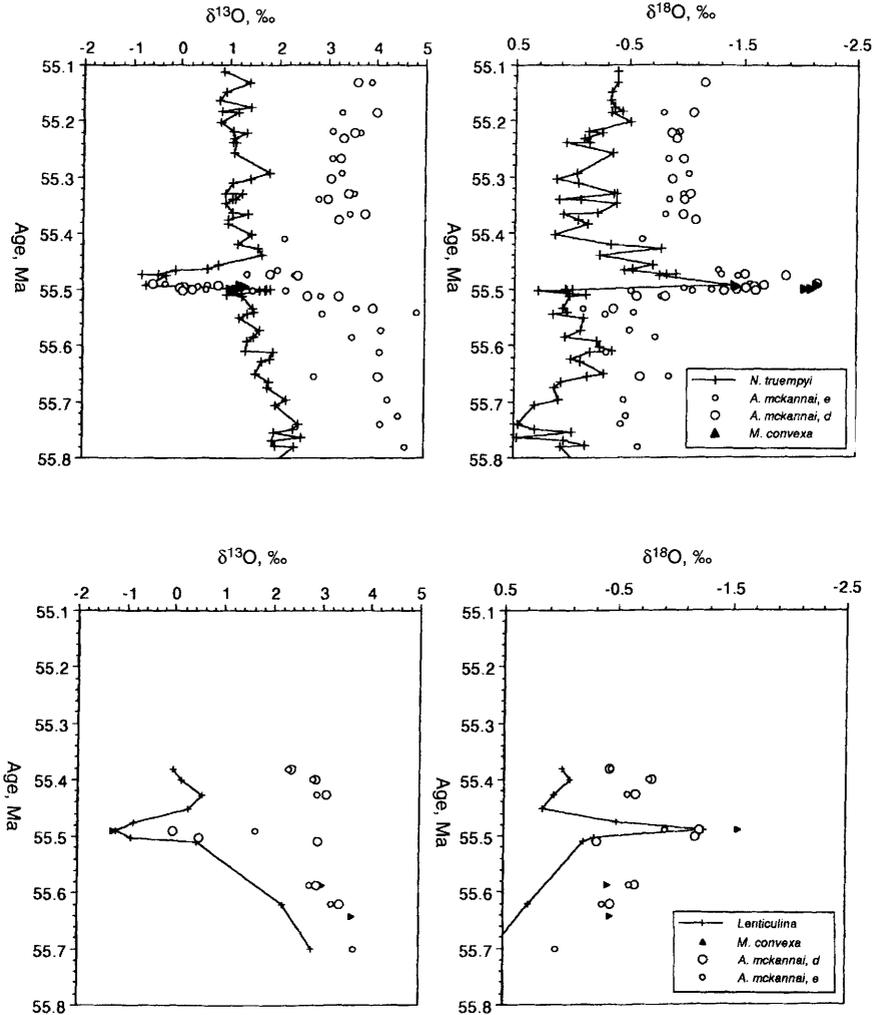


Fig. 11. Carbon and oxygen isotope data for benthic and planktonic foraminifera, Sites 689 and 690. Included are data from Stott *et al.* (1990), Kennett & Stott (1990), and our data.

1992; Rathburn & Corliss 1994), rather than a decrease in dissolved oxygen. This faunal evidence appears to conflict with the carbon isotope data, which show decreased surface-to-deep gradients, suggesting lowered productivity just after the extinction (Kennett & Stott 1991; Fig. 11). We can try to reconcile faunal and isotope data by speculating that the decreased gradients resulted from upwelling of deep, nutrient-rich waters as a result of a changes in deep-water circulation. The upwelling could have caused increased productivity and isotopically lighter total dissolved carbonate in the surface waters. This mechanism can not explain the short-term carbon isotope

excursion, and can be only a part of the explanation for the depressed gradients at Sites 689 and 690. We do not suggest that productivity increased worldwide, only that it increased at Maud Rise, and possibly at other sites at high latitudes, or on continental margins, or close to the equator. We suggest that all these sites may have been characterized by increased deposition of organic matter (thus probably increased surface productivity) after the extinction.

In contrast, benthic foraminifera data suggest that the flux of organic matter to the seafloor decreased after the extinction at the Walvis Ridge sites, in line with suggestions that productivity

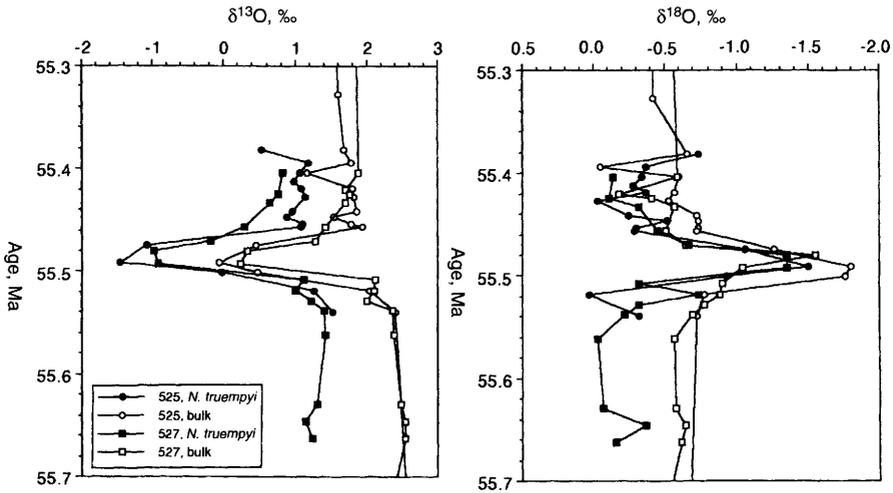


Fig. 12. Carbon and oxygen isotope data for benthic foraminifera and bulk carbonate, Sites 525 and 527. Included are data from Shackleton *et al.* (1984a), Shackleton & Hall (1984), and this paper.

decreased globally at the time (e.g. Corfield & Shackleton 1988; Corfield 1993). Thomas (1990) speculated that this difference might have been apparent only, and the result of the fact that many researchers study such a large size fraction (> 125 μm) that the abundant presence of the small, biserial species would go undetected. Data presented in this study, however, clearly show that the difference is real, not the result of difference in size-fraction studied, nor of the presence of hiatuses, since these different faunas coexisted with the short-lived carbon isotope excursion.

Studies from many locations in the Atlantic Ocean have described post-extinction faunas resembling those at the Walvis Ridge sites, with common to abundant abyssaminid species and *N. truempyi* (Clark & Wright 1984; Tjalsma & Lohmann 1984; Katz & Miller 1991; Müller-Merz & Oberhänsli 1991; Oberhänsli *et al.* 1991; Pak & Miller 1992); similar faunas have also been described from land sections in Spain (Zumaya: G. Keller pers. comm. 1994) and Egypt (Speijer 1994). We cannot be certain that these faunas are indeed coeval with the immediate post-extinction faunas because of the lack of isotopic data or low-resolution sampling. At Site 549 in the Bay of Biscay (Reynolds 1992 MSc Thesis, op. cit.) biserial and triserial species increased in relative abundance after the extinction, as at Maud Rise. In the land section at Caravaca (Spain), a similar faunal pattern was observed (G. Keller pers. comm. 1994).

Faunas from some sites in the Indian Ocean show a pattern similar to that at the Atlantic sites,

but we do not know whether we see the immediate post-extinction faunas because of lack of isotope data, poor recovery, or low time resolution. At Site 762 (palaeodepth 1000–1500 m) the post-extinction faunas were dominated by *N. truempyi*, *Cibicidoides* spp., and abyssaminid species; *T. selmensis* and *A. aragonensis* were present (Thomas, unpub. data; R. Nomura pers. comm. 1994). At Site 747 (Kerguelen Plateau, Indian Ocean; palaeodepth 2000–3000 m) post-extinction faunas were dominated by *N. truempyi*, but at the shallower Site 748 (palaeodepth 600–2000 m) *Stilostomella* spp. and *Lenticulina* spp. dominated (Mackensen & Berggren 1992). At Site 752 (palaeodepth 500–1000 m), post-extinction faunas had common *Anomalinoidea capitatus/danicus*, *N. truempyi* and *Cibicidoides* spp. (Nomura 1991; R. Nomura pers. comm. 1994). At Site 738 (Indian Ocean, Kerguelen Plateau; Fig. 1) recovery was poor, but the short-term isotope excursion occurred in a recovered interval and is similar in magnitude to that at Maud Rise (Lu & Keller 1993). The benthic foraminiferal extinction occurred in the lower part of an interval of dissolution as at Walvis Ridge, and post-extinction faunas coeval with the carbon isotope negative peak had high relative abundances of *Tappanina selmensis* and small buliminids (Thomas, unpub. data) as at Maud Rise. Later faunas became dominated by *N. truempyi* (R. Nomura pers. comm. 1994).

At equatorial Pacific Site 865 (palaeodepth about 1200, palaeolatitude 2°N, Bralower *et al.* 1995a) post-extinction faunas coeval with the short-lived $\delta^{13}\text{C}$ excursion lack *N. truempyi*, have

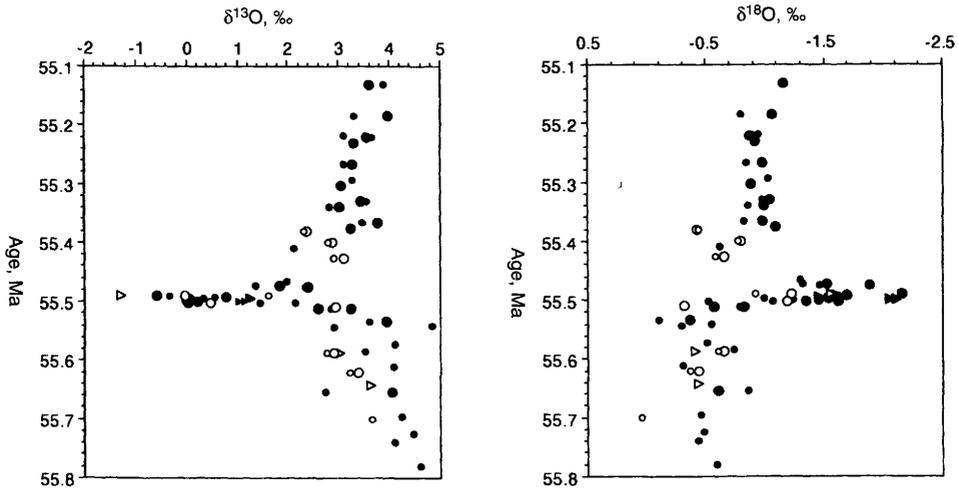


Fig. 13. Carbon and oxygen isotope data for planktonic foraminifera, Sites 689 and 690. Open symbols are data from Site 690, closed symbols are data from Site 689. Circles, *Acarinina mckannai*; triangles, *Morozovella convexa*. Included are data from Stott *et al.* (1990), Kennett & Stott (1990), and this paper.

common *Cibicidoides* spp. and *Bulimina semicostata*, and high relative abundances of bi- and triserial species (Thomas, unpub. data). At Pacific Site 577 (palaeodepth 1500 m) post-extinction faunas show increased relative abundance of *Bulimina semicostata* as well as *N. truempyi* (Miller *et al.* 1987b; Pak & Miller 1992), in contrast to the faunal pattern at Site 865, but there may be a hiatus in the immediate post-extinction interval (Corfield & Cartlidge 1992b; Berggren *et al.* 1995).

The benthic extinction thus was followed by a period of varying faunas at different locations. At all locations *T. selmensis* and *A. aragonensis* are typical for the post-extinction period at bathyal to abyssal depths. These species might probably be seen as opportunists, quickly filling habitats emptied during the extinction. Species that became extinct included common species thought to have lived epifaunally (e.g. *Neoeponides* spp. and *Gavelinella beccariiiformis*), many agglutinant species (e.g. *Dorothia oxycona*, *Tritaxia paleocenica*), but also infaunal species such as *B. thanetensis* and *B. delicatulus*. At some locations infaunal morphotypes increased in relative abundance for a short period (50 000 years), but at many sites in the Atlantic, Pacific and Indian Oceans epifaunal morphotypes increased in abundance, and infaunal morphotypes were replaced by other infaunal morphotypes (e.g. *Bulimina velascoensis* and *Aragonia velascoensis* before the extinction, *Bulimina semicostata* and *Aragonia aragonensis* after the extinction at Pacific sites; Miller *et al.* 1987b; E. Thomas, unpub. data).

It thus appears to be too simplistic to relate the benthic foraminiferal extinction only to increased temperatures with concomitant decreased oxygen levels, although these probably were involved (Thomas 1990; Kaiho 1991). A strong increase in volume of deep to intermediate waters derived from subtropical latitudes rather than from high southern latitudes (as invoked by many authors, including Thomas 1989, 1990, 1992; Kennett & Stott 1990, 1991; Lu & Keller 1993; Zachos *et al.* 1993) could have caused the observed patterns partially by direct and indirect effects on the deep sea benthic faunas: patterns of aging of intermediate and deep waters, and thus their nutrient and oxygen levels, would change as well as temperature. Probably as important for deep sea faunas, a change in direction of deep water circulation might influence the locations where deep, nutrient-rich waters welled up to the surface, and thus where high surface productivity occurred.

If $\delta^{13}\text{C}$ values of deep sea benthic foraminifera during the maximum-excursion were well known from various localities, one might attempt to reconstruct the direction of deep water circulation (e.g. Pak & Miller 1992; Zachos *et al.* 1993), lighter values indicating the presence of 'older' deep to intermediate water masses. Lu & Keller (1993) thus argue that the fact that the benthic carbon isotopic excursion was of smaller magnitude in the Pacific (Pak & Miller 1992) than at Maud Rise (Kennett & Stott 1991) indicates deep-water circulation from Pacific low latitudes to the Antarctic. We do not think that this argument is valid, however, because of the presence of hiatuses

at Site 577 (Corfield & Cartlidge 1992b; Berggren *et al.* 1996), so that we can not be certain that the observed values reflect the maximum excursion. New isotope data from equatorial Pacific Site 865 suggest that deep waters might have been forming at that location just after the extinction (Bralower *et al.* 1995b).

It is not clear whether a change in deep-water circulation could be related to decreased levels of dissolved oxygen in shallower marine areas on continental margins (e.g. North Sea Basin: King 1989; Charnock & Jones 1990; New Jersey continental margin: Gibson *et al.* 1992; southwest Pacific: Moore 1988; Kaiho 1994b; the southern edge of Tethys: Speijer 1994). There might be a link through locally increased productivity (as suggested by Speijer 1994), or decreased levels of dissolved oxygen might have resulted from the high temperatures.

The rapid isotope excursions occurred against the background of longer term events. In the middle Paleocene, bulk carbonate $\delta^{13}\text{C}$ values reached a maximum for the Cenozoic (e.g. Shackleton 1986, 1987), which was not higher than more long-term values during the Late Cretaceous (Corfield *et al.* 1991; Corfield & Cartlidge 1992b). This period of high $\delta^{13}\text{C}$ values could be seen as reflecting return to higher oceanic productivity (Shackleton *et al.* 1985; Shackleton 1986, 1987) after the biosphere recovered from the end-Cretaceous extinction, especially since oceanic $\delta^{13}\text{C}$ gradients were very high at the time (Shackleton *et al.* 1985). This explanation works only if the organic carbon produced was buried in the lithosphere, because $\delta^{13}\text{C}$ changes that persist over periods much longer than the residence time of carbon in the ocean can arise only as a result of long-term changes in the global rate of storage of carbon in organic matter rather than in carbonates (Berger & Vincent 1986; Shackleton 1986, 1987; Corfield & Cartlidge 1992b). It is not obvious where the organic matter required to be deposited in the Paleocene was stored, but extensive organic matter-rich deposits of that age occur in southwestern Pacific (Moore 1988), the North Sea region, and possibly oil shales in the continental United States. Carbon isotopic values decreased over a few million years from the middle Paleocene until the early Eocene (Shackleton 1987; Corfield & Cartlidge 1992b). The long-term decreasing values of $\delta^{13}\text{C}$ have been tentatively explained by increased erosion of organic carbon rich sediments from the Himalyan orogen (Beck *et al.* 1995b).

This long-term decrease occurred at a time when oxygen isotope data suggest increasing temperatures of global deep waters as well as surface waters at high latitudes (Shackleton 1986; Corfield & Cartlidge 1992b; Zachos *et al.* 1994). The high

temperatures have been explained by higher rates of CO_2 emissions from hydrothermal systems (Owen & Rea 1985; Rea *et al.* 1990) or from plateau basalts (Eldholm & Thomas 1993), especially because preliminary results from studies of chemical proxies for palaeo- CO_2 levels suggest that early Eocene $p\text{CO}_2$ levels were 2 to 6 times present values (Freeman & Hayes 1992; Cerling 1991).

What caused the rapid isotope excursions in the latest Paleocene to be superimposed on these long-term trends? Specifically, did sudden events (e.g. emission of large amounts of volcanic CO_2) trigger the rapid isotope excursions, or did gradual changes cross a threshold-value (Zachos *et al.* 1994)? Planktonic foraminiferal data from Sites 689 and 690 (Stott *et al.* 1990, 1996; this paper) and 738 (Lu & Keller 1993) clearly demonstrate that low-latitude, keeled planktonic foraminifera reached high latitudes ($> 60^\circ\text{S}$) before the benthic foraminiferal extinction; extensive warming of high latitude surface waters thus preceded the isotope anomalies. Note that $\delta^{18}\text{O}$ values of keeled planktonics at Site 690 are extremely low, while $\delta^{13}\text{C}$ values of the same specimens have not reached peak values, suggesting that temperature increase slightly pre-dated the $\delta^{13}\text{C}$ excursion (Fig. 11). The rapid oxygen isotope excursion at a time when polar ice-sheets were absent or insignificant should probably be explained by a rapid increase in temperature of high latitude surface waters and deep waters worldwide, by 4–6°C. Such an ocean-wide increase in temperature would have the effect of raising sea-level by about 5–6 m only (Varekamp, pers. comm. 1994), which would not be expected to be noticeable in benthic foraminiferal faunas. Speijer (1994) did not observe a clear sea level signal in neritic sequences in the eastern Mediterranean.

The large, short-term carbon isotope excursion occurred in deep as well as surface waters and in the continental record (Kennett & Stott 1990, 1991; Pak & Miller 1992; Lu & Keller 1993; Stott *et al.* 1996), and was associated with increased CaCO_3 -dissolution over a wide depth and geographical range in the deep oceans (Bay of Biscay: Reynolds 1992 MSc Thesis, op. cit.; Stott *et al.* 1996; Site 738: Lu & Keller 1993; Walvis Ridge: this paper). A large, negative shift in $\delta^{13}\text{C}$ composition of total dissolved inorganic carbon as well as increased dissolution and shallowing of the calcium carbonate compensation depth in the oceans could result from a transfer of isotopically light material from the lithosphere or biosphere into the ocean-atmosphere system (Shackleton 1977; Broecker & Peng 1984). The shape of the excursion (rapid start, more gradual return to almost pre-excursion values) is as modelled for disturbance of the ocean–

atmosphere system by transfer of large amounts of isotopically light carbon into the atmosphere (e.g. Kasting & Walker 1993). The rapidity of the initiation of the excursion (a few thousand years) and its short duration suggest that this transfer is not likely to have been caused by changes in ratio of organic carbon to carbonate deposition or erosion (Shackleton 1987; Corfield & Cartledge 1992*b*; Walker 1993).

One can estimate how much carbon from the terrestrial biosphere must be transferred to the oceans to explain the short-term carbon shift (Shackleton 1977), if one makes assumptions regarding the size of oceanic and atmospheric reservoirs, and the average isotopic composition of plant material, and if one knew the average, global size of the carbon isotope excursion. Estimates of the average isotopic composition of plant material can be made, assuming that only C3 plants were around in the late Paleocene (Cerling 1993). The amplitude of the carbon isotope excursion is not well constrained, because the event was short and many sections have low sedimentation rates and hiatuses (Aubry *et al.* 1996); a value between 1 and 2‰, possibly close to 2‰ seems most probable from present observations (Kennett & Stott 1991; Pak & Miller 1992; Stott *et al.* 1996; Bralower *et al.* 1995*b*).

The sizes of atmospheric and oceanic reservoirs are less easy to estimate: the atmospheric reservoir may well have been several times larger than today, which would have led to a larger oceanic reservoir. On the other hand, higher deep ocean temperatures may have caused a decrease in size of the oceanic reservoir. In addition, the size of the oceanic reservoir may have been influenced by the deep sea ventilation rate (Broecker & Takahashi 1984; Thierstein 1989), leading to decreased oceanic reservoir sizes at decreased ventilation rates. Assumption of various values for these parameters, however, indicates that the excursion was so large that transfer of terrestrial biosphere is very unlikely as the cause: such a transfer would require destruction of continental biosphere much larger than that during the end-Cretaceous extinction, which did not occur according to the fossil record.

It is difficult to explain fully such a large shift by rapid emission into the atmosphere of volcanogenic CO₂, derived from the subaerial plateau basalt eruptions of the North Atlantic Volcanic Province (White 1989; White & MacKenzie 1989). The minimum isotopic composition of mid-oceanic ridge volcanic CO₂ is only of the order of -6 to -7‰ (Eldholm & Thomas 1993), so that very large volumes would be necessary. Calculations of the size of plateau-basalt eruptions and the eruption rate required to cause the observed carbon isotope shift have the same problems as the estimates of

biomass required. Average rates of basalt eruption during Chron 24r (several times 10¹³ g CO₂ per year; see Eldholm & Thomas 1993) were certainly too low to cause the excursion. Maximum eruption rates, however, could have been much higher than that (e.g. White & MacKenzie 1989), and an eruption rate of about 100 times the average rate, persisting over a few thousand years, could have been sufficient (Eldholm & Thomas 1993).

The timing of maximum volcanic activity might be seen as problematic for this tentative explanation. The major ash falls, including the most explosive phase of volcanism (ashes -17, +19), clearly postdated the benthic foraminiferal extinction at the Bay of Biscay sites (Knox & Morton 1983, 1988; Backman *et al.* 1984; Aubry *et al.* 1986; Knox 1990; Eldholm & Thomas 1993; Berggren & Aubry 1996). We would, however, not necessarily expect high rates of CO₂-effusion resulting from the most explosive volcanic activity, because basaltic (thus less explosive) magmas have the highest concentrations of CO₂ (Gerlach & Taylor 1990). Ashes and basalt flows are present in sediments from nannofossil zone NP9 and the dinoflagellate *Apectodinium hyperacanthum* zone (Boulter & Manum 1989; Knox 1990; Ali *et al.* 1992), and Berggren & Aubry (1996) tentatively place the initiation of phase 2 volcanic activity in the lower part of zone NP9. The extinction may be very close in age to the change of volcanic activity from western to eastern Greenland. This change may have triggered the break-up of the continent along the East Greenland margin with ensuing volcanism over very large areas adjacent to the line of break-up, which occurred close the chron C25n/C24r boundary (Larsen *et al.* 1992).

We can speculate about the following scenario: high latitude surface ocean temperatures increased, possibly as a result of increased atmospheric pCO₂. These high temperatures may have been increased even more by high CO₂ levels from massive plateau basalt eruptions at the initiation of continental break-up, and they caused a decrease in density of surface waters at high latitudes so that a low density lid formed. Therefore formation of deep to intermediate waters at these latitudes decreased in volume, so that waters sinking at subtropical latitudes could increase in volume. The circulation changes caused changes in the location where deep, nutrient-rich waters could well up to the surface, thus influencing productivity of planktonic organisms, and extinction of deep sea benthic foraminifera. The latter extinctions were thus caused by a combination of increased temperature and concomitant decrease in dissolved oxygen concentration, as well as changes in local productivity – either increased or decreased productivity, depending upon location.

In this scenario, dominant deep water formation at subtropical latitudes was probably a short-term phenomenon, lasting not much longer than 50 000 to maybe 100 000 years, because of the self-limiting effect of deep water circulation from lower to higher latitudes (Pak & Miller 1992; Lu & Keller 1993). Such a circulation pattern would increase heat transport to high latitudes, thus further decreasing thermal gradients; but very low gradients would mean decreased speed of atmospheric as well as oceanic currents (as observed in the decreased intensity of dust transport at the time of the benthic extinction; Miller *et al.* 1987; Rea *et al.* 1990; Hovan & Rea 1992). This decreased transport would then cause cooling of high latitudes (Walker & Sloan 1992; Sloan *et al.* 1995), so that deep to intermediate waters could again form there.

There are problems with this simplistic and speculative scenario. We do not think that volcanic CO₂ input can cause the short-term negative δ¹³C anomaly by itself. It is not clear whether the total dissolved CO₂ content of the oceans would increase or decrease under this scenario: increased temperatures would lead to a decreased reservoir size, but increased atmospheric pCO₂ would lead to an increased oceanic reservoir (Broecker & Takahashi 1984; Walker 1993). Degassing of the ocean when it was warming up would be expected to lead to increased oceanic δ¹³C values, contrary to observations. The increased CaCO₃ dissolution suggests increasing dissolved carbonate levels in the oceans (Broecker & Peng 1984; Kasting & Walker 1993). Stott (1992) suggested, on the basis of changes in isotopic composition of organic material in planktonic foraminifera, that atmospheric pCO₂ decreased sharply during the isotope anomalies, possibly as a result of oceanic CO₂ uptake. But data on carbon isotopic composition of organic marine material cannot simply be interpreted as resulting from changes in atmospheric pCO₂, especially for times when other oceanic parameters (temperature, possible alkalinity) were changing (Goericke & Fry 1994; Hinga *et al.* 1994). In addition, we do not know the full effects of the speculative scenario on oceanic productivity and we do not know the average, global effects on productivity changes that vary from place to place.

At the present time, we do not understand the extreme carbon isotope anomalies in the late Paleocene to early Eocene ocean-atmosphere system. More data on the anomalies, including a better understanding of their maximum extent and geographical variability, may lead to a better under-

standing of this major disturbance of the global carbon cycle, and thus of the functioning of the carbon cycle itself.

Conclusions

1. The deep sea benthic foraminiferal extinction in the latest Paleocene was coeval with short-term, extreme, negative excursions of δ¹³C in planktonic and benthic foraminifera and bulk carbonate, in the South Atlantic and the Weddell Sea, over a range of palaeodepths from 1100–3400 m.

2. The benthic carbon isotope excursion had the same magnitude in both areas, about –2‰. The excursion was a very rapid (beginning in a few thousand years, duration about 50 000 years), and detailed sampling is necessary to evaluate differences in its magnitude in different areas.

3. We do not have enough geographical coverage of the minimum benthic δ¹³C values to use these for reconstruction of deep water circulation.

4. Oxygen isotope data indicate rapid warming of high latitude surface waters as well as global deep waters. It is not clear whether this warming was caused by a change in dominant source region of deep to intermediate waters from high to subtropical latitudes, or by increased surface water temperatures at high latitudes.

5. The biogeography of benthic foraminiferal assemblages after the extinction suggests that a change in dominant source area of deep water formation was involved.

6. The short-term carbon isotope excursion was so large that it probably could not have been caused by transfer of terrestrial biomass into the ocean-atmosphere system; it was so rapid that it probably could not have been caused by a change in deposition or erosion rates of carbon in carbonate as compared to carbonate in organic matter.

7. We cannot explain the short-term negative carbon isotope excursion, which probably is a complex signal, having global as well as local (productivity) components.

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Appendix 2. Faunal data, Site 6690

sample	depth, mbsf	age, Ma	number of species	number of species (100)	number of specimens	Abyssamina poagi	Abyssamina quadrata	Alatamina creta	Allomorphina bigona	Anomalinoidea acutus	Anomalinoidea capitatus	Anomalinoidea semicibratus	Anomalinoidea spissiformis	Anomalinoidea spp.	Aragonia aragonensis	Aragonia velascoensis	Bolivinoidea cf. decorata	Bolivinoidea delicatulus	Bolivinoidea laevigatus	Bulimina macilenta	Bulimina midwayensis	Bulimina ovula	Bulimina cf. semicostata	Bulimina simplex	Bulimina thacensis	Bulimina trinitatensis	Buliminella beaumonti	Caracabulimina sp., small	
16H-1, 41-43	138.21	54.690	56	42	318	1				1							12					18	27						
16H-2, 41-43	139.71	54.729	57	39	326	7			1								6					35	14						
16H-3, 41-43	141.21	54.767	45	31	334	11											3					57	5		4	2			
16H-4, 41-43	142.72	54.805	52	37	314	1	3							1			20					15	30	4				1	
16H-5, 41-43	144.21	54.843	58	38	321	8	3			5	1						9					28	15	2				1	
16H-6, 41-43	145.71	54.881	49	34	328	6	3			1							14					19	7	3				3	
16HCC	147.50	54.926	35	27	291		2										3					80	20	46	3			1	
17H-1, 40-42	147.91	54.937	58	38	339		3			1							12					33	19	1	10				
17H-2, 40-45	149.41	54.974	50	32	337		3			1							6					13	8	84	1	3			
17H-3, 40-42	150.91	55.022	46	34	330		4			1							10					22	24	41	6				
17H-4, 41-44	152.42	55.059	58	35	348		6		1								2		1			30	9	3	1			3	
17H-5, 40-42	153.91	55.095	51	33	313		1		1								8					6	8	1	1	1		5	
17H-6, 40-42	155.44	55.132	56	37	365	5	3		7								3					35	7	3				3	
17H,CC	157.20	55.175	41	32	278		1										10					8	12	4				5	
18H-1, 42-44	157.63	55.185	46	29	320	2	6										9					26	12	2					
18H-2, 40-42	159.11	55.221	42	28	324		1										8					17	18	4	11				
18H-3, 40-42	160.61	55.257	54	37	342		3										8		3			10	27	21	10			3	
18H-4, 40-42	162.12	55.294	47	29	342	1	3										6					33	28	22	29				
18H-5, 40-42	163.61	55.330	40	28	335	3											8					10	13	5	1	4		5	
18H-6, 40-42	165.11	55.366	30	22	344												16					3	19	5	3			1	
18H,CC	168.65	55.403	30	23	289		2										16					16	22	12	7			11	
19H-1, 40-42	167.31	55.420	25	20	318	8											14					28	73	1					
19H-1, 74-76	167.65	55.428	31	26	335	2	4										8					16	30	74	3				
19H-1,114-120	168.07	55.436	30	23	303	1	19										13		1			5	29	76	5				
19H-2, 40-42	168.81	55.456	35	26	335	1											17					49	7	14	117	4	2	1	
19H-2, 74-76	169.16	55.464	34	22	323	3											5					33	8	8	103	12		2	
19H-2,118-124	169.58	55.474	35	24	323	8	5				1						1					5	16	71	14				
19H-3, 40-42	170.31	55.492	34	25	305	18		2									26					26	28	10	3				
19H-3, 51-53	170.42	55.495	40	31	311		19										12	1				12	35	6	10	2			
19H-3, 60-62	170.51	55.497	33	25	313	20				2							14					14	43	4	15	1			
19H-3, 66-68	170.57	55.499	37	29	334	1	27			1							16	8				5	52	5	16				
19H-3, 72-74	170.63	55.500	66	46	318	19						2	4				5	5		1		5	19	26	28			1	
19H-3, 74-76	170.65	55.501	63	43	322	1						5	1				6	2				2	7	11	43				
19H-3,115-121	171.05	55.511	71	44	302	4			1			2	5				12	2				8	9		1				
19H-4, 42-44	171.85	55.533	64	39	324	3					1						1	22				1	13					1	
19H-4, 74-76	172.18	55.542	76	47	312	1			3		2	3					1	3	2			2	9	9	1	5			
19H-5, 40-42	173.31	55.573	66	46	302						1	2					1	2	4			5	5	2	19	1			
19H-5, 74-76	173.69	55.584	60	38	338						2	1					1	11				6		45	3				
19H,CC	174.30	55.601	51	35	273		1				2	1					7					4	2	83					
20H-1, 40-42	174.71	55.612	52	37	296						1	3					7					18	24	24					
20H-2, 40-42	176.21	55.654	73	48	342						4						3	7				10	2	4	6			4	
20H-3, 40-42	177.71	55.696	65	42	307	2			3		1						5	6				1	4	27	1				
20H-4, 41-43	179.22	55.739	61	37	343				2		1						2	3	5			12	5	61				1	
20H,CC	180.26	55.768	52	34	308						1						1					6	8	1	6	25			1
21H-1, 40-42	180.71	55.780	61	41	297		1										1	3				1	3	15	5	27		2	
21H-2, 40-42	182.21	55.822	67	38	334		1										6	2				6	2	7	7				
21H-3, 40-42	183.71	55.864	62	44	316			1			4						5					5	5	4	5	1		1	
21H-4, 5-7	184.66	55.896	55	36	334						1						3	6				3	6	2	22			1	
21H,CC	185.20	55.900	54	36	261												1	11				7	5	56				1	
22H-1, 40-42	185.71	55.915	61	38	323						1	6					2					26	11	9	7				
22H-2, 42-44	187.13	55.962	56	35	330			2	1			8					2					5	11	4	10	4			
22H-3, 42-44	188.63	56.070	61	40	317			1			1	1					2					18	7		2				
22H,CC	191.15	56.163	57	43	269						1	1					11	6				3	2	6	6	1			

Appendix 3. Faunal data, Site 525

sample	depth, mbsf	age, Ma	number of species	number of species (100)	number of specimens	Abyssamina posagi	Abyssamina quadrata	Atabamina creta	Alomorphina trigona	Anomalinoides acutus	Anomalinoides capitatus	Anomalinoides semicribratus	Anomalinoides spissiformis	Anomalinoides spp.	Aragonia aragonensis	Aragonia velascoensis	Bolivinooides cf. decorata	Bolivinooides delicatulus	Bolivinooides laevigatus	Bulimina callahani	Bulimina macienta	Bulimina midwayensis	Bulimina ovula	Bulimina cf. semicostata	Bulimina simplex	Bulimina thanetensis	Bulimina trinitatis	Bulimella baumonti	Ceratobulimina sp. small
32-4, 75-77	388.86	55.383	46	35	313	12						1	2		2					3		26		66	11				
32-4, 125-127	389.36	55.394	47	33	322	4							4	4							4		7		51	12			
32-5, 25-27	389.86	55.404	42	31	312	10							3		6							5	18	6	56				
32-5, 48-50	390.08	55.412	47	31	329	5								9	2						3		14		86	4			
32-5, 77-79	390.38	55.419	39	32	330	5							3		9		2					15		16		59	5		
32-5, 115-117	390.76	55.427	34	25	309	14								4	2							10		17		30	1		
32-6, 25-27	391.36	55.441	52	40	350	17								34		1						2	17		27	10		1	
32-6, 50-52	391.61	55.447	43	32	324	20						1		29	2	5						6		13		18	2		
32-6, 76-78	391.87	55.454	49	35	326	9		1					7	6	5							3		13		11	17		
32-6, 90-92	392.01	55.457	53	40	313	7							4		21		4					2		13		13	11		
32-6, 113-115	392.24	55.475	54	39	343	44									11		1						19		14	1	8	1	
32-6, 130-132	392.41	55.491	34	27	328	13									34		5						3		18	11	2		
32-6, 145-147	392.58	55.501	60	40	317	37	1					1	9		10	2	1						3	9	6	7	8	1	
32-7, 22-24	392.81	55.519	81	56	352	2	2						5	6			3						1	20	3	4	1	3	
32-7, 53-55	393.11	55.540	73	52	309	4	1					8	5				3	1					3	13		5			

sample	depth, mbsf	age, Ma	number of species	number of species (100)	number of specimens	Karenella subglabra	Lenticulina spp.	Neoflabellina reticulata	Neosponides hillebrandi	Neosponides lunata	Nonion havanense	Nonion longicamerata	Nonionella robusta	Nuttallides umbonifera	Nuttallides trowmpti	Nuttallides sp. flat	Nuttallides sp. high	Nuttallinella florealis	Ordozalis nitidus	Ordozalis umbonatus	Orthonomophina spp.	Oscingularia reverrona	Oscingularia velascoensis	polymorphinid species	pleurostomellid species	Patellina corrugata	Pseudoparrella sp.	Pseudopatelellinoides sp.	Pullenia bulioides	Pullenia coryelli	Pullenia jarvisi
32-4, 75-77	388.86	55.383	46	35	313		19													10	2										
32-4, 125-127	389.36	55.394	47	33	322		12			4	1				23					11	7										
32-5, 25-27	389.86	55.404	42	31	312		11			8			2	27						8	2			1	20						
32-5, 48-50	390.08	55.412	47	31	329		17			6	6	2	18							12	1			1	24						
32-5, 77-79	390.38	55.419	39	32	330		19			6	5	3	20							15	8			1	28						
32-5, 115-117	390.76	55.427	34	25	309		12			15	6	2	50							7					12						
32-6, 25-27	391.36	55.441	52	40	350		14			7	1	2	56							1		2	1	27			1				
32-6, 50-52	391.61	55.447	43	32	324		18			14	4	4	38							6	1			22							
32-6, 76-78	391.87	55.454	49	35	326		13			12	2	4	65							10	2			2	23						
32-6, 90-92	392.01	55.457	53	40	313		12			15	1	6	40							9	4			1	23						
32-6, 113-115	392.24	55.475	54	38	343		10			10	8		45							8	2			3	35						
32-6, 130-132	392.41	55.491	34	27	328		10			4	7	108	2							17	2		1	22							
32-6, 145-147	392.58	55.501	60	40	317		4		2	6	5	5	18	5	4				5	48	1		1	4	14		1	5			
32-7, 22-24	392.81	55.519	81	56	352		12	1	1	4	4	1	8	8	2				2	9			1	7	20		1	5			
32-7, 53-55	393.11	55.540	73	52	309		14	1		1	6	5	5	8	2	18				4	10	4		11	8					8	

Appendix 4. Faunal data, Site 527

sample	depth, mbsf	age, Ma	number of species	number of species (100)	number of specimens	Abyssamina poogi	Abyssamina quadrata	Alabamina creta	Allomorphina trigona	Anomalinoidea acutus	Anomalinoidea capitatus	Anomalinoidea semicibratus	Anomalinoidea spissiformis	Anomalinoidea spp.	Aragonia aragonensis	Aragonia velascoensis	Bolivinoidea cf. decorata	Bolivinoidea delicatulus	Bolivinoidea laevigatus	Bulimina callahani	Bulimina macilenta	Bulimina midwayensis	Bulimina ovula	Bulimina cf. semicosata	Bulimina simplex	Bulimina tharretensis	Bulimina trinitatis	Buliminella beaumonti	Ceratobulimina sp., small
24-1. 62- 64	199.63	55.504	31	24	320		20						20																
24-1. 85- 89	199.86	55.424	37	28	313		13						21																
24-1.102-104	200.03	55.433	39	27	324		20					12			8		16												
24-1.136-138	200.37	55.457	32	27	304		30					11			18		8												
24-2. 5- 7	200.56	55.470	31	22	315		50					26			19		15												
24-2. 20-22	200.71	55.480	19	15	342		58					20			1		2												
24-2. 38- 40	200.89	55.493	24	20	216		64					12			3		1												
24-2. 56-58	201.07	55.507	53	37	349		25	1		3		3	17		2		12												
24-2. 68- 70	201.19	55.519	62	41	344		10					3	3				12												
24-2. 80- 82	201.31	55.529	57	40	320		18					1	19				7												
24-2. 88- 90	201.39	55.538	63	42	321		7					11					7												
24-2.115-117	201.66	55.561	60	42	312		6		2			6	2				4	1	1										
24-3. 28- 30	202.29	55.629	55	41	332		7					5					2	1											
24-3. 48-50	202.49	55.648	60	42	316		2	1				1	6				6		2										
24-3. 75- 77	202.76	55.662	62	42	317		2					4	8				6		1										

sample	depth, mbsf	age, Ma	number of species	number of species (100)	number of specimens	Kanerehella subglabra	Lenticulina spp.	Necrolabellina reticulata	Neosponoides hillebrandi	Neosponoides lunata	Nonion havanense	Nonionella longicamerata	Nonionella robusta	Nuttallides umbonifera	Nuttallides truempyi	Nuttallides sp., flat	Nuttallides sp., high	Nuttallinella florealis	Orkdesalis nitidus	Ondosalis umbonatus	Orthomorphina spp.	Ossangulana navarroana	Ossangulana velascoensis	polymorphinid species	pleurostomellid species	Patellina corrugata	Pseudoparrella sp.	Pseudopatellinelloides sp.	Pullenia bulloides	Pullenia coryelli	Pullenia jarvisi
24-1. 62- 64	199.63	55.504	31	24	320							15	8	6	6					18											
24-1. 85- 89	199.86	55.424	37	28	313							15	5	61						14											
24-1.102-104	200.03	55.433	39	27	324						3	1	8	12	63					7				2							
24-1.136-138	200.37	55.457	32	27	304						4		6	13	54					13	3			2							
24-2. 5- 7	200.56	55.470	31	22	315								11	2	51					11				2							
24-2. 20-22	200.71	55.480	19	15	342								5	26	87					24											
24-2. 38- 40	200.89	55.493	24	20	216					3			3	7	50					6											
24-2. 56-58	201.07	55.507	53	37	349		3		23	36	10	13	33	8	30				10	17			1								
24-2. 68- 70	201.19	55.519	62	41	344		6		4	4	12	1	8	37	14	1			4	18	2			11							
24-2. 80- 82	201.31	55.529	57	40	320		5		6	12	7	5	35	6	9	2	4	1		2				7							
24-2. 88- 90	201.39	55.538	63	42	321		7	1	2	5	8	4	28	7	4	2	6			6				10							
24-2.115-117	201.66	55.561	60	42	312				3	4	1	5	16	3	7	3	5			3	5		2	15							
24-3. 28- 30	202.29	55.629	55	41	332		7			5	5	6	28	8	4				3	10				9							
24-3. 48-50	202.49	55.648	60	42	316		3			8	12	9	14	15	6				5	1				7			1				
24-3. 75- 77	202.76	55.662	62	42	317		4	2		4	3	11	28	16					5	8				17							

